

Frontiers in Ab Initio Nuclear Structure Theory

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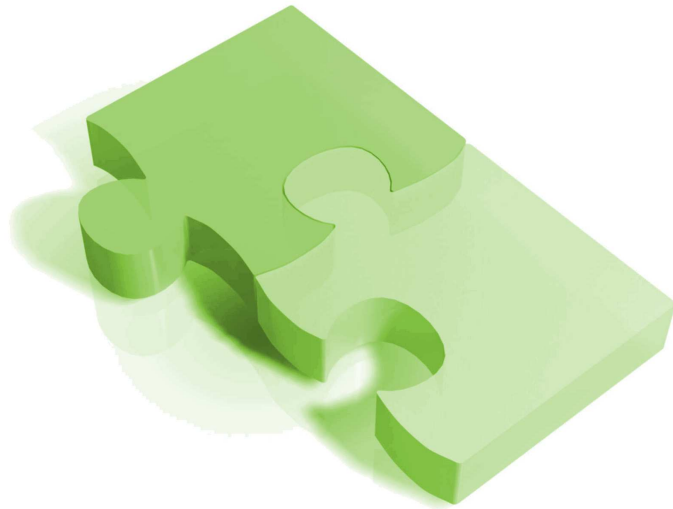
New Era of Nuclear Structure Theory

- **QCD at low energies**

improved understanding through effective field theories & lattice simulations



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- **quantum many-body methods**

advances in ab initio treatment of the nuclear many-body problem

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- **computing & algorithms**

increase of computational resources & improved algorithms

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improved understanding through effective field theories & lattice simulations

- **quantum many-body methods**

advances in ab initio treatment of the nuclear many-body problem

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increase of computational resources & improved algorithms

- **experimental facilities**

amazing perspectives for the study of nuclei far-off stability

Ab Initio Nuclear Structure

Nuclear Structure Observables

Lattice QCD

quarks & gluon on a lattice

Lattice EFT

nucleons & pions on a lattice

Exact Solutions
solve nuclear many-body problem with converged truncations

Controlled Approx.
treat many-body problem with controlled & improvable approximations

Similarity Transformations

physics-conserving unitary transformation to adapt Hamiltonian to limited model space

Chiral EFT Hamiltonians

consistent NN,3N,... interactions & current operators

Chiral Effective Field Theory

based on relevant degrees of freedom & symmetries of QCD

Energy-Density Funct.

guided by chiral EFT

Low-Energy Quantum Chromodynamics

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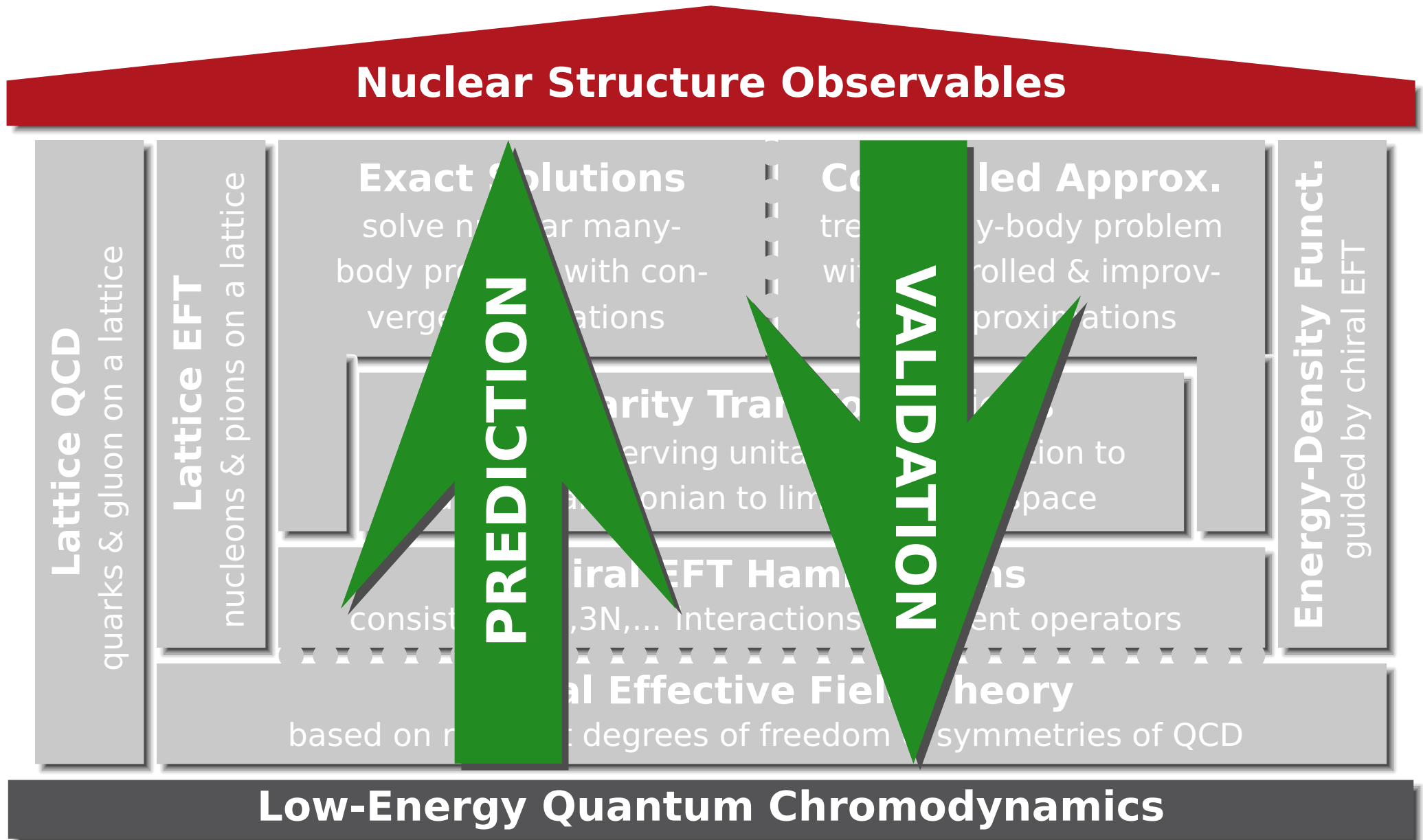
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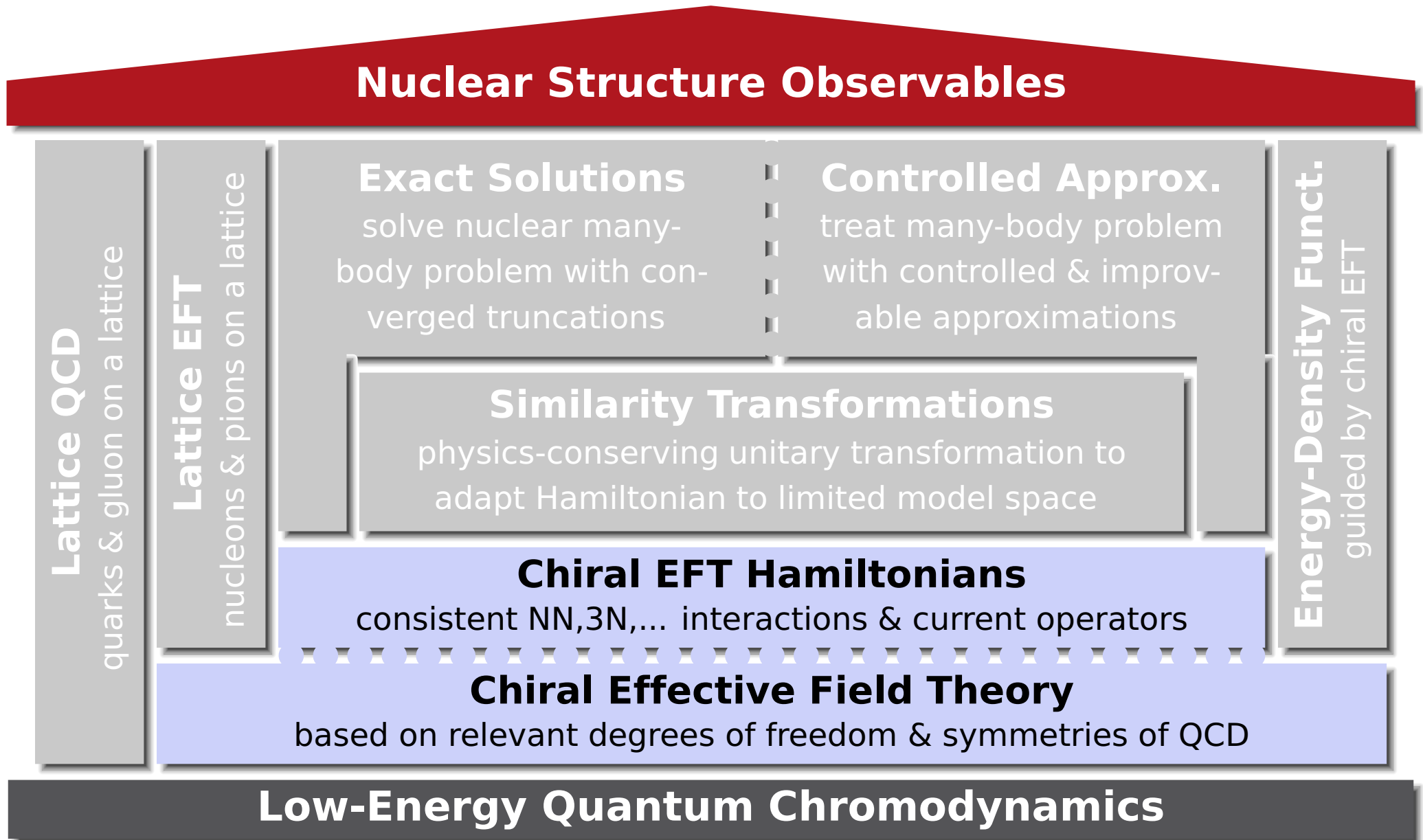
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Low-Energy Quantum Chromodynamics

Ab Initio Nuclear Structure



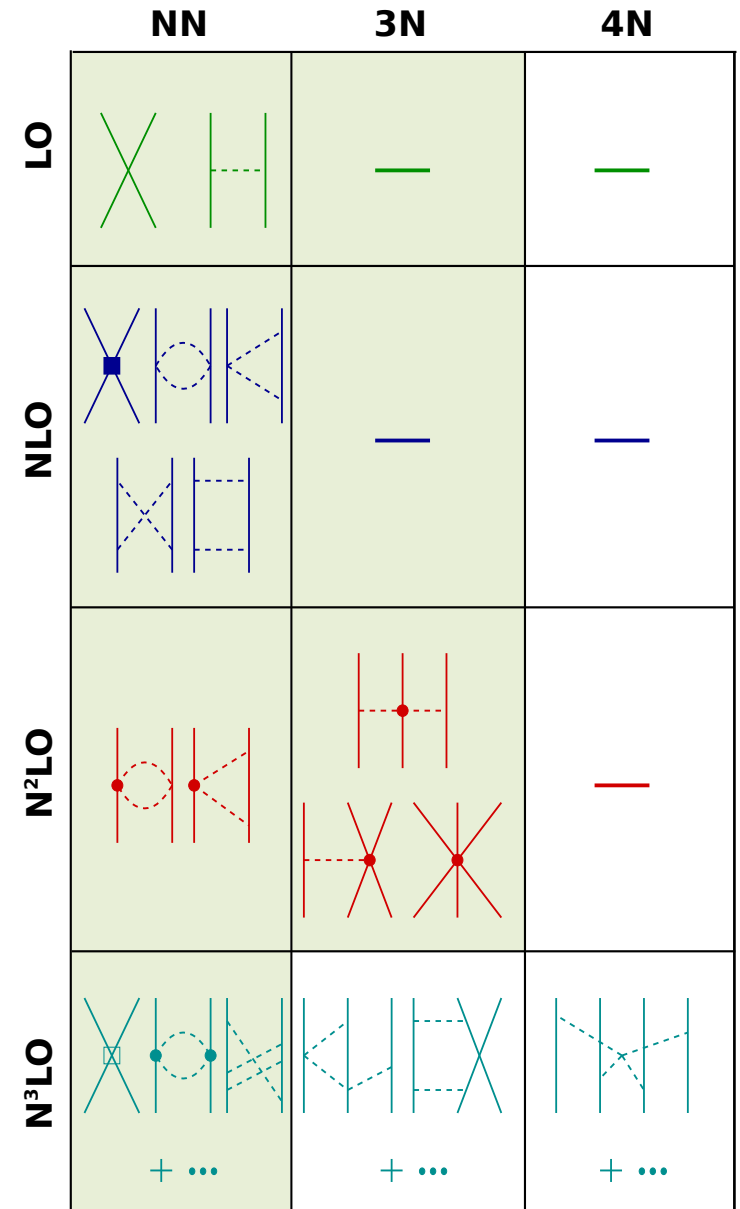
Ab Initio Nuclear Structure



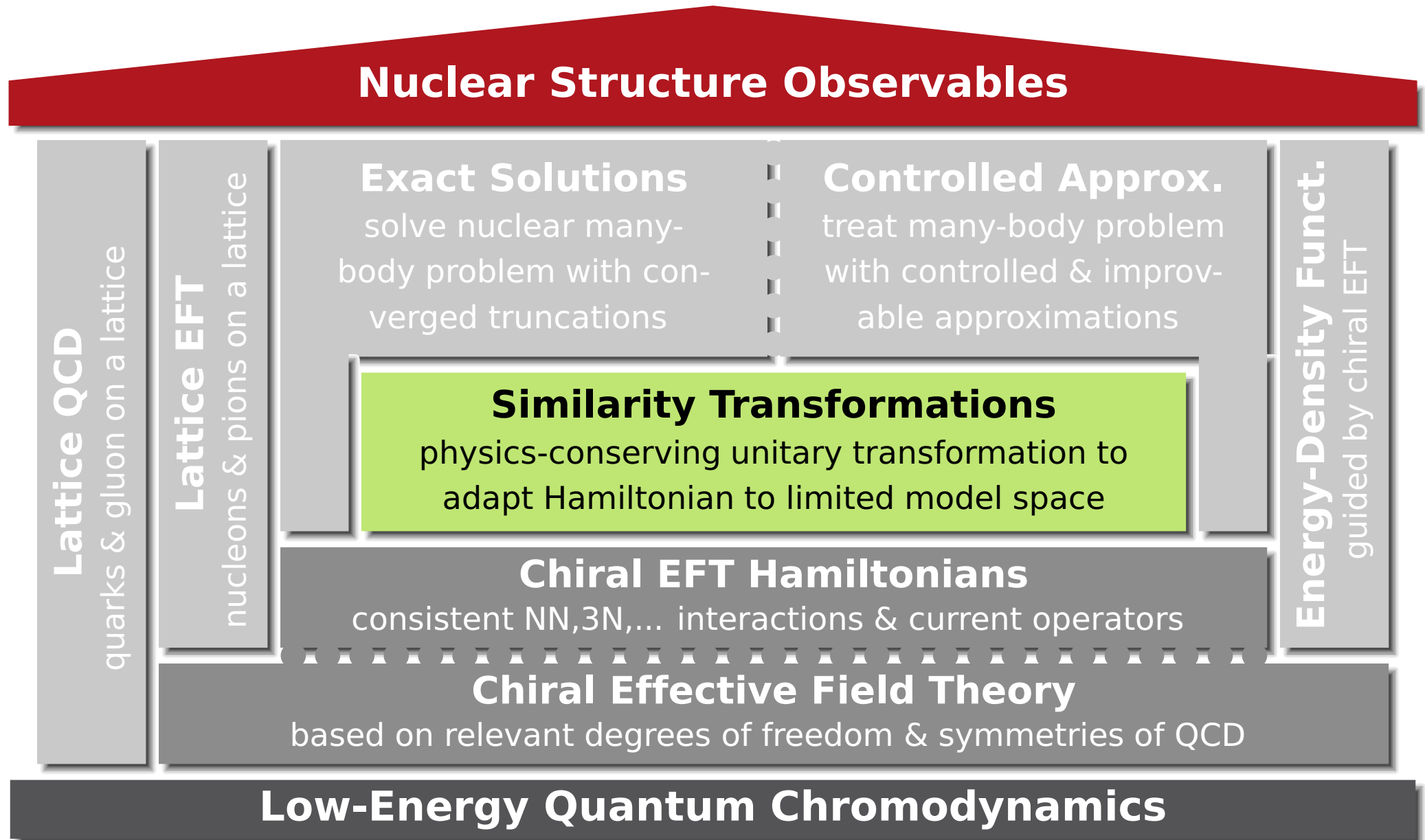
Nuclear Interactions from Chiral EFT

Weinberg, van Kolck, Machleidt, Entem, Meißner, Epelbaum, Krebs, Bernard,...

