Light Neutron-Rich Hypernuclei from the IT-NCSM



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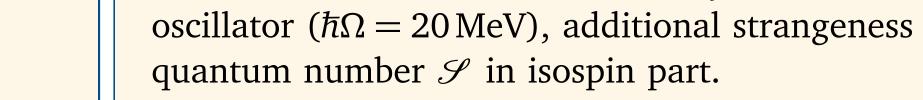
| Motivation | Including Induced YNN Forces | Hypernuclear Importance-Truncated NCSM [2] |
|--|---|--|
| • Explore (broken) symmetries of strongly-interacting matter with strange baryons (hyperons). | • Similarity Renormalization Group (SRG) transformation of the Hamiltonian induces strong repulsive YNN terms \tilde{Y}_{1} | • Starting point: SRG-evolved Hamiltonian with NN+3N interaction [3, 4], YN interaction [5, $\Lambda_Y = 700 \text{ MeV/c}$]: |
| • Apply χ EFT ideas to general baryon-baryon interaction, treating π, K, η as pseudo-Goldstone bosons. | V_{YNN} [1]: $T_{int} + V_{NN} + V_{YN} + V_{3N}$ | $H = T_{int} + M + V^{[2]} + V^{[3]}.$ |
| Few scattering data, determination of LECs challenging. | ↓ SRG | • Expand Hamiltonian on finite Slater Determinant basis. |
| • But: structure of hypernuclei experimentally accessible | $	ilde{T}_{	ext{int}} + 	ilde{V}_{NN} + 	ilde{V}_{YN} + 	ilde{V}_{3N} + 	ilde{V}_{YNN}.$ | • Single-particle basis $ n(ls)jm, \mathcal{S}tm_t\rangle$: harmonic |

- at bildetaie of hyperinderer experimentally through multiple reaction channels
- \Rightarrow Hyperon separation energies.
- \Rightarrow Excited levels via decay γ rays.
- Baryon-baryon interaction has impact on appearance of hyperons in neutron-star matter
- \Rightarrow Influence on mass-radius relation and maximum mass.
- \Rightarrow Connection to and constraints from astrophysics.
- Use *ab initio* framework as link between data and interactions to select and improve models.
- Leading-order χ EFT hyperon-nucleon interaction provides surprisingly good description of observables. But: need explicit YNN terms for SRG-evolved YN.
- Explore experimentally inaccessible parts of the hypernuclear landscape
- \Rightarrow Go neutron-rich.

• Include terms in IT-NCSM calculation explicitly. \Rightarrow Solve SRG flow equation in three-body space:

 $\partial_{\alpha} H_{\alpha} = (2m_N)^2 [[T_{\text{int}}, H_{\alpha}], H_{\alpha}], H_{\alpha} = H.$

- Use Jacobi HO basis to separate center-of-mass d.o.f. and keep basis sizes manageable, truncate total energy $E \leq E_{3,\max}$.
- Isolate genuine three-body part by subtracting two-body interaction evolved in two-body space.
- Adapt HO basis parameter $\hbar\Omega$ from value optimal for SRG evolution to value providing optimal N_{max} convergence of target observables.
- Convert to single-particle coordinates.
- Decouple to *m* scheme during many-body calculation.



- Include all particle species combinations with correct charge and strangeness, e.g. $np\Lambda$, $pp\Sigma^-$, $np\Sigma^0$, $nn\Sigma^+$.
- Limit number of HO excitation quanta to N_{max}
- Compute matrix representation of *H*, diagonalize.

Importance Truncation

- Many basis states contribute very little to low-lying states.
- \Rightarrow Neglecting introduces only small error.
- Estimate contribution for basis state $|\phi_i\rangle$ from 1st-order perturbation theory

