Advances in holmium-based neutrino mass experiments

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International School of Nuclear Physics - 43rd Course
Neutrinos in Cosmology, in Astro-, Particle- and Nuclear Physics
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Outline

- direct neutrino mass measurements
- calorimetric measurements with low temperature detectors
- low Q beta decay experiments
- ^{163}Ho EC decay calorimetric experiments
  - decay spectrum
  - statistical sensitivity
- HOLMES and ECHo experiments
  - detectors
  - isotope production and embedding
  - arrays readout, data acquisition and processing
  - present results
- future of holmium based experiments

Neutrino properties

- neutrinos are massive fermions
- there are 3 active neutrino **flavors: e, μ, τ**
- neutrino flavor states are mixtures of 3 **mass states**

neutrino mixing matrix

neutrino **mass** eigenstate

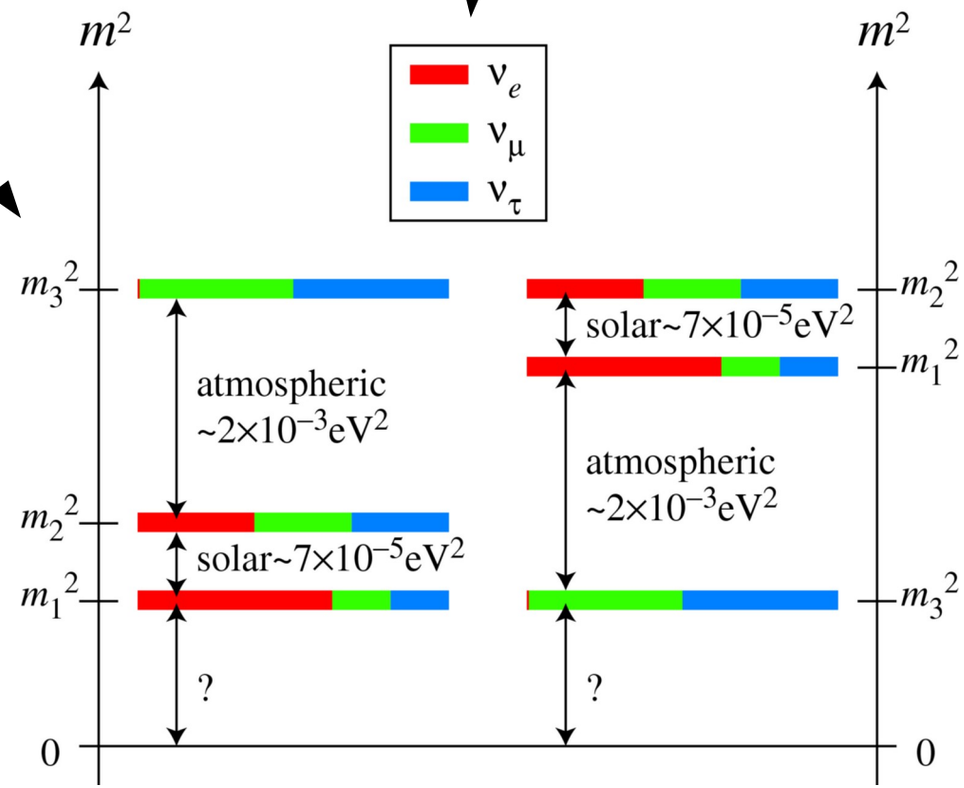
$$|\nu_l\rangle = \sum_k U_{lk} |\nu_k\rangle$$