- low-energy **effective field theory** for relevant degrees of freedom (π, N) based on symmetries of QCD
- long-range **pion dynamics** explicitly, short-range physics absorbed in **contact terms** fitted to data ($NN, \pi N, \dots$)
- hierarchy of **consistent NN, 3N, ... interactions** plus currents
- **standard Hamiltonian:**
 - NN at N3LO: Entem & Machleidt, 500 MeV cutoff
 - 3N at N2LO: Navrátil, A=3 fit, 500 MeV cutoff
- many ongoing developments



Ab Initio Nuclear Structure



Similarity Renormalization Group

Wegner, Glazek, Wilson, Perry, Bogner, Furnstahl, Hergert, Roth, Jurgenson, Navratil,...

continuous transformation driving
Hamiltonian to band-diagonal form
with respect to a uncorrelated basis

- **unitary transformation** of Hamiltonian

$$H_\alpha = U_\alpha^\dagger H U_\alpha$$

- **evolution equations** for H_α and U_α

$$\frac{d}{d\alpha} H_\alpha = [\eta_\alpha, H_\alpha]$$

- **dynamic generator**: commutator with the operator in whose eigenbasis H_α shall be diagonalized

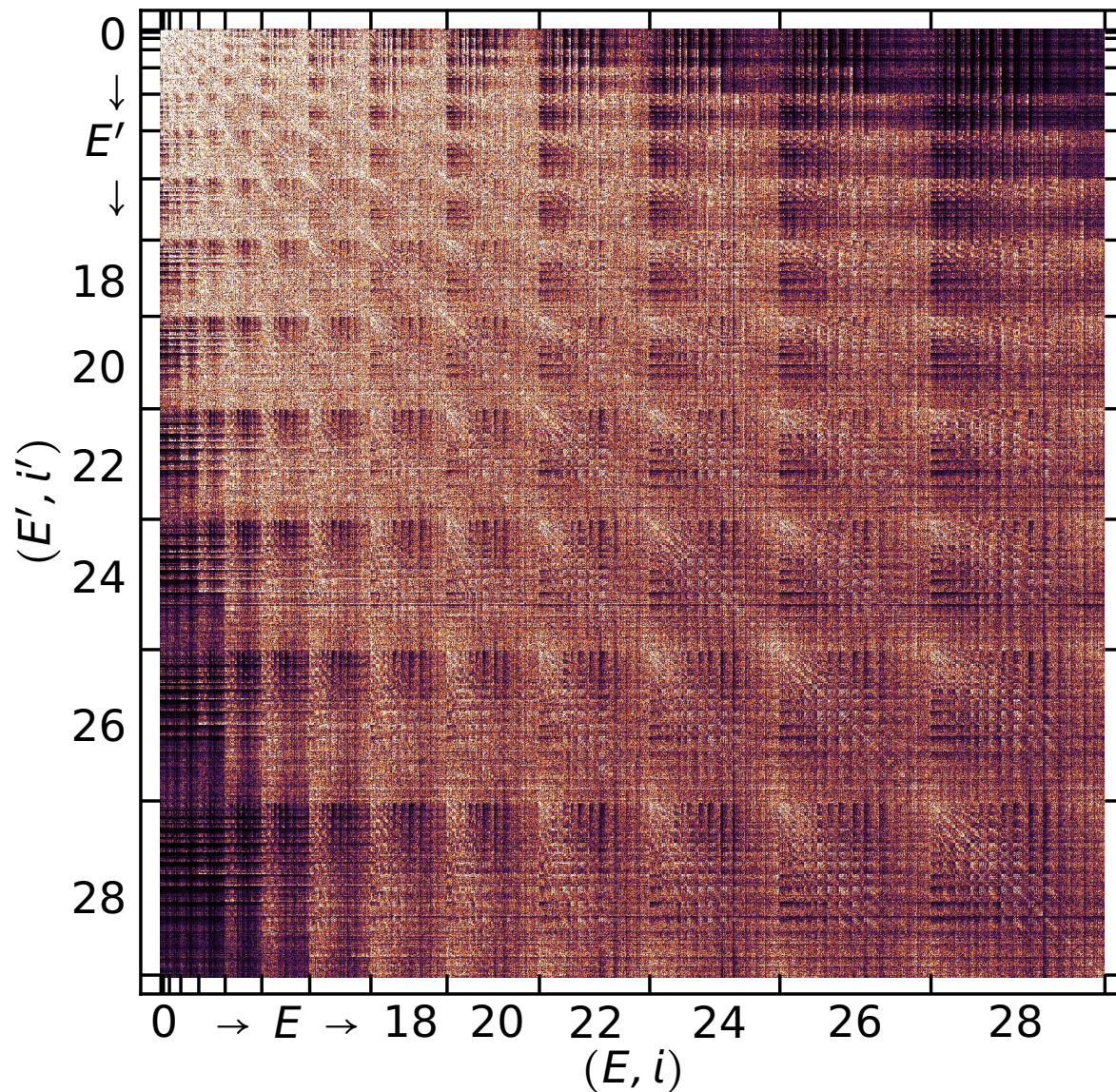
$$\eta_\alpha = (2\mu)^2 [T_{\text{int}}, H_\alpha]$$

simplicity and flexibility
are great advantages of
the SRG approach

solve SRG evolution
equations using two-,
three- & four-body matrix
representation

SRG Evolution in Three-Body Space

3B-Jacobi HO matrix elements

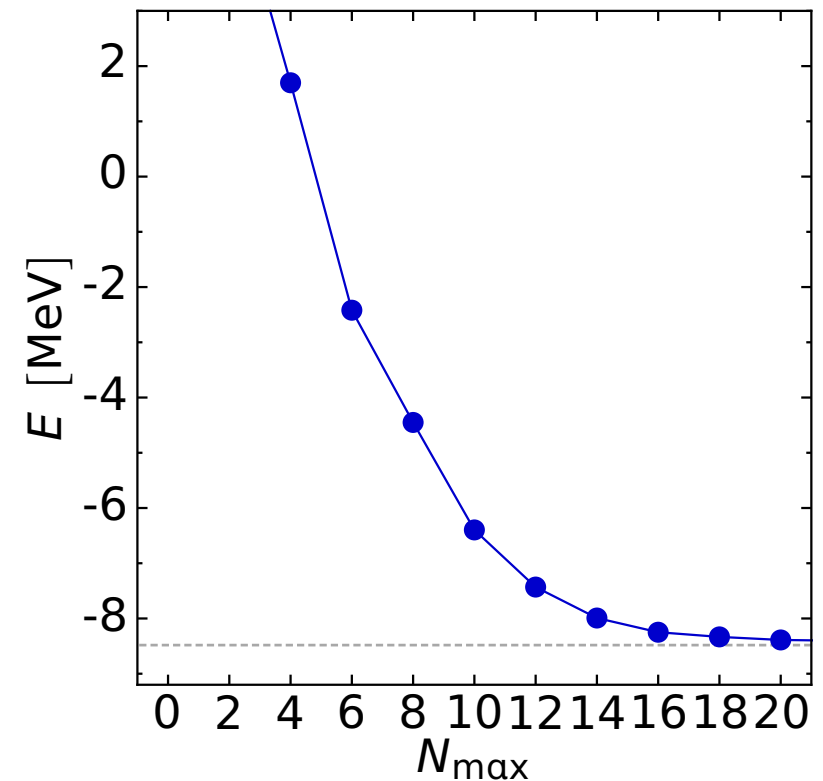


$$\alpha = 0.000 \text{ fm}^4$$

$$\Lambda = \infty \text{ fm}^{-1}$$

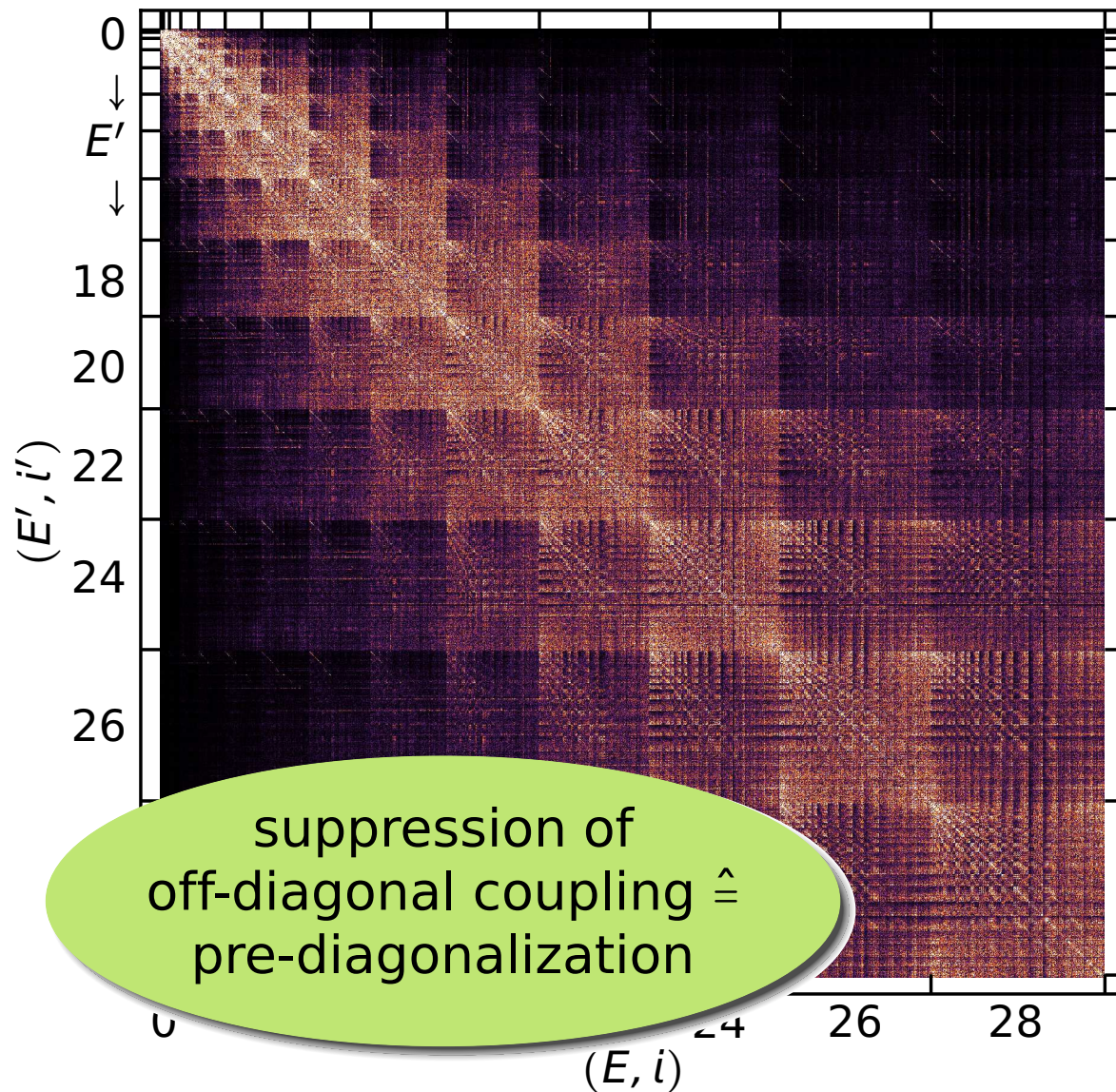
$$J^\pi = \frac{1}{2}^+, T = \frac{1}{2}, h\Omega = 28 \text{ MeV}$$

NCSM ground state ${}^3\text{H}$



SRG Evolution in Three-Body Space

3B-Jacobi HO matrix elements

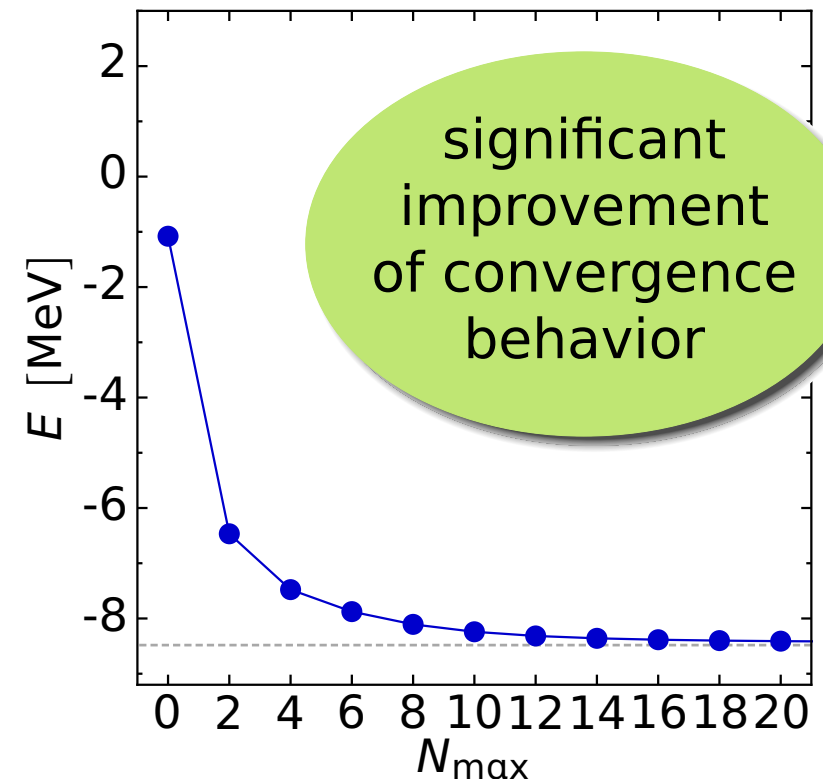


$$\alpha = 0.320 \text{ fm}^4$$

$$\Lambda = 1.33 \text{ fm}^{-1}$$

$$J^\pi = \frac{1}{2}^+, T = \frac{1}{2}, h\Omega = 28 \text{ MeV}$$

NCSM ground state ${}^3\text{H}$



Hamiltonian in A-Body Space

- evolution **induces n -body contributions** $H_{\alpha}^{[n]}$ to Hamiltonian

$$H_{\alpha} = H_{\alpha}^{[1]} + H_{\alpha}^{[2]} + H_{\alpha}^{[3]} + H_{\alpha}^{[4]} + H_{\alpha}^{[5]} + \dots$$

- **truncation of cluster series** formally destroys unitarity and invariance of energy eigenvalues (independence of α)
- flow-parameter α provides **diagnostic tool** to assess neglected higher-order contributions

SRG-Evolved Hamiltonians

NN_{only}	use initial NN, keep evolved NN
NN + 3N_{ind}	use initial NN, keep evolved NN+3N
NN + 3N_{full}	use initial NN+3N, keep evolved NN+3N
NN + 3N_{full} + 4N_{ind}	use initial NN+3N, keep evolved NN+3N+4N

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No-Core Shell Model

Barrett, Vary, Navratil, Maris, Nogga, Roth,...