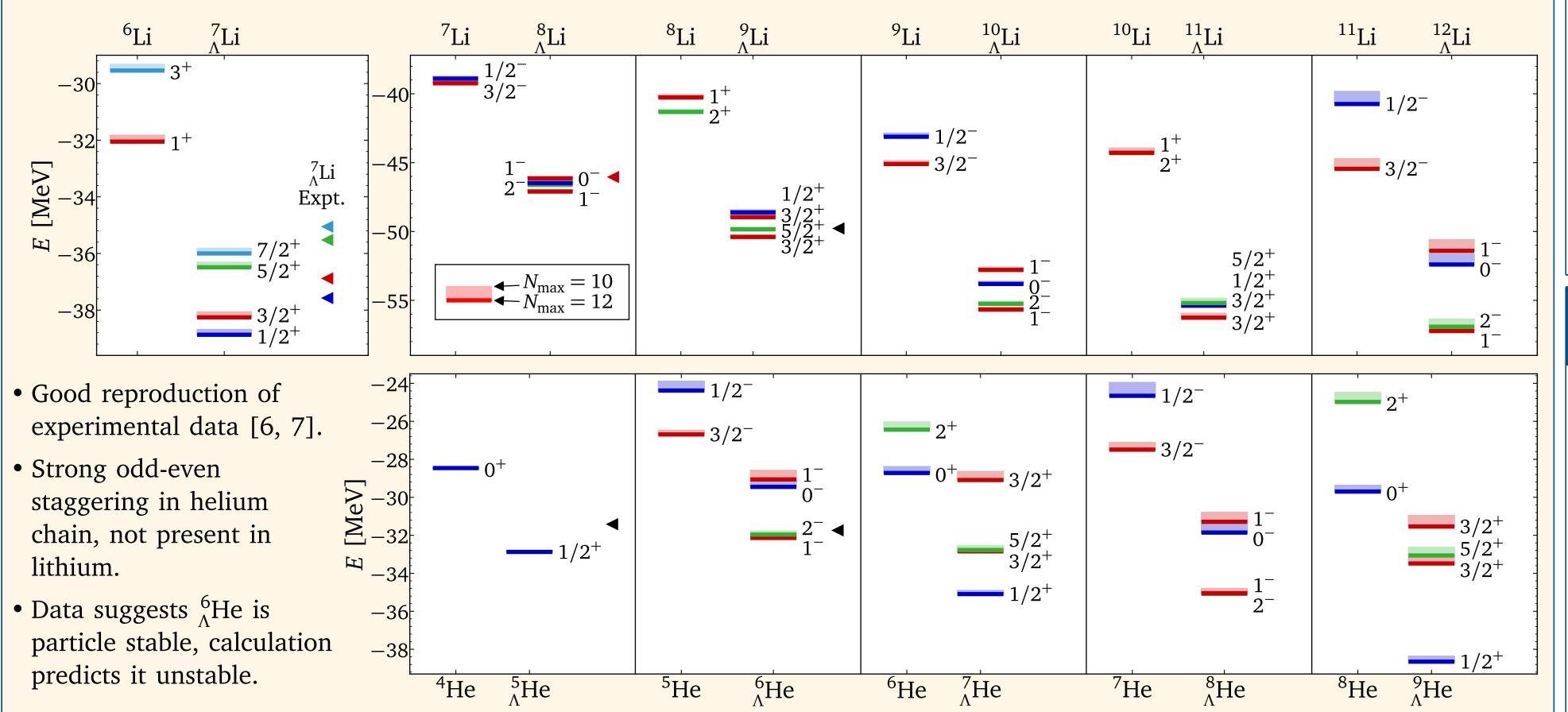
$$\kappa(|\phi_i\rangle) = -\frac{\langle \phi_i | \boldsymbol{H} | \psi_{\text{ref}} \rangle}{\Delta \epsilon_i}.$$

State $|\psi_{ref}\rangle$: approximation to target state from smaller model space. Unperturbed energy difference $\Delta \epsilon_i$ contains Λ - Σ mass difference.

• Build IT model space $\mathcal{M}(\kappa_{\min}) = \{ |\phi_i\rangle : |\kappa(|\phi_i\rangle) | \ge \kappa_{\min} \}.$