neutrino **flavor** weak eigenstate

from neutrino oscillation experiments

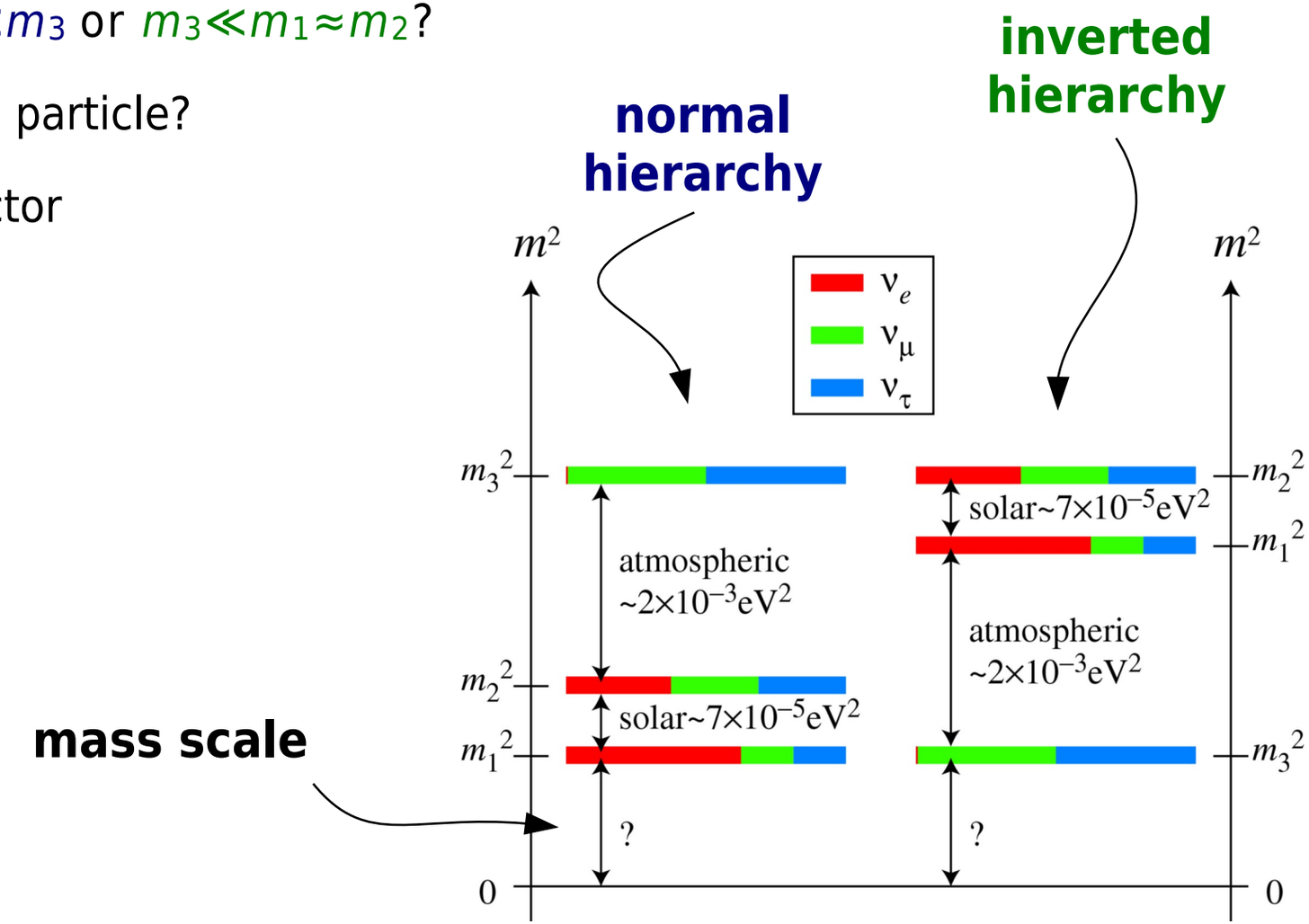
$$\Delta m_{ik}^2 = |m_i^2 - m_k^2|$$

$$\sin^2 2\vartheta_{lk} = f(|U_{lk}|^2)$$



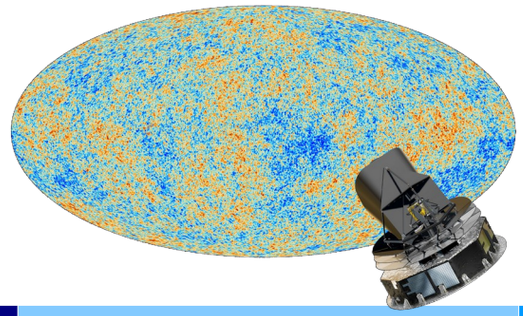
Neutrino open questions

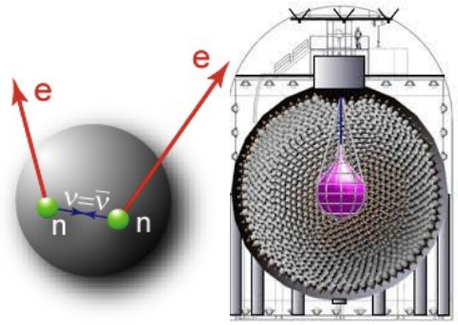
- mass scale: i.e. mass of the lightest ν
- degenerate ($m_1 \approx m_2 \approx m_3$) or hierarchical masses
  - mass hierarchy: $m_1 < m_2 \ll m_3$ or $m_3 \ll m_1 \approx m_2$?
- $\nu = \bar{\nu}$? i.e. Dirac or Majorana particle?