NCSM is one of the most powerful and universal exact ab-initio methods

- construct matrix representation of Hamiltonian using a **basis of HO Slater determinants** truncated w.r.t. HO excitation energy $N_{\max} \hbar \Omega$
- solve **large-scale eigenvalue problem** for a few extremal eigenvalues
- **all relevant observables** can be computed from the eigenstates
- range of applicability limited by **factorial growth** of basis with N_{\max} & A
- adaptive **importance truncation** extends the range of NCSM by reducing the model space to physically relevant states

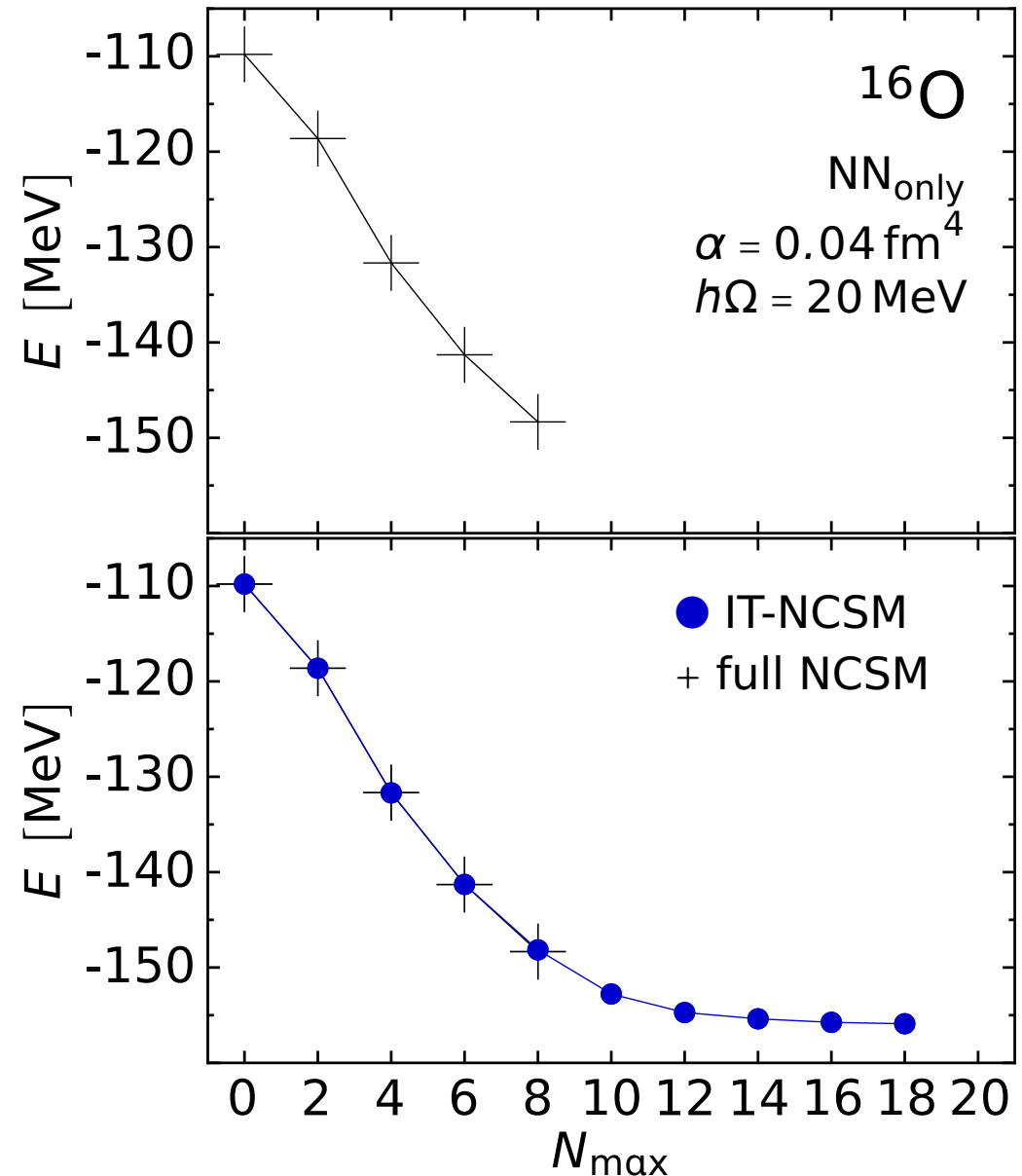
Importance Truncated NCSM

Roth, PRC 79, 064324 (2009); PRL 99, 092501 (2007)

- converged NCSM calculations essentially restricted to lower/mid p-shell
- full $N_{\max} = 10$ calculation for ^{16}O very difficult (basis dimension $> 10^{10}$)

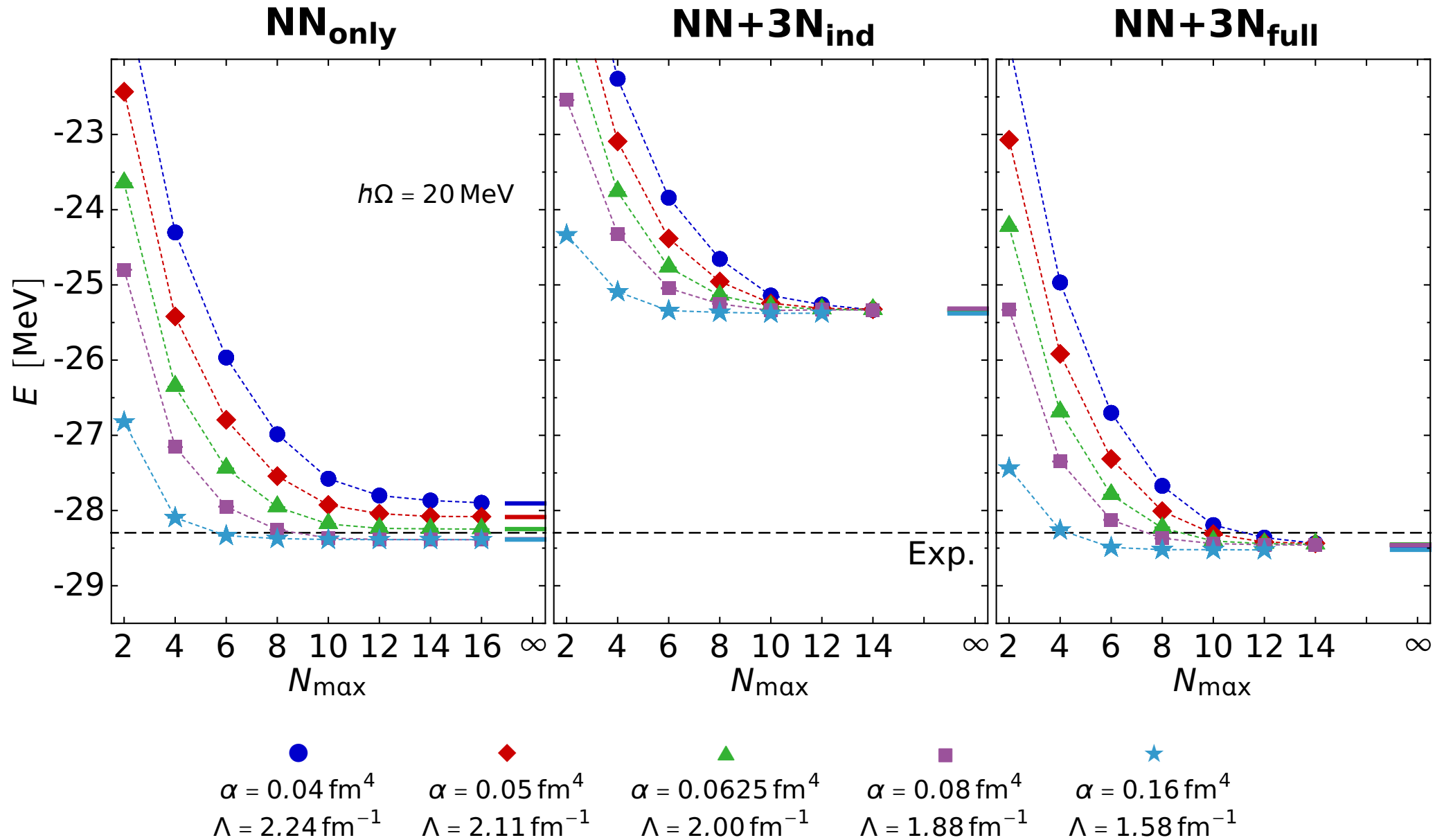
Importance Truncation

reduce model space to the relevant basis states using an **a priori importance measure** derived from MBPT



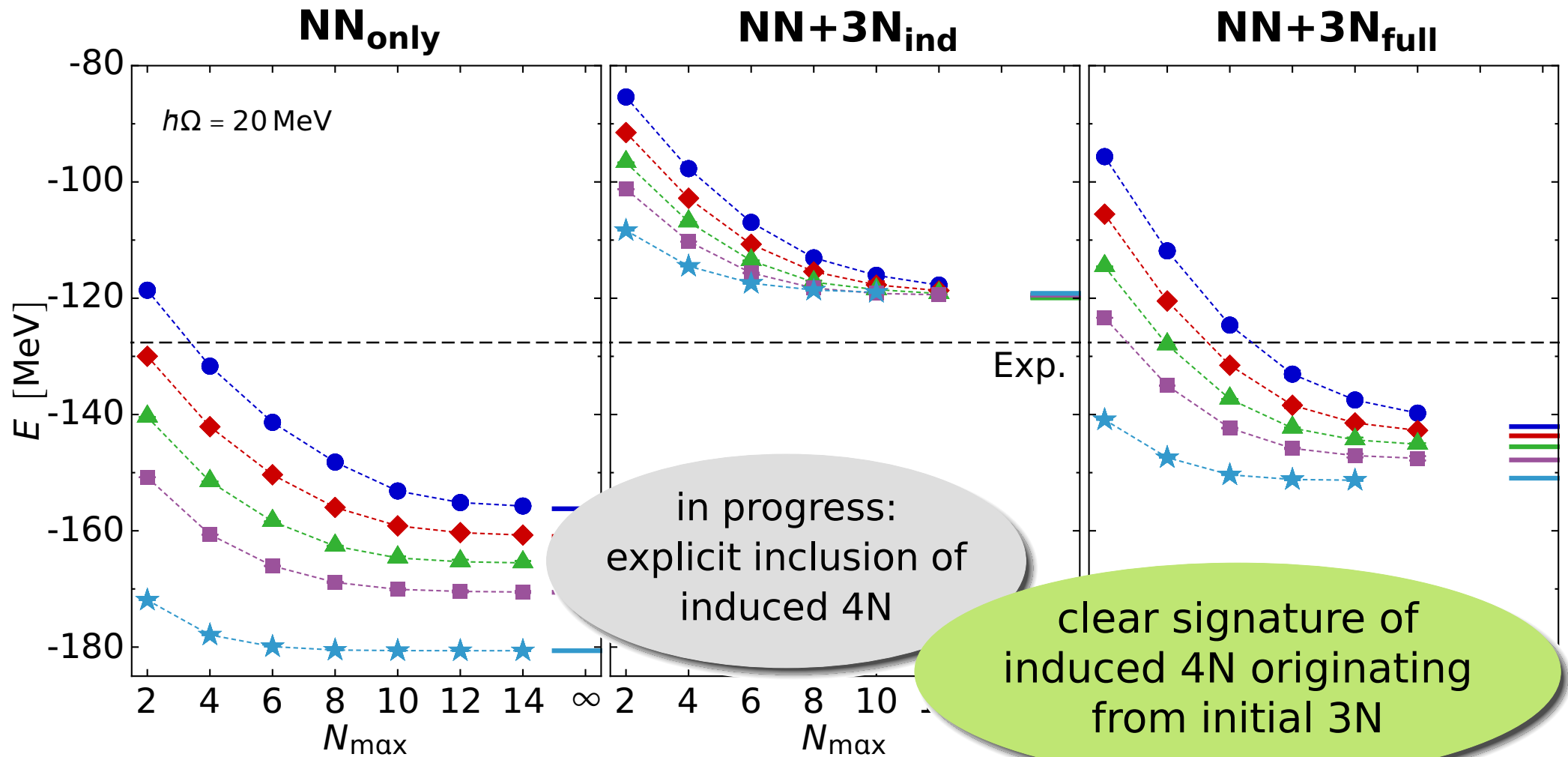
^4He : Ground-State Energies

Roth, et al; PRL 107, 072501 (2011)



