



Technische Universitaet Muenchen Excellence Cluster "Origin and Structure of the Universe" <u>https://www.e12.ph.tum.de/groups/kcluster</u> Utrecht University WS2014-15

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- K_{s}^{0} in p+A reactions: detailed studies of the production and interaction
- Λ-p femtoscopy in p+A collisions: a new method to test the Hyperon-Nucleon interactions



NIVERSITÄ



- Test-bed of the strong interaction in few body systems
- Strange quarks are intermediate between "light" and "heavy"
- -> Interplay between spontaneous and explicit chiral symmetry breaking in low energy QCD.
- Testing ground: K-N and K-N interactions





Hyperon Star





It all depends upon the Λ -N and Λ -NN interaction and whether or not it has a repulsive core This repulsive core could stiffen again the EOS allowing for heavy neutron stars







Kaons in the Nucleus



p+Nb, 3.5 GeV



Kaon interactions in nuclear medium:

- Elastic scattering
- Charge Exchange
- Inelastic reactions
- π-induced secondary reactions...





Data are interpreted with the GiBUU transport model

O. Buss et al., Phys. Rept. 512, 1 (2012) http://gibuu.physik.uni-giessen.de/GiBUU/





4-body states produced via Δ -resonances

9

- Final states with two pions (5-body) added to the model
- via NN $\rightarrow \Delta^{++}$ Y* K, Y* is $\Sigma(1385)$ or $\Lambda(1405)$.
- Good description of the elementary reference.





In-medium kaon potential



T. Gaitanos, K. Lapidus

ChPT potential, ~35 MeV ($\varrho = \varrho_0$, k=0)

$$m_{K}^{*} = \sqrt{m_{K}^{2} - \frac{\Sigma_{KN}}{f_{\pi}^{2}}\rho_{s} + V_{\mu}V^{\mu}}$$

$$V_{\mu} = \frac{3}{8f_{\pi}^{*2}j_{\mu}}$$

$$E^{*} = \sqrt{\mathbf{k}^{*2} + m_{K}^{*2}} + V_{0}$$

$$\mathbf{k}^{*} = \mathbf{k} - \mathbf{V}$$

$$\Sigma_{KN} = 250 - 450 \text{ MeV}$$

$$f_{\pi} = 93 \text{ MeV}, f_{\pi}^{*2} = 0.6f_{\pi}^{2}$$

$$U = E^* - E_{vac.} = E^* - \sqrt{p^2 + m^2}$$

For nuclear matter at rest $\langle V_{1,2,3} \rangle = 0 \Rightarrow k^* = k$

$$U = \sqrt{\mathbf{k}^{*2} + m^{*2}} + V_0 - \sqrt{\mathbf{k}^2 + m_{vac.}^2}, \ m^* < m$$





- Systematical modification of pt-spectra owe to the repulsive potential.
- Uncertainties in the model parameters (np cross sections, ...).











All the results



System (energy)	Experiment	Kaon potential at ρ p=0 [MeV]
π+A (1.02 GeV)	FOPI	20 ± 5
p+A (2.3 GeV)	ANKE	20 ± 5
Ar+KCI (1.76 GeV)	HADES 1)	39
p+Nb (3.5 GeV)	HADES 2)	40 ± 5

FOPI: M. Benabderrahmane et al., Phys. Rev. Lett. 102 (2009) 182501.

HADES: 1) G. Agakishiev et al., Phys. Rev. C 82 (2010) 044907; 2) Phys. Rev. C90 (2014) 054906.