- CP violation in the lepton sector



Direct ν mass measurements: the status

three complementary tools available



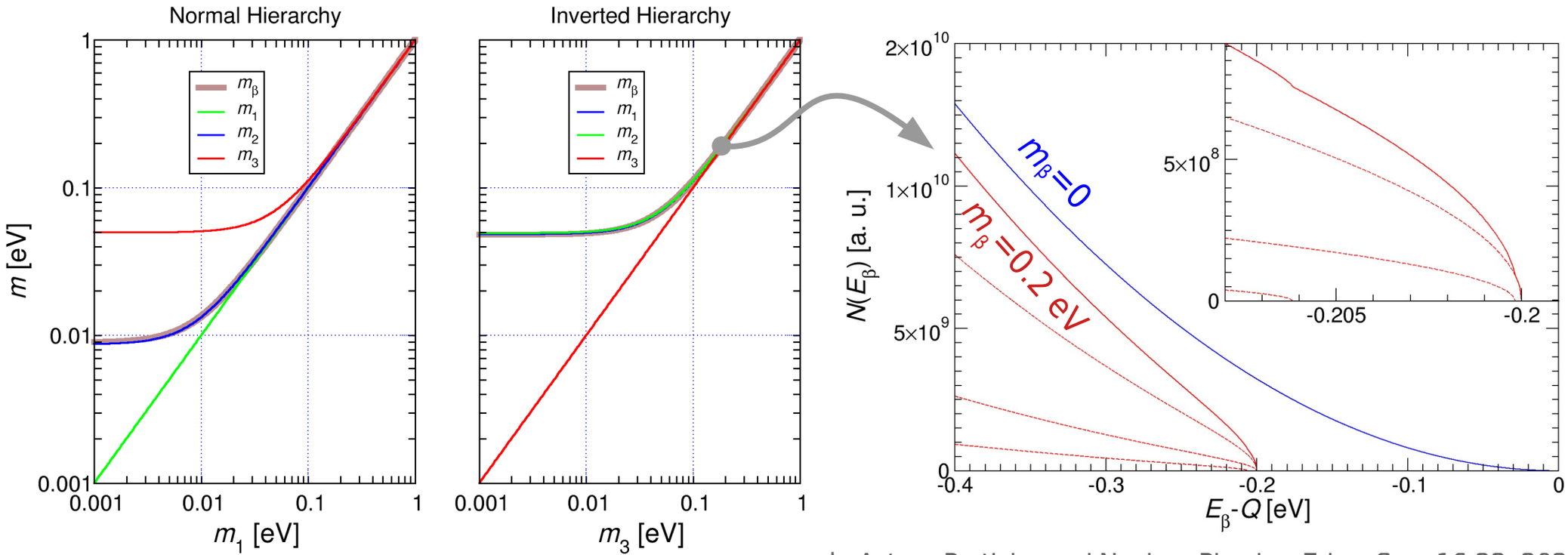


tool	Cosmology CMB+LSS+...	Neutrinoless Double Beta decay	Beta decay end-point
observable	$m_\Sigma=\sum_k m_{\nu_k}$	$m_{\beta\beta}=\lvert\sum_k m_{\nu_k} U_{ek}^2\rvert$	$m_\beta=(\sum_k m_{\nu_k}^2 \lvert U_{ek}\rvert^2)^{1/2}$
present sensitivity	≈0.1 eV	≈0.1 eV	≈1 eV
≈10y future sensitivity	**≈0.05 eV**	**≈0.01 eV**	**≈0.1 eV**
model dependency	yes ☹	yes ☹	no ☺
systematics	large ☹	some 😐	large ☹