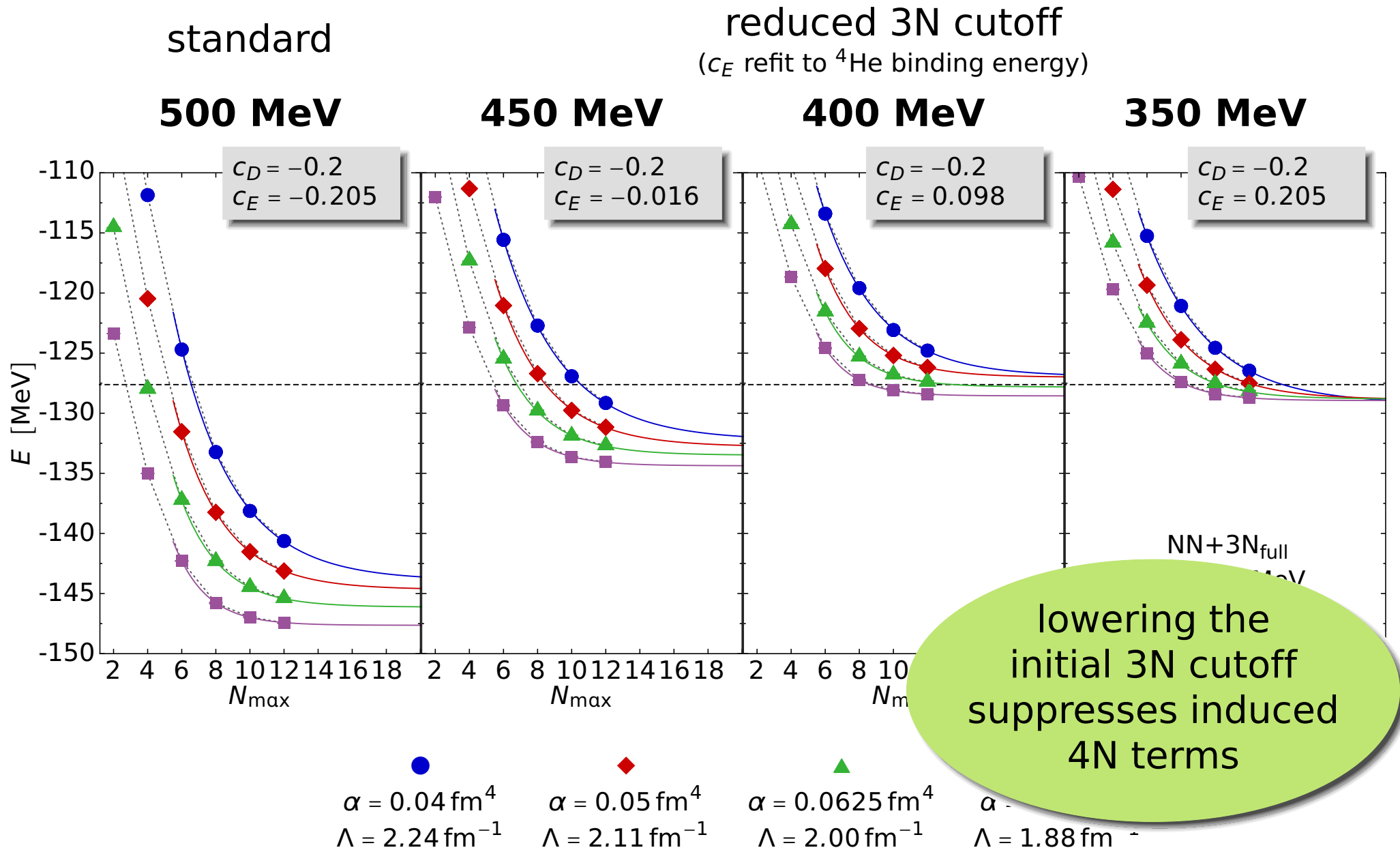
^{16}O : Ground-State Energies

Roth, et al; PRL 107, 072501 (2011)



●	◆	▲	■	★
$\alpha = 0.04 \text{ fm}^4$	$\alpha = 0.05 \text{ fm}^4$	$\alpha = 0.0625 \text{ fm}^4$	$\alpha = 0.08 \text{ fm}^4$	$\alpha = 0.16 \text{ fm}^4$
$\Lambda = 2.24 \text{ fm}^{-1}$	$\Lambda = 2.11 \text{ fm}^{-1}$	$\Lambda = 2.00 \text{ fm}^{-1}$	$\Lambda = 1.88 \text{ fm}^{-1}$	$\Lambda = 1.58 \text{ fm}^{-1}$

^{16}O : Lowering the Initial 3N Cutoff



Ground States of Oxygen Isotopes

- **oxygen isotopic chain** has received significant attention and documents the **rapid progress** over the past years

Otsuka, Suzuki, Holt, Schwenk, Akaishi, PRL 105, 032501 (2010)

- 2010: **shell-model calculations** with 3N effects highlighting the role of 3N interaction for drip line physics

Hagen, Hjorth-Jensen, Jansen, Machleidt, Papenbrock, PRL 108, 242501 (2012)

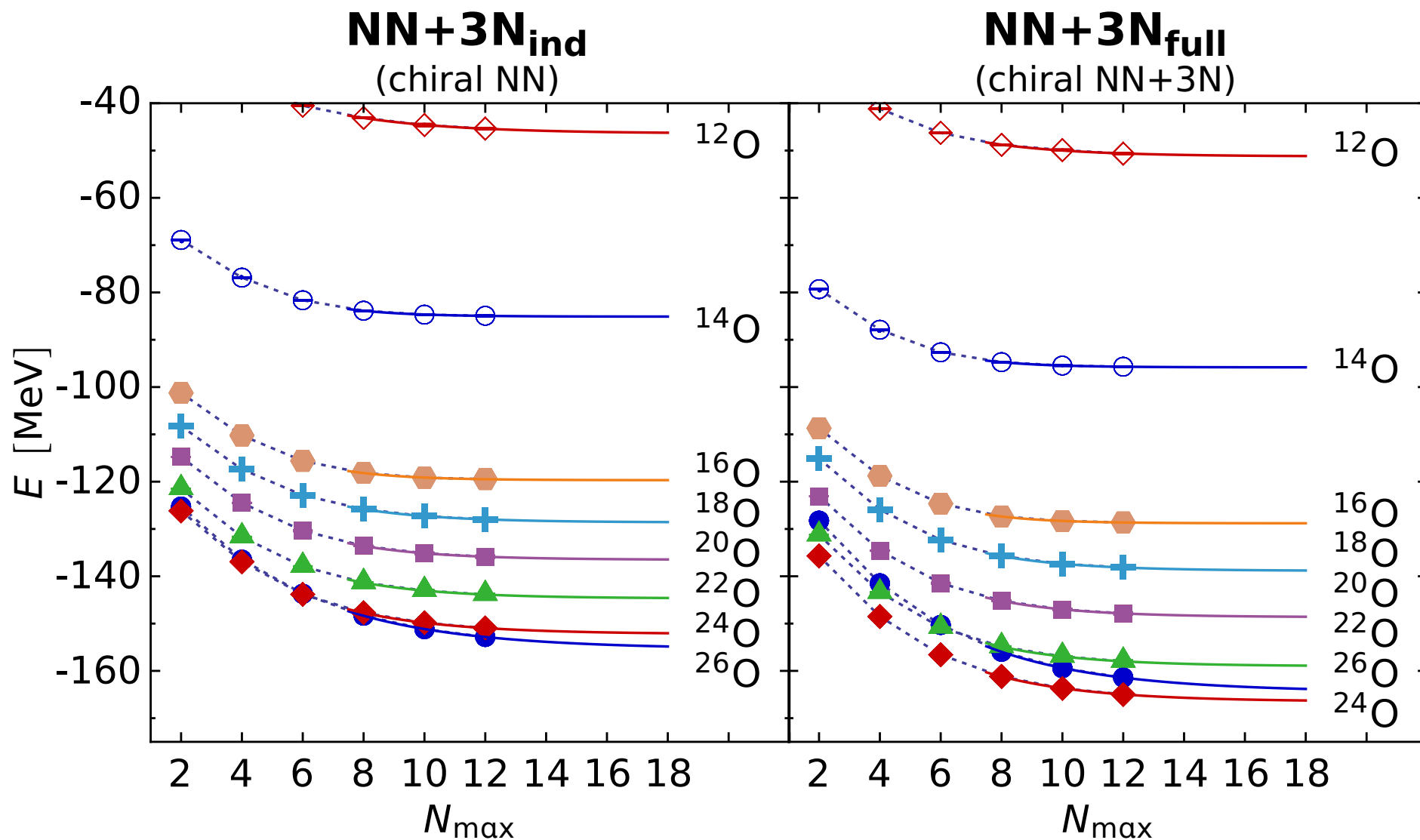
- 2012: **coupled-cluster calculations** with phenomenological two-body correction simulating chiral 3N forces

Hergert, Binder, Calci, Langhammer, Roth, PRL 110, 242501 (2013)

- 2013: **ab initio IT-NCSM** with explicit chiral 3N interactions...

Ground States of Oxygen Isotopes

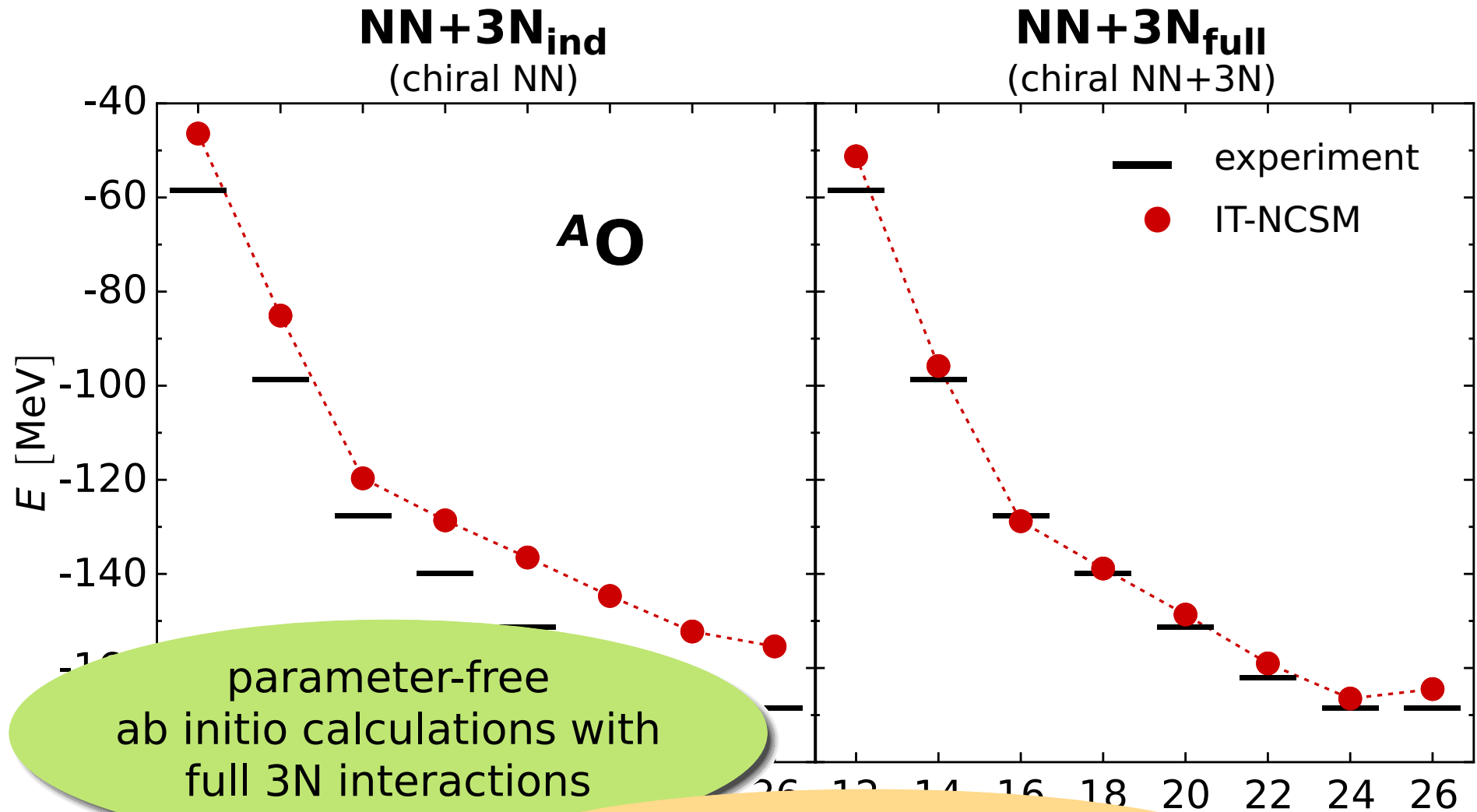
Hergert et al., PRL 110, 242501 (2013)



$\Lambda_{3N} = 400 \text{ MeV}$, $\alpha = 0.08 \text{ fm}^4$, $E_{3\text{max}} = 14$, optimal $\hbar\Omega$

Ground States of Oxygen Isotopes

Hergert et al., PRL 110, 242501 (2013)