Hyp-N: Experimental Evidence I





Λ- or Σ - Hypernuclei

Λ -Nucleon Potential



U~ -30 MeV (attractive) from Hypernuclei No idea yet about the momentum and density dependence

Σ -Nucleon Potential



No Idea at all





Λ

Hyp-N: Experimental Evidence II



 Λ -p Σ -p scattering Λ and Σ beams from K-+p collisions "seen" by Bubble chambers





The Femtoscopy Method





Distinguishable and Undistinguishable pairs of particles emitted from a common source

Correlation function is a measure of the source size and also of the particle interaction

F. Wang, and S.Pratt, Phys. Rev. Lett. 83 (1999) 3138

$$C(\vec{p}_{a},\vec{p}_{b}) = \frac{\mathcal{P}(\vec{p}_{a},\vec{p}_{b})}{\mathcal{P}(\vec{p}_{a})\mathcal{P}(\vec{p}_{b})} \approx \frac{\int d^{4}x_{a} d^{4}x_{b}S(p_{a},x_{a})S(p_{b},x_{b})|\phi_{rel}(\vec{p}_{b}-\vec{p}_{a})|^{2}}{\int d^{4}x_{a} d^{4}x_{b}S_{a}(\vec{p}_{a},x_{a})S_{b}(\vec{p}_{b},x_{b})}$$

Theoretical Function

 $C(k) = \mathcal{N} \frac{N(\mathbf{p}_1, \mathbf{p}_2)_{\text{same}}}{N(\mathbf{p}_1, \mathbf{p}_2)_{\text{mixed}}} \quad k = \frac{1}{2} |\mathbf{p}_1 - \mathbf{p}_2|$

Experimental Observable

p-p-, п-п Correlations in Ar+KCl at 1.75 AGeV Hades, EPJA 47, 63 (2011) A-p Correlations in Ar+KCl at 1.75 AGeV Hades, [PRC 82, 021901 (2010)].





Femtoscopy in p+A reactions (GeV)

p+Nb, 3.5 GeV

Interaction

kinematic freeze-our surface

Can be determined for p-p and Ap pairs via Transport Calculation (UrQMD) -> The Source is hence known and the measured correlation provides the interaction strength.









Ap Scattering Length



Experimental DistributionFit Function $C(k) = \mathcal{N} \frac{N(\mathbf{p}_1, \mathbf{p}_2)_{same}}{N(\mathbf{p}_1, \mathbf{p}_2)_{mixed}}$ $k = \frac{1}{2} |\mathbf{p}_1 - \mathbf{p}_2|$ $C(k) = 1 + \sum_S \rho_S \left[\frac{1}{2} \left| \frac{f^S(k)}{r_0} \right|^2 \frac{2\mathcal{R} f^S(k)}{\sqrt{\pi r_0}} F_1(Qr_0) - \frac{\mathcal{I} f^S(k)}{r_0} F_2(Qr_0) \right]$ Preliminary**PreliminaryPreliminaryPreliminaryPreliminary**

1.65

1.65

3.5

2.5

d0_{1S0} [fm]

1.7

R_G [fm]

1.75

1.75

R_G [fm]

1.8

1.8

k [MeV/c] First Fit of the correlation delivers parameters in perfect agreement with scattering data !!!! Further Improvements:

200

100

- Source-Radius Determination,
- Improved S/B Ratio,
- Fit for S=0 and S=1 components.



1.85





Example for p-p correlations



After-Burner which includes the relevant Interactions



p+Nb reaction simulated in UrQMD + CRAB afterburner

Coulomb +











Summary and Outlook



K⁰**s**-nucleons attractive potential verified with p+Nb cactions at 3.5 GeV $\int_{0}^{10} \int_{0}^{10} \int_{0}^{$

Test the Hyperon-Nucleon Interaction with <u>Femtoscopy</u> at intermediate and high energies (compare to Lattice potential or other calculations)











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The people in Munich





Thanks



Experiments with π Beams





- π-absorption mostly on the nucleus surface
- less model dependent

Study of Hadron-nucleon interaction

Not so easy to measure, since π -beams are secondary beams with large emittance



CERBEROS: 3-heads dog at the HADES entrance



Experiments with π Beams





- π-absorption mostly on the nucleus surface
- less model dependent

Study of Hadron-nucleon interaction





First Measurement of K absorption in normal nuclear matter



Hyperon Star





Strangeness violation possible due to large time scale of NS Appearance of Hyperon already starting at $2\varrho_0$ This scenario might also be problematic for large masses (~ $2M_{\odot}$) since the hyperon appearance implies new degree of freedom and hence a softening of the EOS