Direct neutrino mass measurements

- **kinematics of weak decays**
  - in beta and electron capture decays where $\bar{\nu}_e$ or ν_e are emitted $|\nu_e\rangle = \sum_k U_{ek}|\nu_k\rangle$
  - non zero neutrino masses $m_{\nu k}$ modify the phase space
  - for nuclear β decay $N(E) \propto p_\beta E_\beta (Q - E_\beta) \sum_k |U_{ek}|^2 \sqrt{(Q-E_\beta)^2 - m_{\nu_k}^2}\, F(Z, E_\beta) S(E_\beta)$
- for degenerate masses ($m_{\nu 1} \approx m_{\nu 2} \approx m_{\nu 3}$)

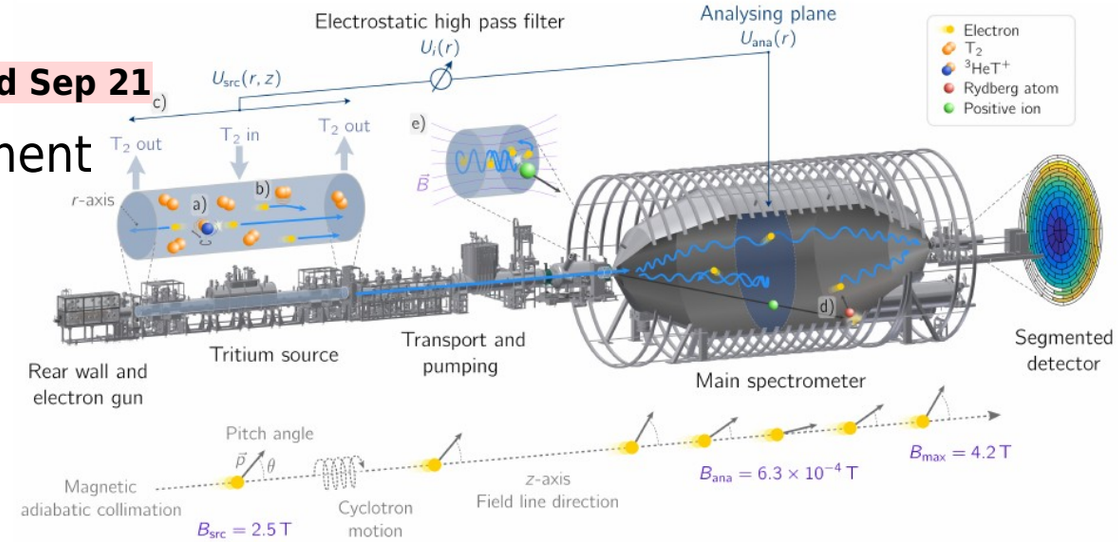
$$N(E) \approx p_\beta E_\beta (Q - E_\beta) \sqrt{(Q-E_\beta)^2 - m_\beta^2}\, F(Z, E_\beta) S(E_\beta) \quad \text{with} \quad m_\beta = \sqrt{\sum_k m_{\nu_k}^2 |U_{ek}|^2}$$



Tritium experiments

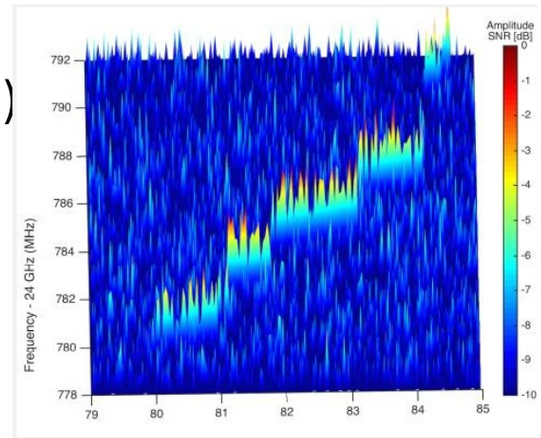
KATRIN *Nat. Phys. 18, 160–166 (2022)* **Lokhov + Schwemmer on Wed Sep 21**