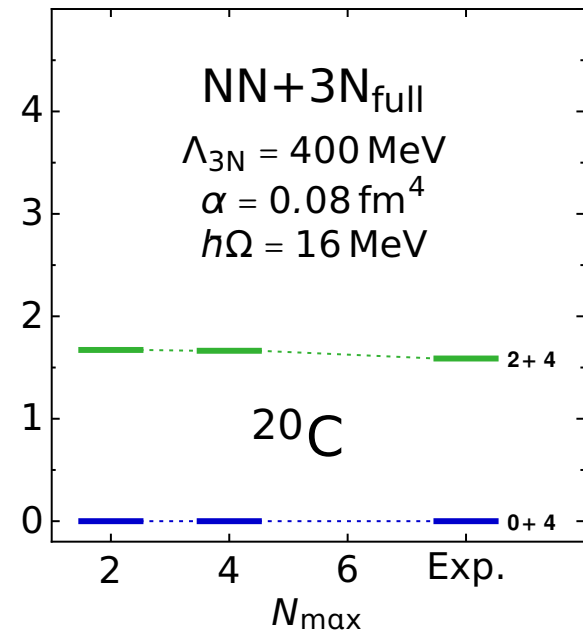
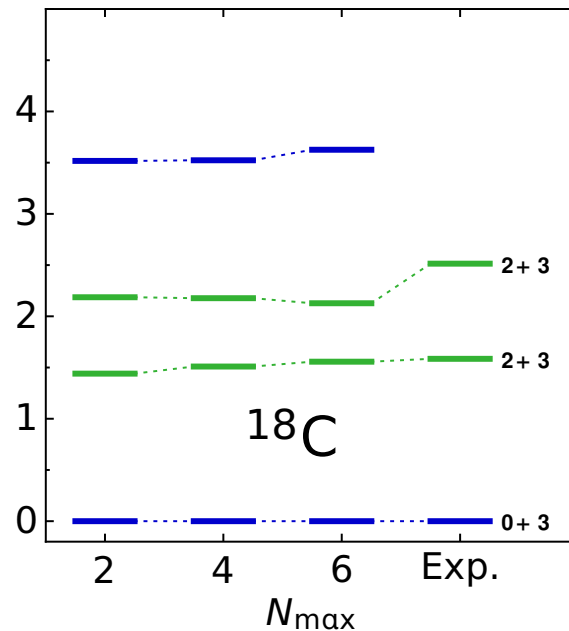
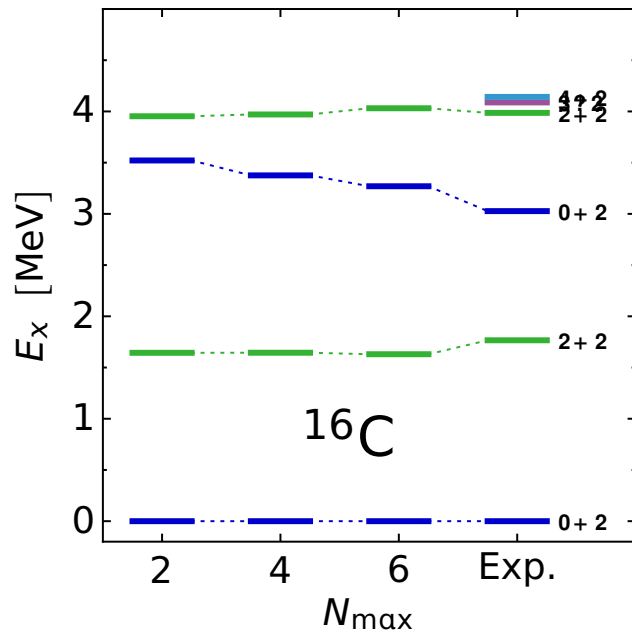
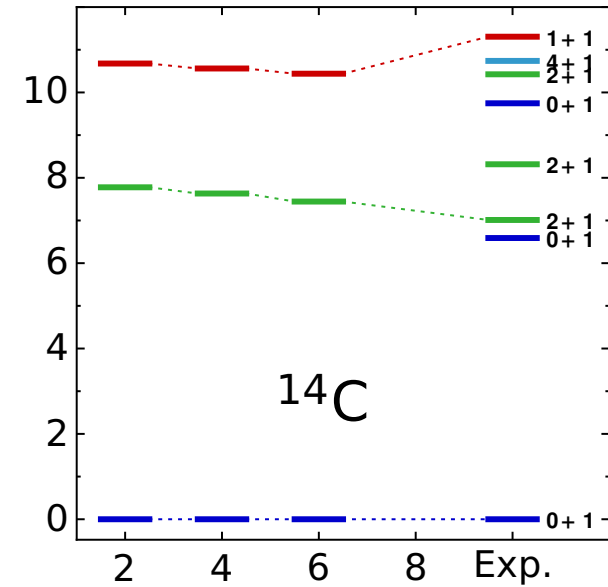
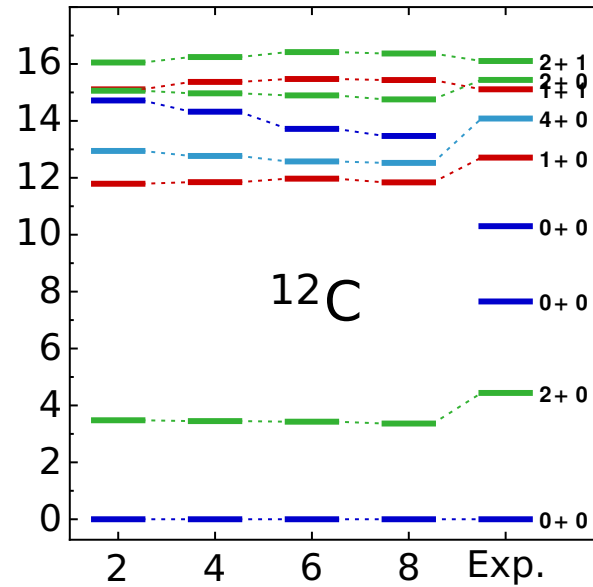
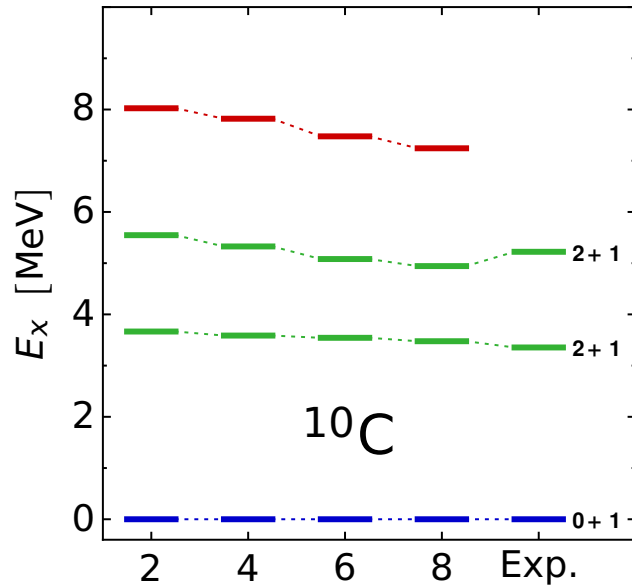
parameter-free
ab initio calculations with
full 3N interactions

highlights predictive power
of chiral NN+3N Hamiltonians

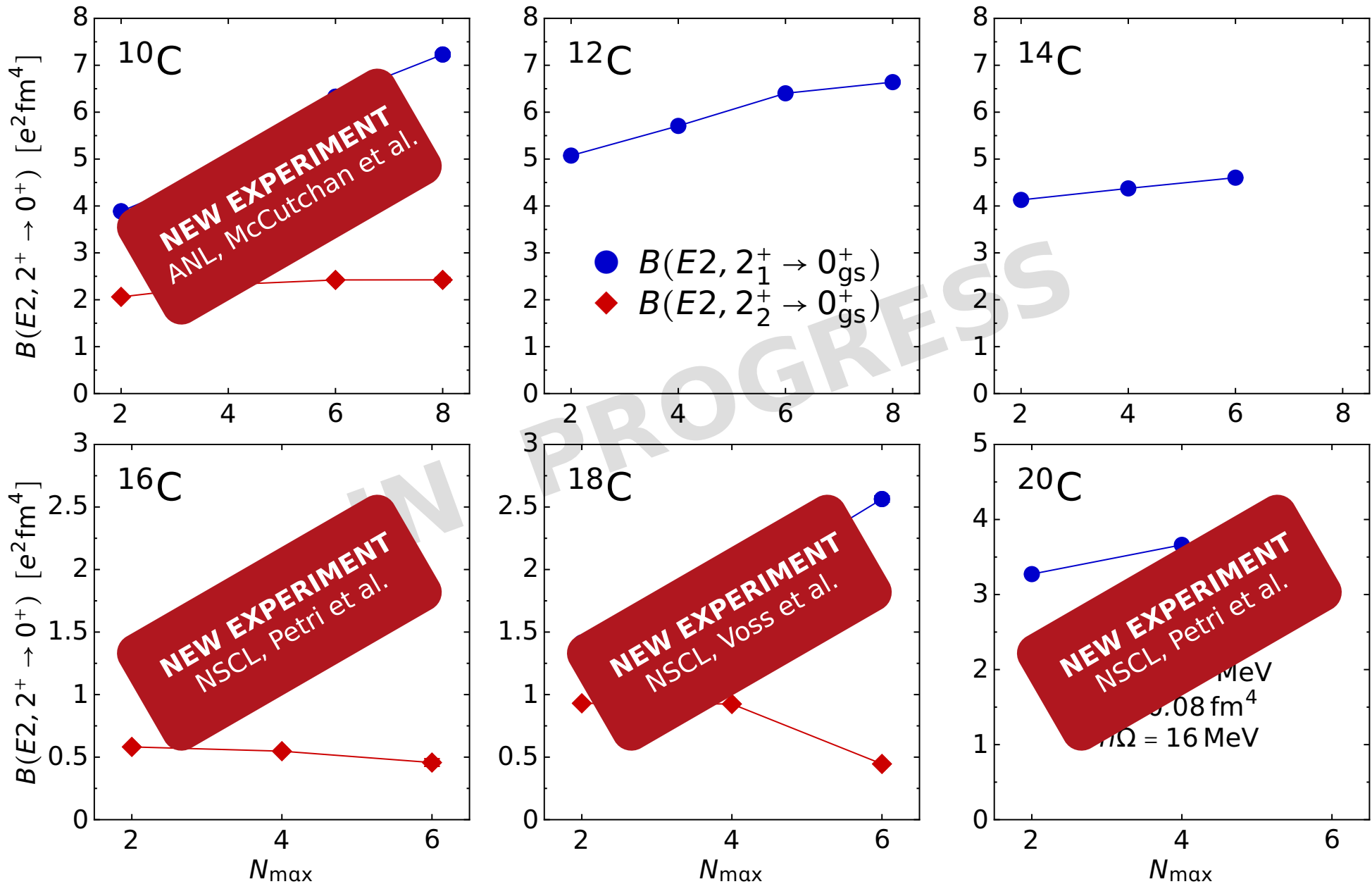
$\Lambda_{3N} = 400 \text{ MeV}$, $\alpha = 0.1$, $\mu = 0.015 \text{ fm}^{-1}$, $\beta = 0.015 \text{ fm}^{-1}$, $\gamma = 0.015 \text{ fm}^{-1}$, $\delta = 0.015 \text{ fm}^{-1}$, $\epsilon = 0.015 \text{ fm}^{-1}$, $\zeta = 0.015 \text{ fm}^{-1}$, $\eta = 0.015 \text{ fm}^{-1}$, $\theta = 0.015 \text{ fm}^{-1}$, $\iota = 0.015 \text{ fm}^{-1}$, $\kappa = 0.015 \text{ fm}^{-1}$, $\lambda = 0.015 \text{ fm}^{-1}$, $\mu = 0.015 \text{ fm}^{-1}$, $\nu = 0.015 \text{ fm}^{-1}$, $\xi = 0.015 \text{ fm}^{-1}$, $\omicron = 0.015 \text{ fm}^{-1}$, $\pi = 0.015 \text{ fm}^{-1}$, $\rho = 0.015 \text{ fm}^{-1}$, $\sigma = 0.015 \text{ fm}^{-1}$, $\tau = 0.015 \text{ fm}^{-1}$, $\upsilon = 0.015 \text{ fm}^{-1}$, $\phi = 0.015 \text{ fm}^{-1}$, $\chi = 0.015 \text{ fm}^{-1}$, $\psi = 0.015 \text{ fm}^{-1}$, $\omega = 0.015 \text{ fm}^{-1}$, $\eta = 0.015 \text{ fm}^{-1}$, $\theta = 0.015 \text{ fm}^{-1}$, $\iota = 0.015 \text{ fm}^{-1}$, $\kappa = 0.015 \text{ fm}^{-1}$, $\lambda = 0.015 \text{ fm}^{-1}$, $\mu = 0.015 \text{ fm}^{-1}$, $\nu = 0.015 \text{ fm}^{-1}$, $\xi = 0.015 \text{ fm}^{-1}$, $\omicron = 0.015 \text{ fm}^{-1}$, $\pi = 0.015 \text{ fm}^{-1}$, $\rho = 0.015 \text{ fm}^{-1}$, $\sigma = 0.015 \text{ fm}^{-1}$, $\tau = 0.015 \text{ fm}^{-1}$, $\upsilon = 0.015 \text{ fm}^{-1}$, $\phi = 0.015 \text{ fm}^{-1}$, $\chi = 0.015 \text{ fm}^{-1}$, $\psi = 0.015 \text{ fm}^{-1}$, $\omega = 0.015 \text{ fm}^{-1}$

Spectroscopy of Carbon Isotopes

Forsen et al., JPG 40, 055105 (2013); Roth et al., in prep.



Spectroscopy of Carbon Isotopes



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Frontier: Medium-Mass Nuclei

advent of novel ab initio many-body approaches applicable in the medium-mass regime

Hagen, Papenbrock, Dean, Piecuch, Binder,...

- **coupled-cluster theory**: ground-state parametrized by exponential wave operator applied to single-determinant reference state

- truncation at doubles level (CCSD) plus triple excitations (CCSD(T))
- equations of motion for excited states

- **in-medium SRG**: many-body perturbation theory (MBPT) with normal-ordering and renormalization

- normal-ordering of Hamiltonian truncated at two-body level
- both close to ground states; excitations via EOM or SM

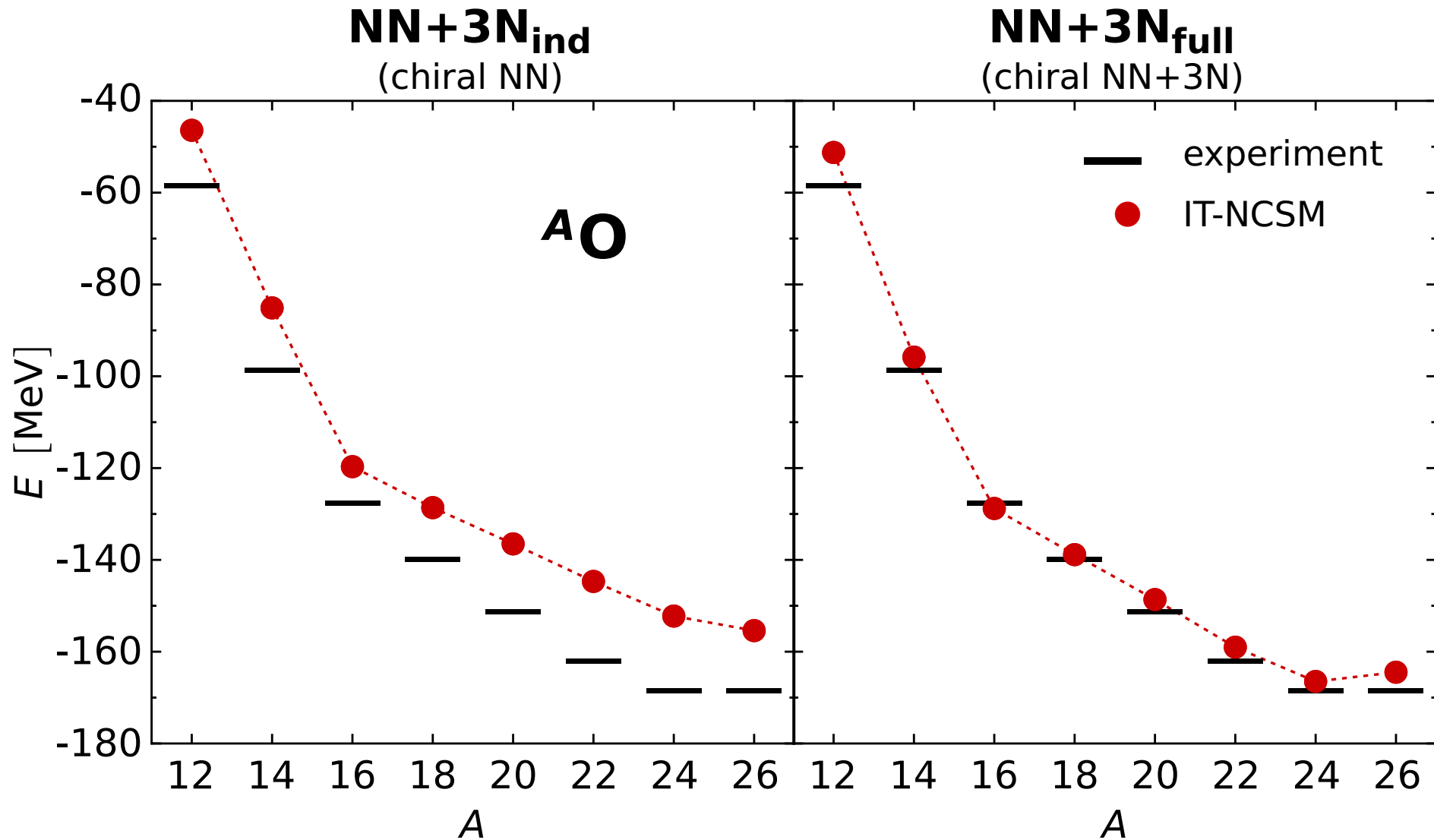
Barbieri, Soma, Duguet,...