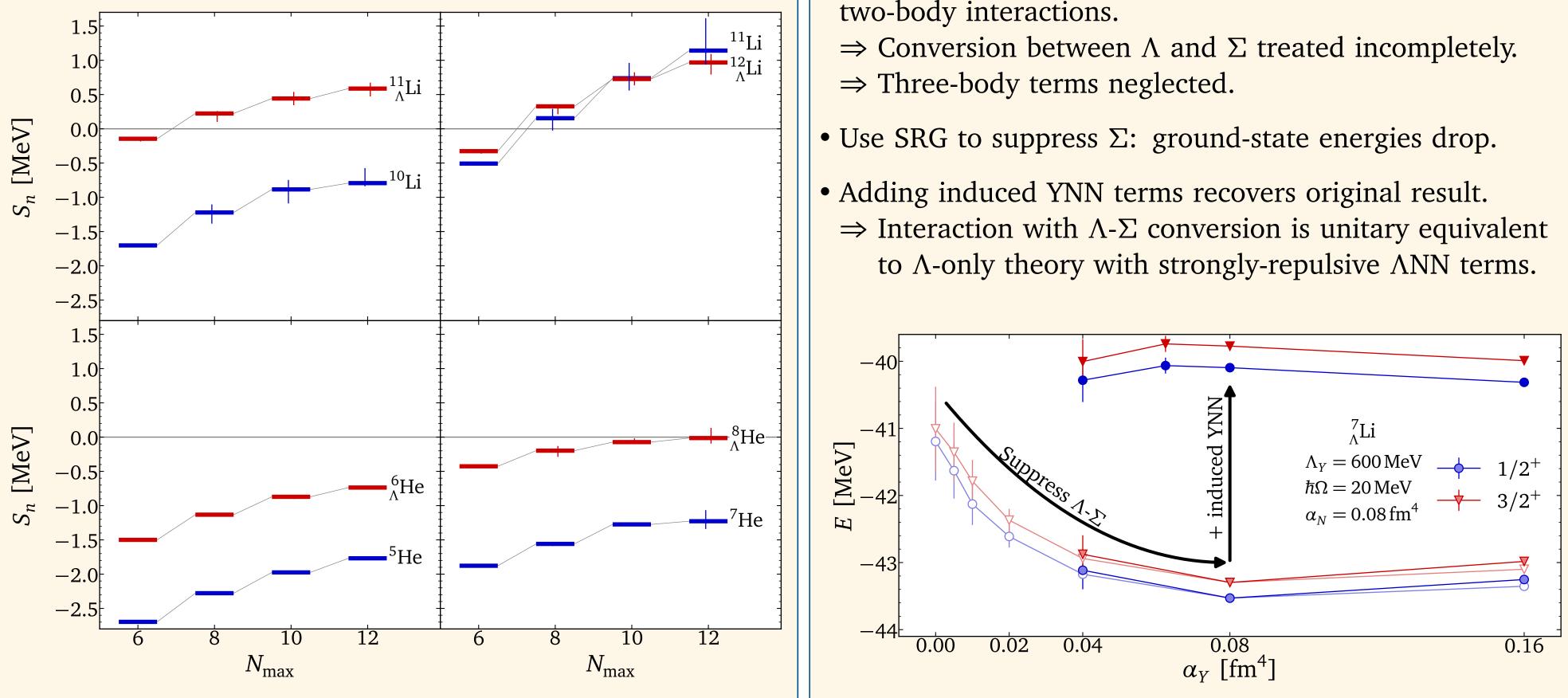
• Diagonalize for multiple thresholds κ_{\min} , extrapolate $\kappa_{\min} \rightarrow 0$ to recover full-space result.

Ground- and Excited-State Energies



Shifting the Neutron Dripline with Hyperons

- Presence of a hyperon can strongly modify properties of hypernucleus compared to nucleonic parent.
- Additional attraction provided by hyperon-nucleon interaction can stabilize particle-unstable cores.
- Hyperon separation energy increases by approx. 1 MeV per additional nucleon; effect more pronounced in



• Raise N_{max} until convergence, use eigenstate $|\psi\rangle$ from $N_{\rm max}$ as $|\psi_{\rm ref}\rangle$ for $N_{\rm max} + 2$.

The Hyperon Puzzle

- Nuclear matter at high tends to favor conversion of nucleons to hyperons: less energy needed to add low-momentum hyperon than for a nucleon at k_F .
- Hyperon-nucleon interaction can enhance or suppress conversion, also modifies compressibility.
- Conventional calculations: conversion causes softening of matter EoS, very small maximum neutron-star masses (hyperon puzzle).
- Often solved by adding strongly-repulsive YNN terms.
- But: calculations either use schemes without Σ hyperons or with Σ hyperons, but with *G*-matrix renormalized two-body interactions.



• Sample cases:

- $-{}^{6}_{\Lambda}$ He: not enough to stabilize.
- ${}^{6}_{\Lambda}$ He & ${}^{11}_{\Lambda}$ Li: hyperon provides additional binding to stabilize nucleonic cores.
- $-\frac{12}{4}$ Li: No additional neutron binding. Indication of proximity to neutron drip line?
- Caveat: YN force overbinds ${}^{5}_{\Lambda}$ He \Rightarrow S_n of ${}^{6}_{\Lambda}$ He too small

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[1] R. Wirth, and R. Roth, Phys. Rev. Lett. 117, 182501 (2016). [2] R. Wirth, et al., Phys. Rev. Lett. **113**, 192502 (2014). [3] D. R. Entem, and R. Machleidt, Phys. Rev. C 68, 041001 (2003). [4] P. Navrátil, Few-Body Syst. 41, 117–140 (2007). [5] H. Polinder, et al., Nucl. Phys. A 779, 244–266 (2006). [6] O. Hashimoto, and H. Tamura, Prog. Part. Nucl. Phys. 57, 564–653 (2006)

[7] D. H. Davis, Nucl. Phys. A 754, 3–13 (2005).