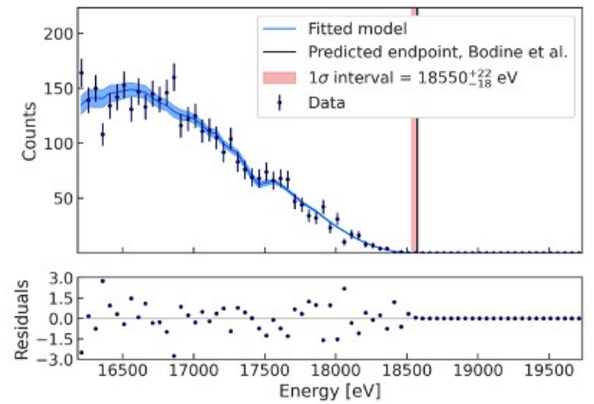
- MAC-E filter: ultimate integral spectrometer experiment
- sensitivity goal: 0.2 eV 90% CL
- energy resolution <3 eV @18 keV.
- first high-purity tritium campaign in 2019
  → m_ν **< 0.8 eV** 90% CL



Project8 *arXiv:2203.07349*

- Cyclotron Resonance Electron Spectrometry (CRES)
- sensitivity goal: 40 meV 90% CL
- 4 different experimental phases: phase III ongoing
- energy resolution ≈2 eV @18 keV

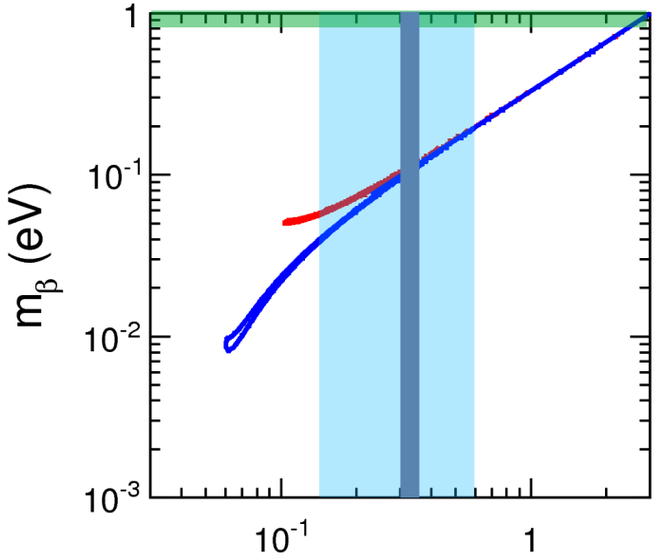




PTOLEMY *M.G. Betti et al., Prog. Part. Nucl. Phys, 106 (2019)*

- project to measure the Cosmic Neutrino Background via neutrino capture on tritium
- differential spectrometer combining CRES with an EM dynamic filter and hi-res microcalorimeters
- sensitivity potential: *O*(10) meV
- presently: small prototype R&D

Direct ν mass measurements: 2022 status



NH and **IH** 2σ bands from oscillation parameters

G.Fogli et al., Phys. Rev. D 86 (2012) 013012 (only minor updates in 2022)

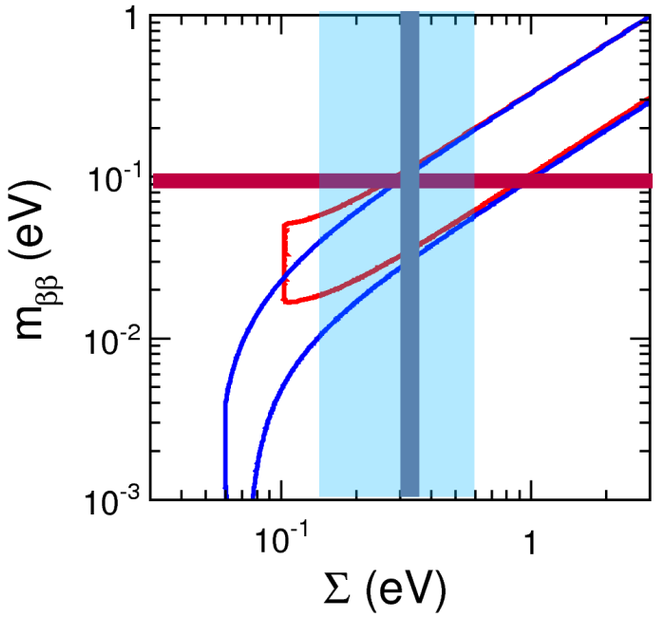
$m_{\beta\beta}$ from 0ν double beta decay experiments