- self-consistent Green's function approaches and others...

controlling and quantifying the uncertainties due to various truncations is a major challenge

Ground States of Oxygen Isotopes

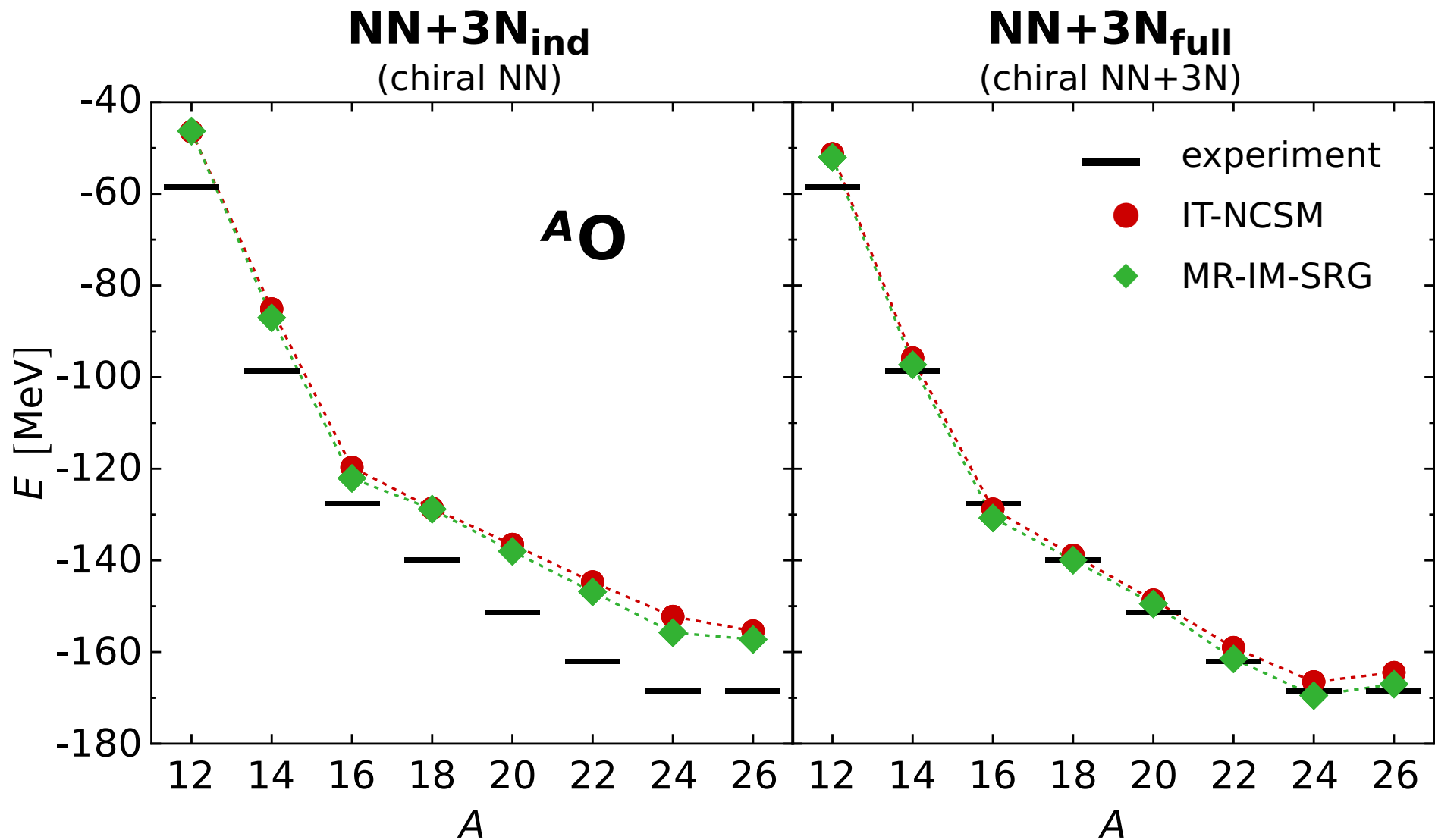
Hergert et al., PRL 110, 242501 (2013)



$\Lambda_{3N} = 400$ MeV, $\alpha = 0.08$ fm⁴, $E_{3\text{max}} = 14$, optimal $\hbar\Omega$

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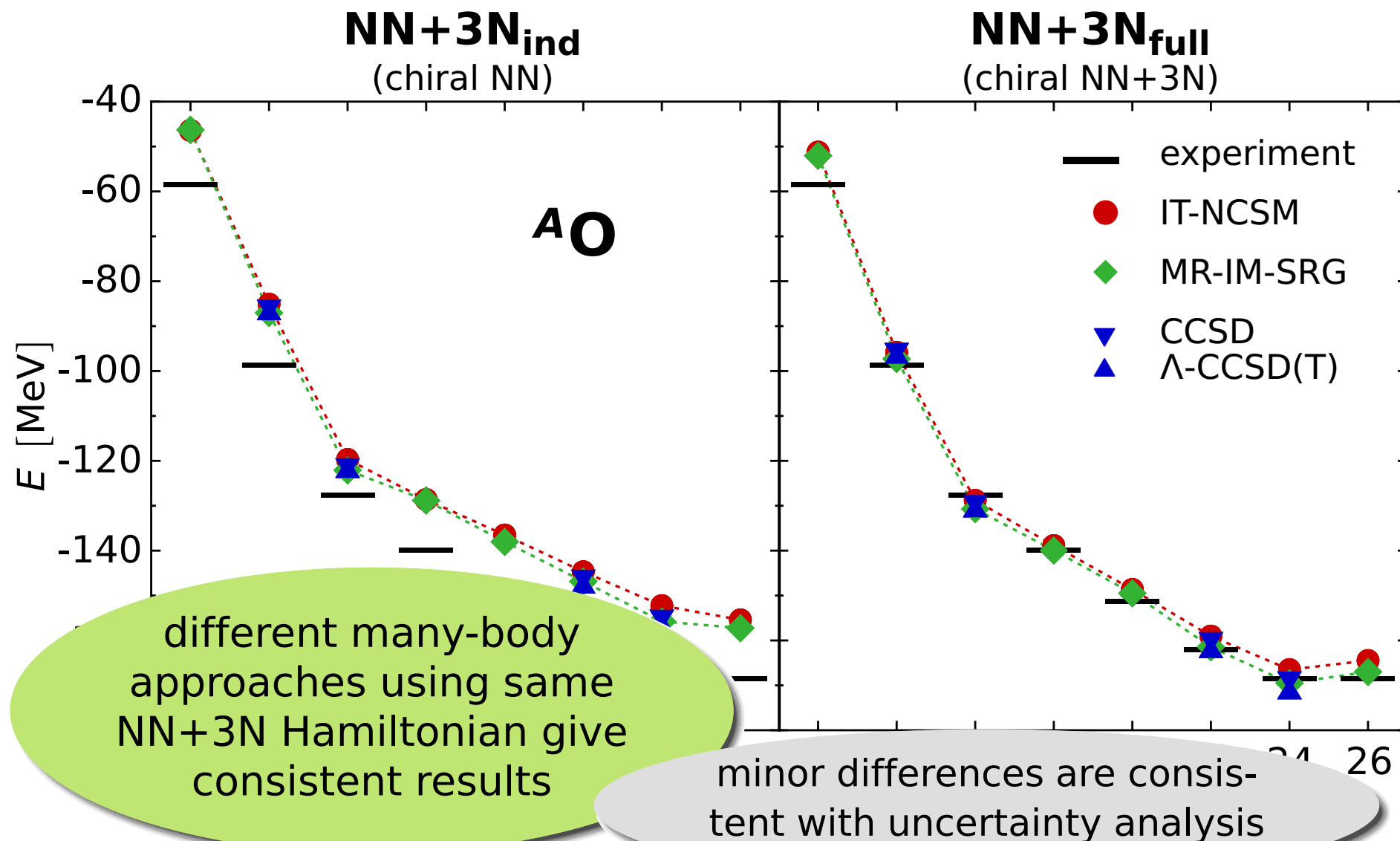
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$\Lambda_{3N} = 400 \text{ MeV}$, $\alpha = 0.08 \text{ fm}^4$, $E_{3\text{max}} = 14$, optimal $h\Omega$

Ground States of Oxygen Isotopes

Hergert et al., PRL 110, 242501 (2013)



$\Lambda_{3N} = 400$ MeV, $\alpha = 0.08$ fm⁻¹, $E_{3\text{max}} = 1\tau$, spatial Ω

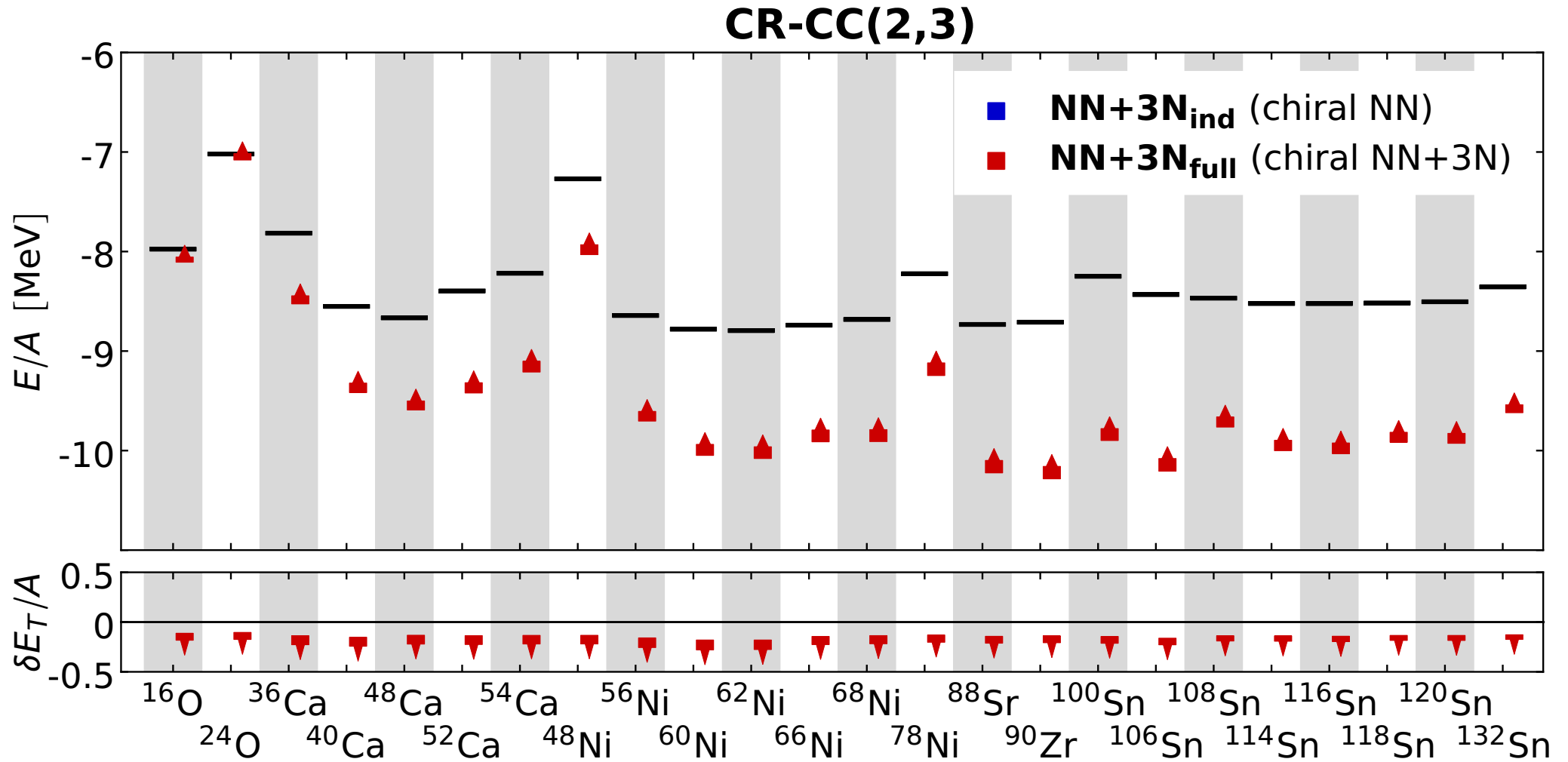
Towards Heavy Nuclei - Ab Initio ?

Roth, et al., PRL 109, 052501 (2012); Binder et al., PRC 87, 021303(R) (2013); PRC 88, 054319 (2013); arXiv:1312.5685 (2013)

- calculations for medium-mass and heavy nuclei are **computationally feasible** with CC or IM-SRG
- however, many of the **technical truncations** that are good in light nuclei **fail for heavier systems**
- we **analysed and improved** all of these truncations...
- **2% residual uncertainty** of the many-body approach for $A \lesssim 130$

Towards Heavy Nuclei - Ab Initio !