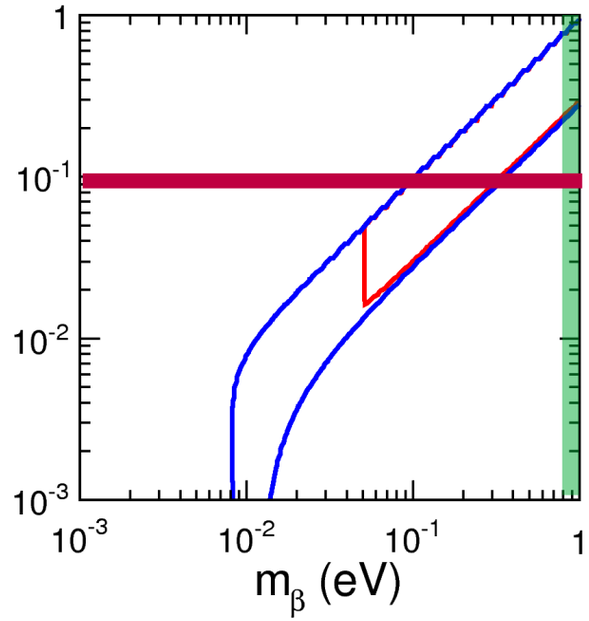
E. Lisi, A. Marrone, Phys.Rev.D 106 (2022) 1, 013009. arXiv:2204.09569

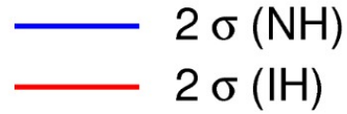
Σ from cosmic microwave background measurements

F. Capozzi, et al. Phys. Rev. D 104, 083031. arXiv:2107.00532

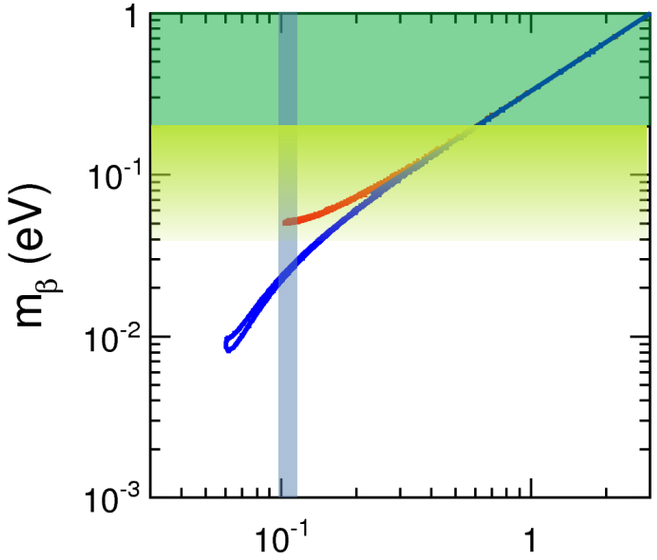
m_β from KATRIN *KATRIN Coll., Nat. Phys. 18, 160–166 (2022)*







Direct ν mass measurements: role of kinematic exp.

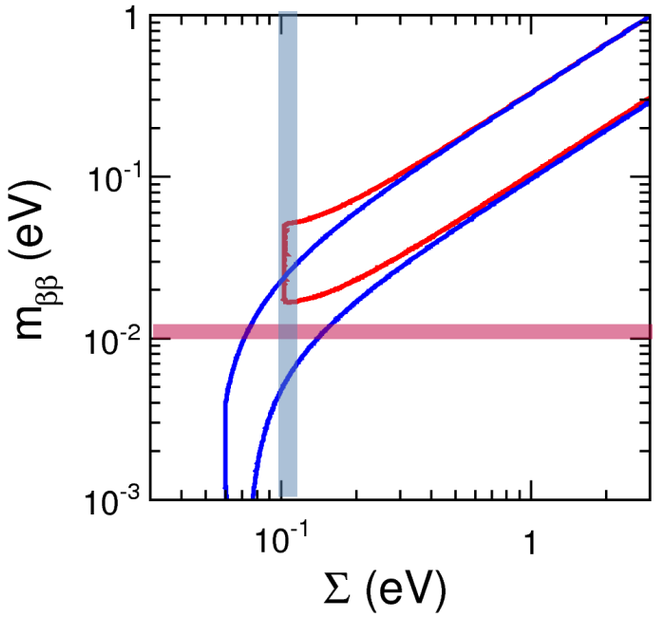


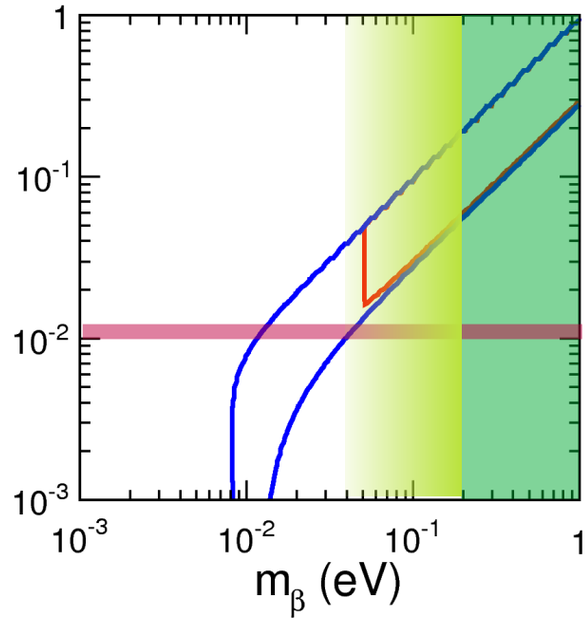
NH and **IH** 2σ bands from oscillation parameters

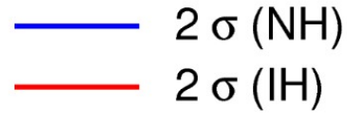
G.Fogli et al., Phys. Rev. D 86 (2012) 013012 (only minor updates in 2022)

m_β KATRIN goal

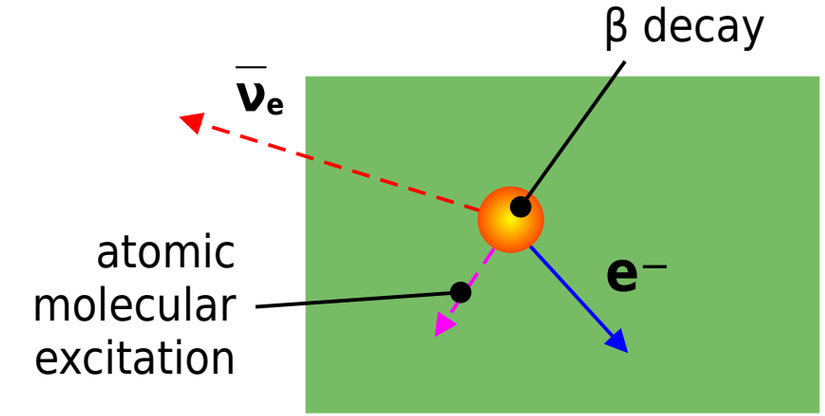
m_β $m_{\beta\beta}$ Σ next generation experiments







Calorimetric experiments



ideal calorimetric experiment

- radioactive source embedded in the detector(s)
- only the neutrino energy escapes detection