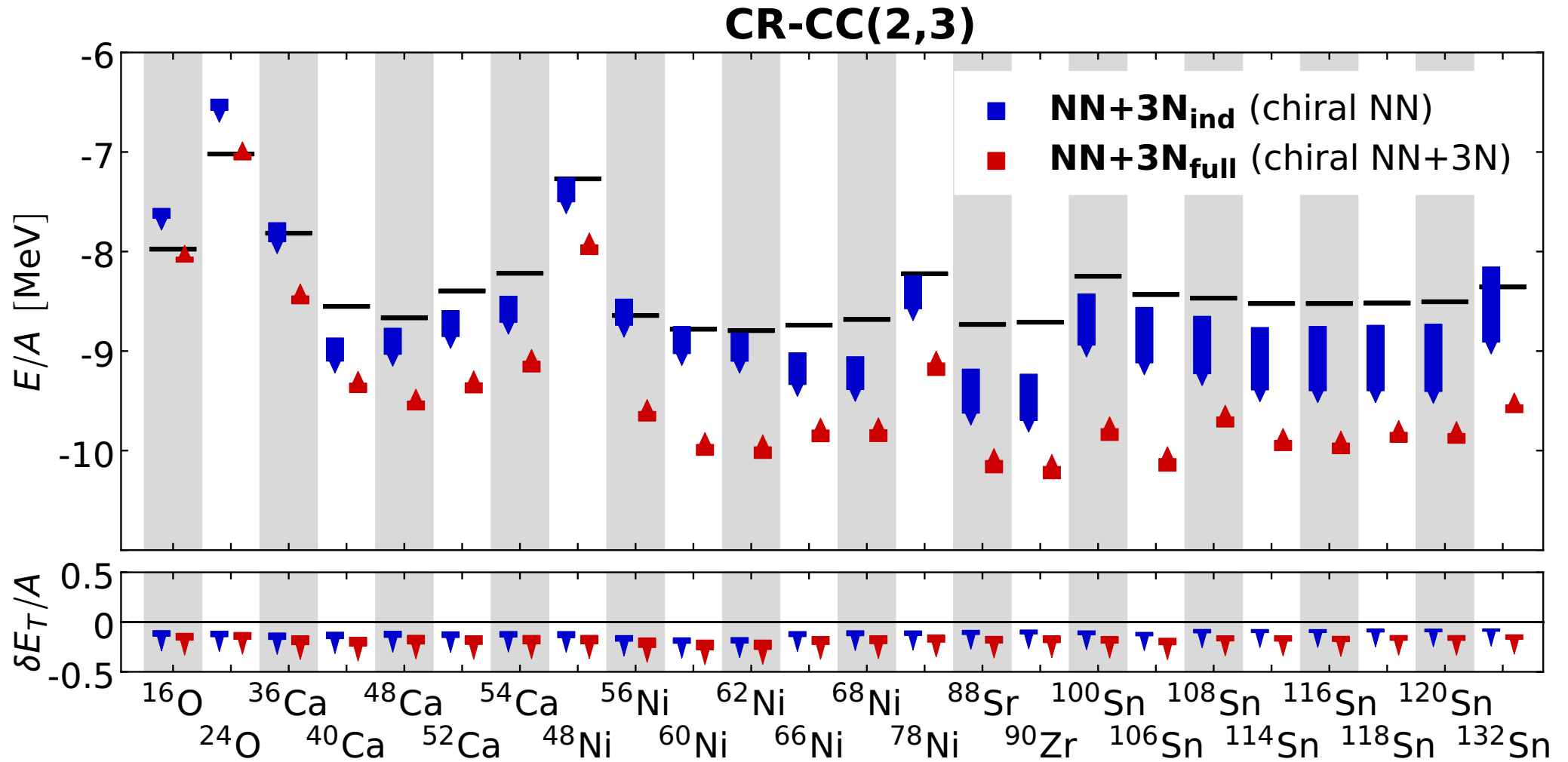
Binder et al., arXiv:1312.5685 (2013)



$\Lambda_{3N} = 400$ MeV, $\alpha = 0.08 \rightarrow 0.04 \text{ fm}^4$, $E_{3\text{max}} = 18$, optimal $h\Omega$

Towards Heavy Nuclei - Ab Initio !

Binder et al., arXiv:1312.5685 (2013)



$$\Lambda_{3\text{N}} = 400 \text{ MeV}, \quad \alpha = 0.08 \rightarrow 0.04 \text{ fm}^4, \quad E_{3\text{max}} = 18, \quad \text{optimal } h\Omega$$

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Ab Initio *Hyper*-Nuclear Structure

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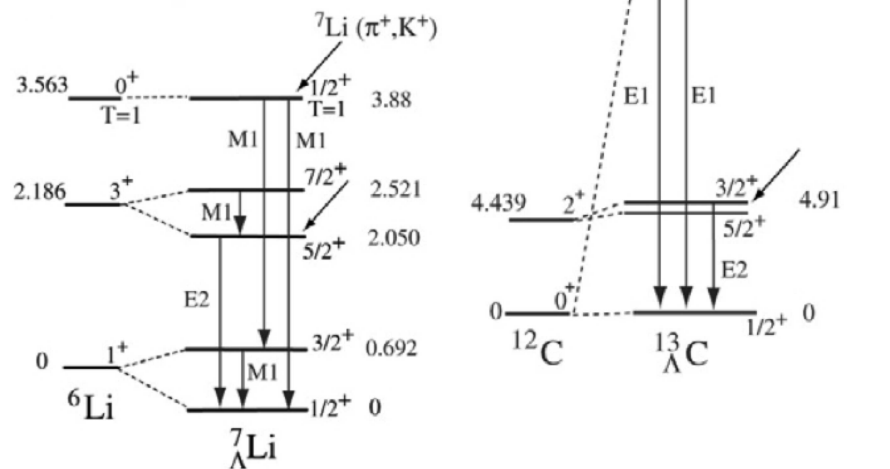
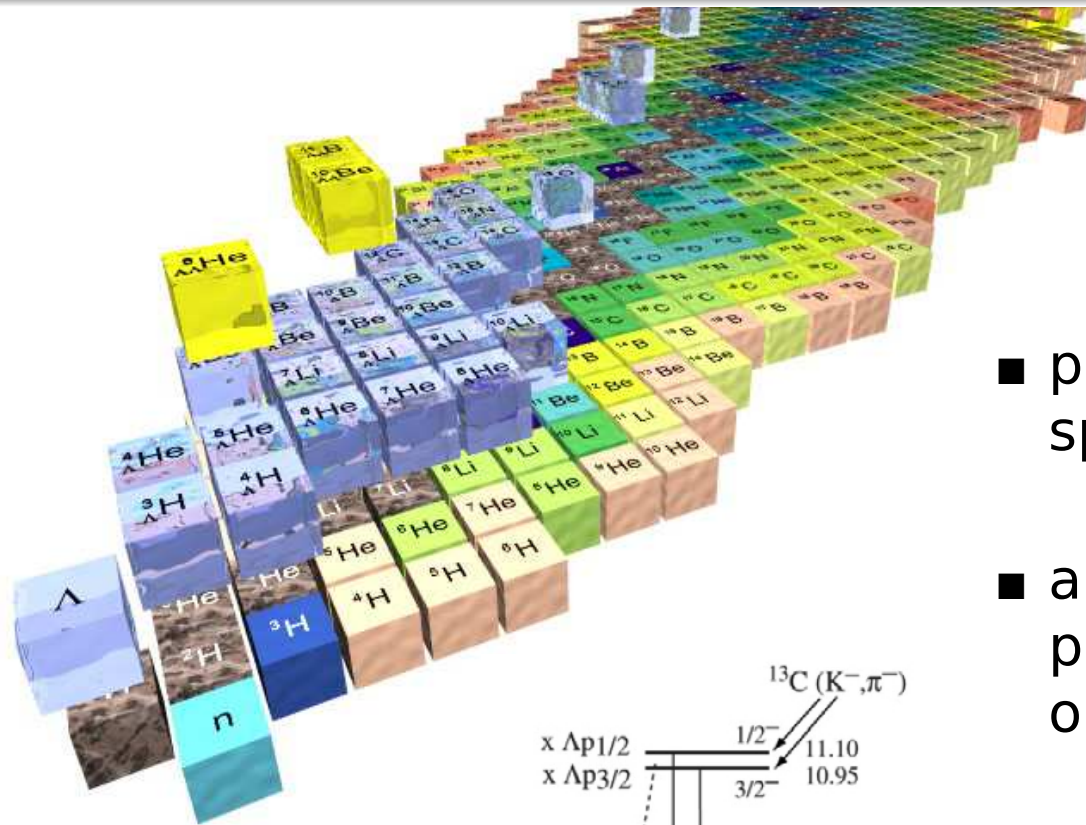
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Ab Initio Hyper-Nuclear Structure



- precise data on ground states & spectroscopy of hyper-nuclei
- ab initio few-body ($A \lesssim 4$) and phenomenological shell model or cluster calculations
- chiral YN & YY interactions at (N)LO are available
- constrain YN & YY interaction by ab initio hyper-nuclear structure calculations

Ab Initio Toolbox

■ Hamiltonian from chiral EFT

- NN+3N: standard chiral Hamiltonian (Entem&Machleidt, Navrátil)
- YN: LO chiral interaction (Haidenbauer et al.), NLO in progress

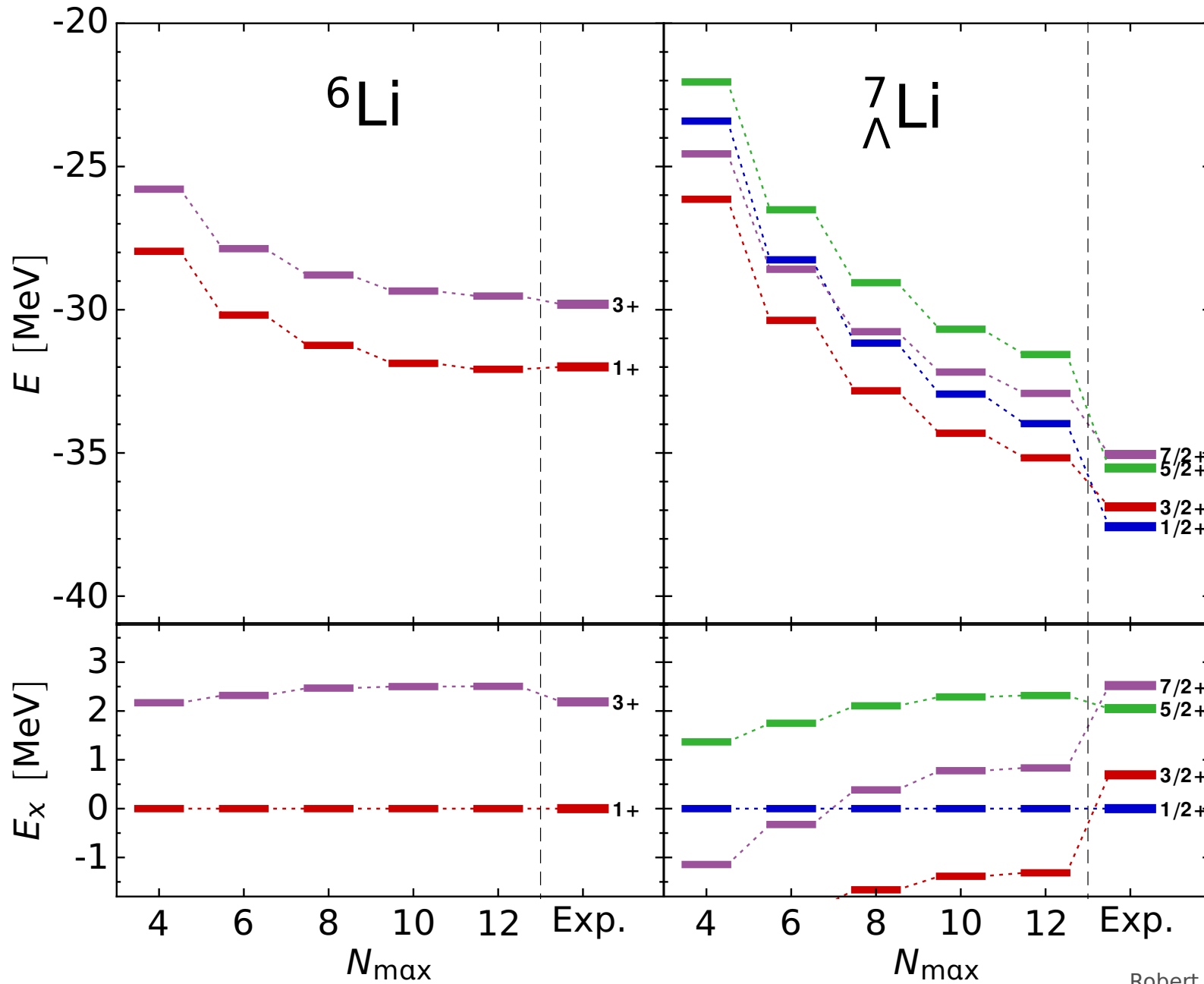
■ Similarity Renormalization Group

- consistent SRG-evolution of NN, 3N, YN interactions
- using particle basis and including $\Lambda\Sigma$ -coupling (larger matrices)
- Λ - Σ mass difference and $p\Sigma^\pm$ Coulomb included consistently

■ Importance Truncated No-Core Shell Model

- include explicit $(p, n, \Lambda, \Sigma^+, \Sigma^0, \Sigma^-)$ with physical masses
- larger model spaces easily tractable with importance truncation
- all p-shell single- Λ hypernuclei are accessible

Application: ${}^7_{\Lambda}\text{Li}$



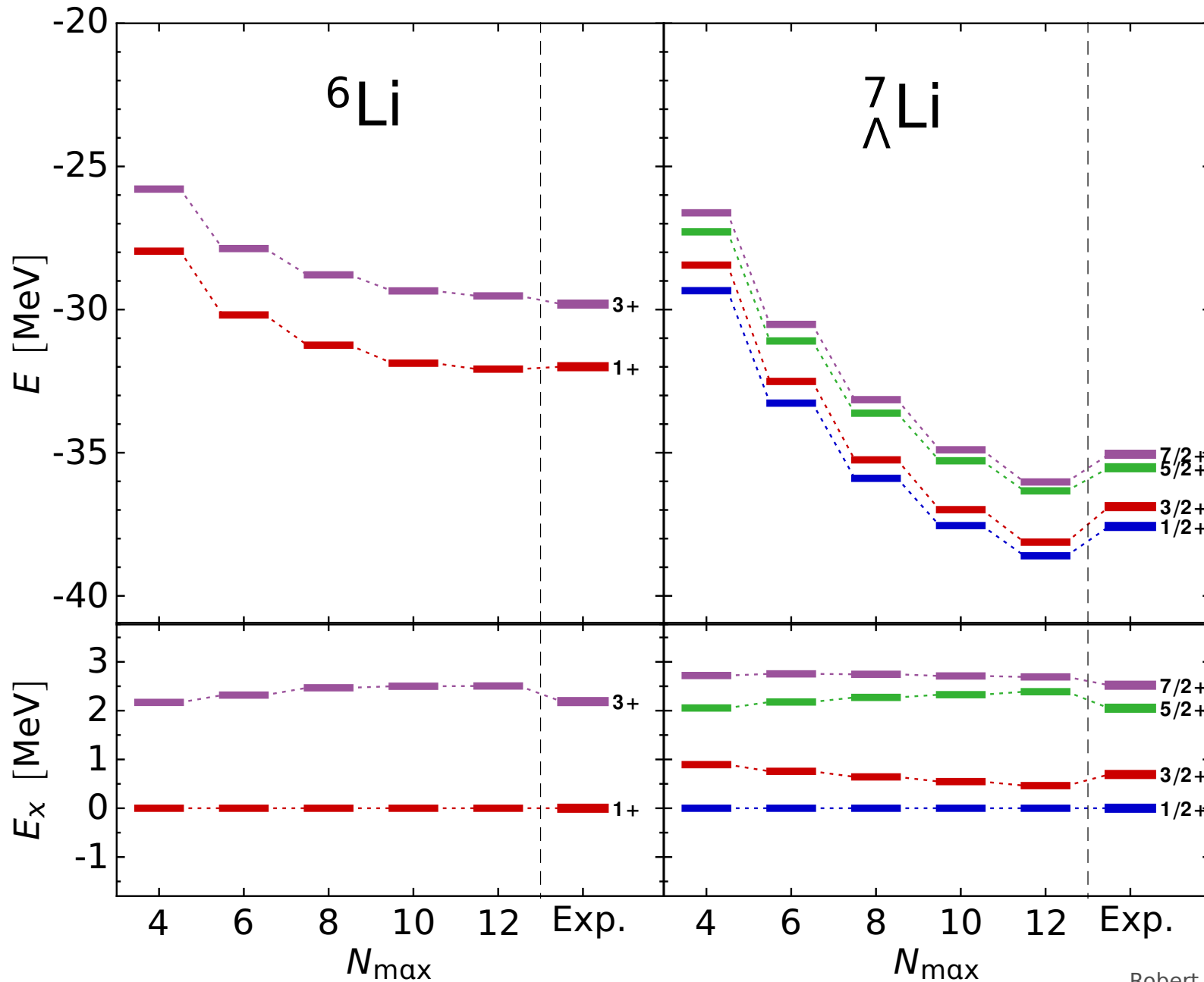
NN @ N3LO
 $\Lambda_{NN} = 500 \text{ MeV}$
 Entem&Machleidt

3N @ N2LO
 $\Lambda_{3N} = 500 \text{ MeV}$
 Navratil
 A = 3 fit

Jülich'04
 Haidenbauer et al.
 scatt. & hypertriton

$\alpha_{NN} = 0.08 \text{ fm}^4$
 $\alpha_{YN} = 0.00 \text{ fm}^4$
 $h\Omega = 20 \text{ MeV}$

Application: ${}^7_{\Lambda}\text{Li}$



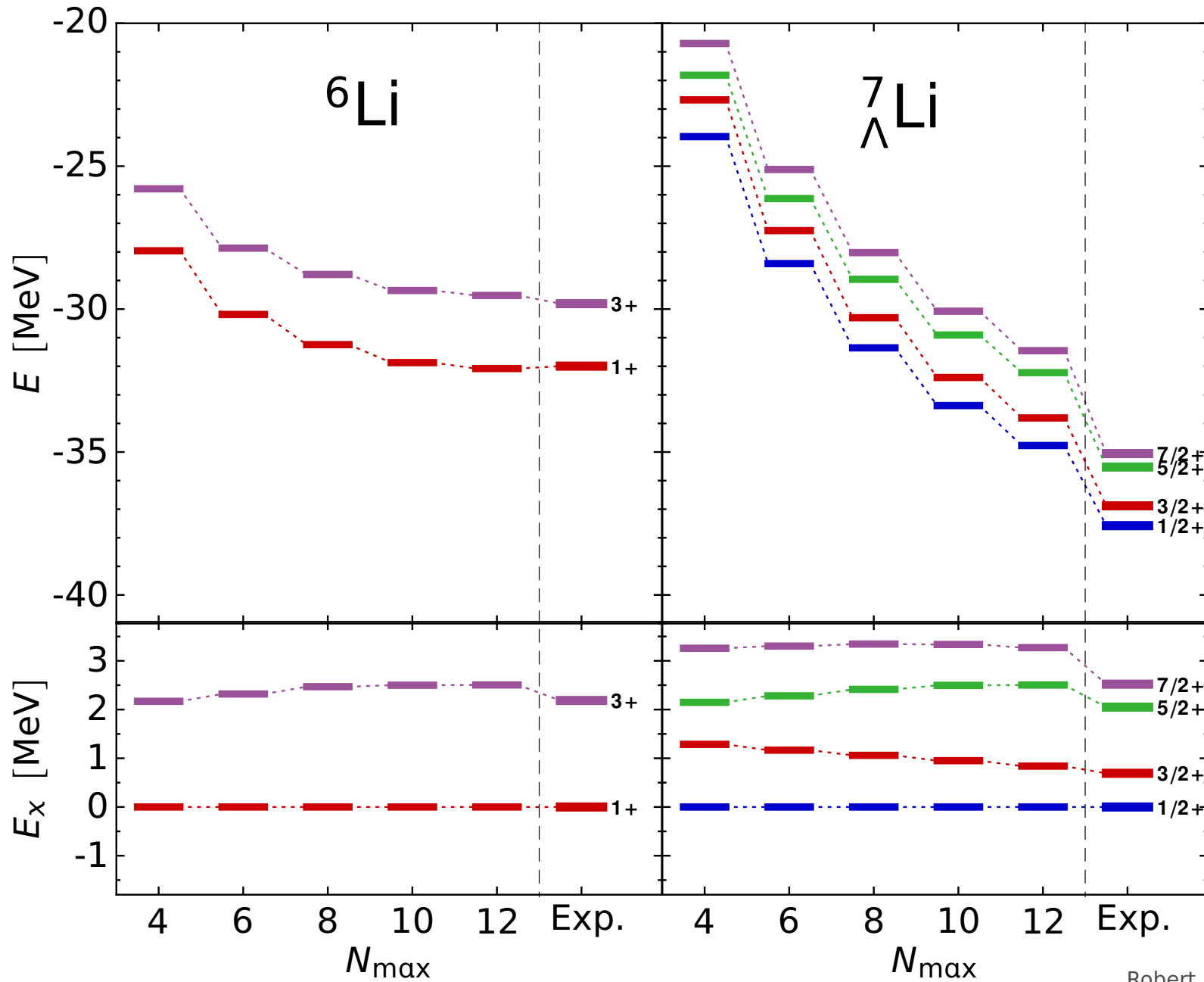
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YN @ LO
 $\Lambda_{YN} = 600 \text{ MeV}$
 Haidenbauer et al.
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$\alpha_{NN} = 0.08 \text{ fm}^4$
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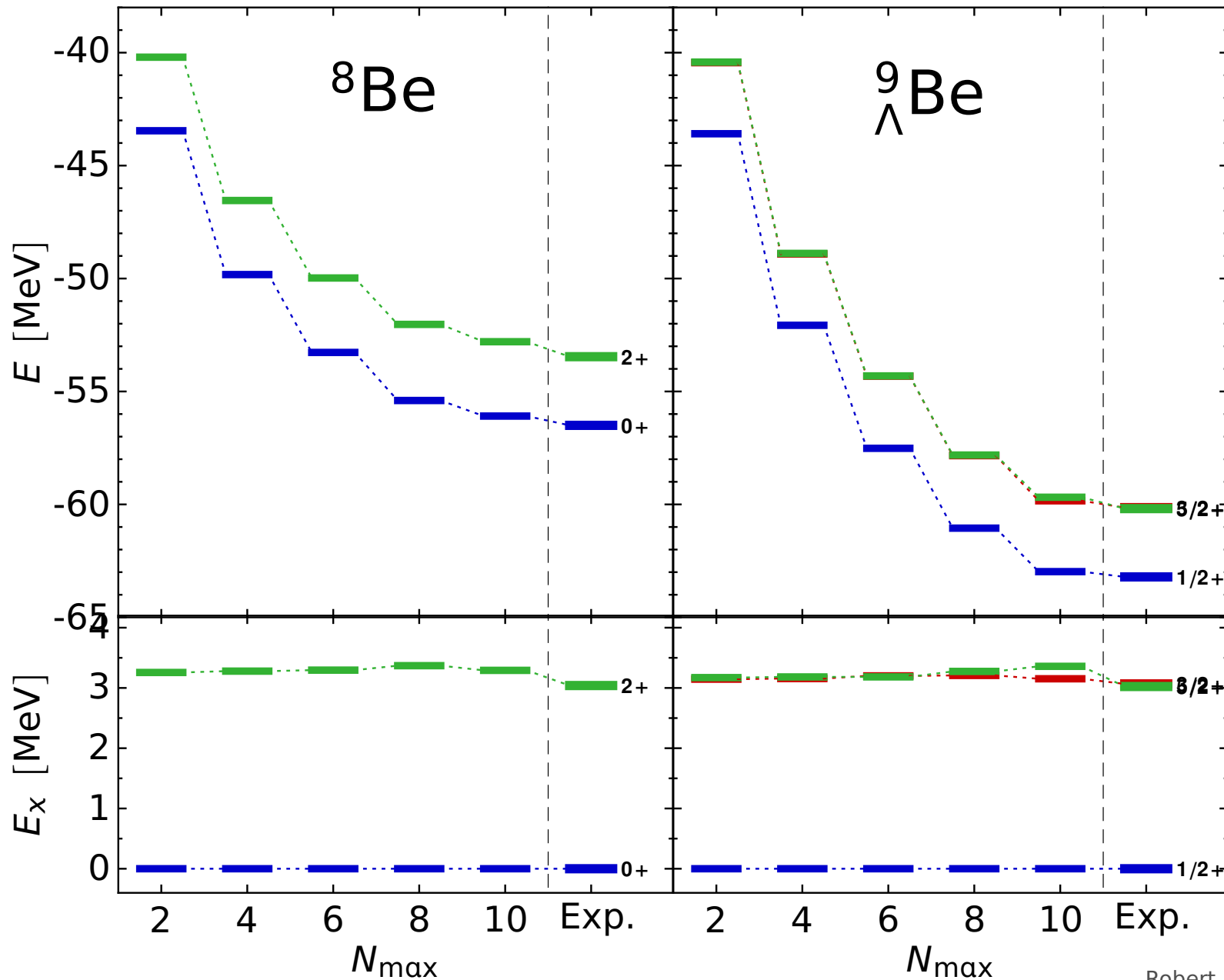
NN @ N3LO
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3N @ N2LO
 $\Lambda_{3N} = 500 \text{ MeV}$
 Navratil
 A = 3 fit

YN @ LO
 $\Lambda_{YN} = 700 \text{ MeV}$
 Haidenbauer et al.
 scatt. & hypertriton

$\alpha_{NN} = 0.08 \text{ fm}^4$
 $\alpha_{YN} = 0.00 \text{ fm}^4$
 $h\Omega = 20 \text{ MeV}$

Application: ${}^9_{\Lambda}\text{Be}$



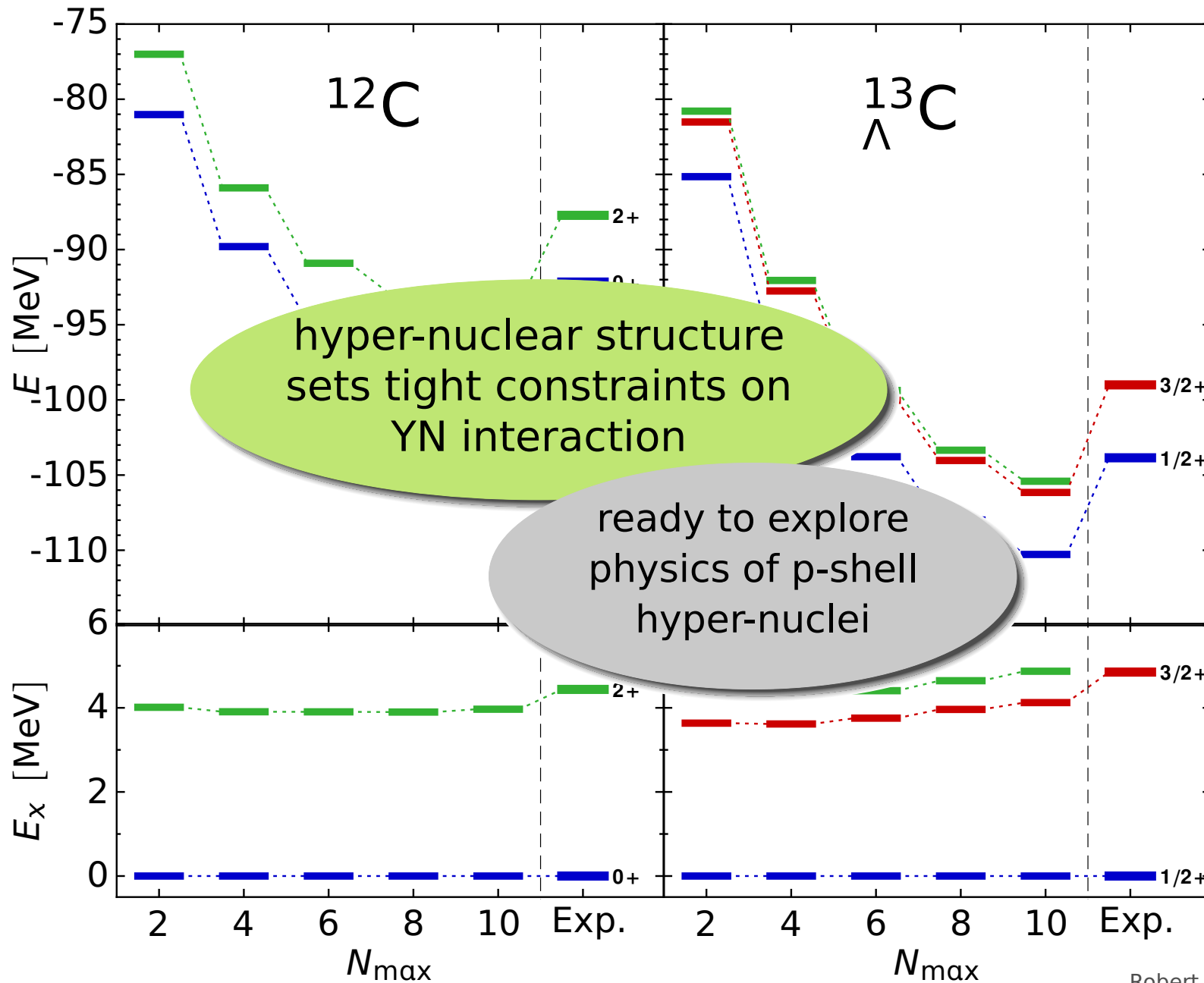
NN @ N3LO
 $\Lambda_{NN} = 500$ MeV
 Entem&Machleidt

3N @ N2LO
 $\Lambda_{3N} = 500$ MeV
 Navratil
 A = 3 fit

YN @ LO
 $\Lambda_{YN} = 600$ MeV
 Haidenbauer et al.
 scatt. & hypertriton

$\alpha_{NN} = 0.08 \text{ fm}^4$
 $\alpha_{YN} = 0.00 \text{ fm}^4$
 $h\Omega = 20$ MeV

Application: ${}_{\Lambda}^{13}\text{C}$



NN @ N3LO
 $\Lambda_{NN} = 500 \text{ MeV}$
 Entem&Machleidt

3N @ N2LO
 $\Lambda_{3N} = 500 \text{ MeV}$
 Navratil
 A = 3 fit

YN @ LO
 $\Lambda_{YN} = 600 \text{ MeV}$
 Haidenbauer et al.
 scatt. & hypertriton

$\alpha_{NN} = 0.08 \text{ fm}^4$
 $\alpha_{YN} = 0.00 \text{ fm}^4$
 $h\Omega = 20 \text{ MeV}$

■ **ab initio theory is entering new territory...**

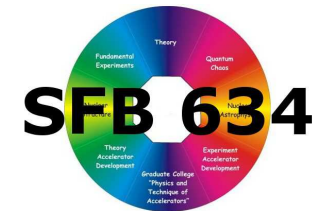
- **QCD frontier**
nuclear structure connected systematically to QCD via chiral EFT
- **precision frontier**
precision spectroscopy of light nuclei, including current contribution
- **mass frontier**
ab initio calculations up to heavy nuclei with quantified uncertainties
- **open-shell frontier**
extend to medium-mass open-shell nuclei and their excitation spectrum
- **continuum frontier**
include continuum effects and scattering observables consistently
- **strangeness frontier**
ab initio predictions for hyper-nuclear structure & spectroscopy

...providing a coherent theoretical framework for nuclear structure & reaction calculations

Epilogue

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