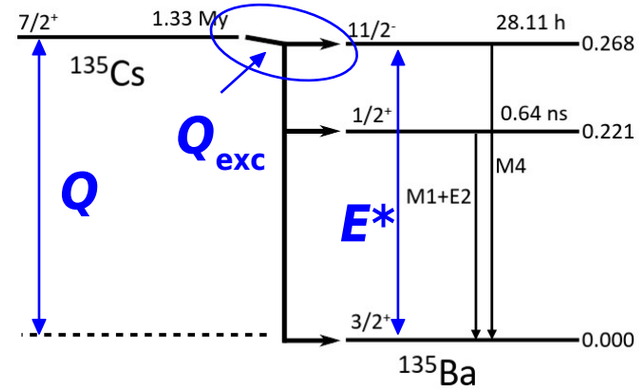
→ $E_c = Q - E_\nu$

- ▲ no backscattering
- ▲ no energy losses in source
- ▲ no decay final state effects
- ▲ no solid state excitation
- ▼ low activity → limited statistics
- ▼ pile-up background

ideal isotope has

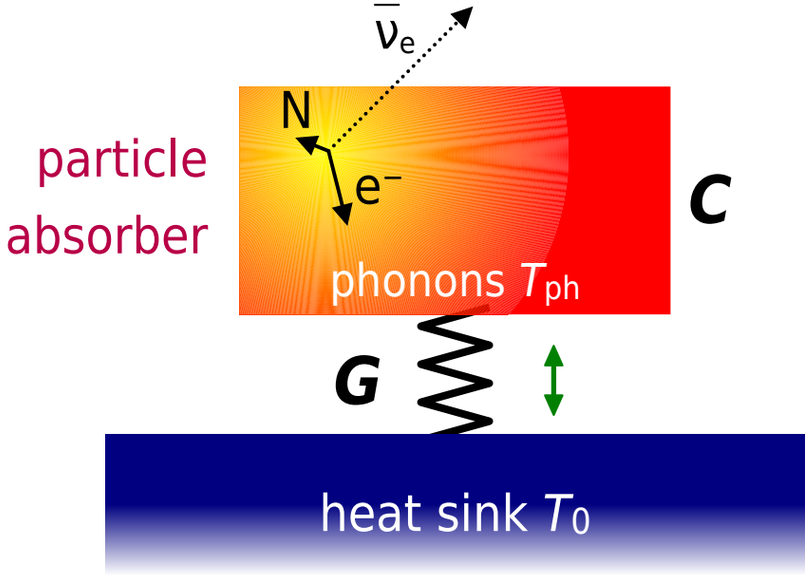
- low Q
  → larger fraction of decays in ROI
  → easier calorimetry
- for EC: capture peak close to end-point
- short decay time

isotope	Q [eV]	$\tau_{1/2}$ [y]	decay	B.R.	experiments
^{3}H	18592.01(7)	12	β^-	1	Simpsons's
^{187}Re	2470.9(13)	4.3×10^{10}	β^-	1	MANU, MIBETA
^{163}Ho	2833(30)	4570	EC	1	Holmes, ECHo
^{135}Cs	440	8.0×10^{11}	β^-	1.6×10^{-6}	-
^{115}In	155	4.3×10^{20}	β^-	1.1×10^{-6}	-



A. de Roubin et al. PRL. 124, 222503 (2020)

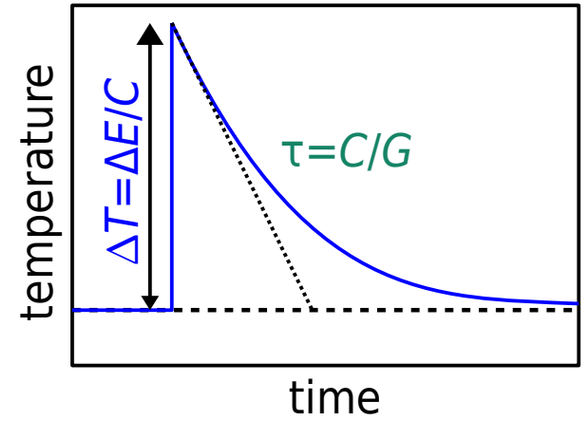
Low temperature detector principles



$$C(T_{ph})\frac{dT_{ph}}{dt}+G(T_{ph},T_0)=P(t)$$

$$P(t)=\Delta E\,\delta(t) \rightarrow T_{ph}(t)=T_0+\frac{\Delta E}{C}e^{-t/\tau}$$

for $t>0$ and with $\tau=C/G$



energy resolution limited by thermodynamic fluctuation noise TFN

$$N_{ph}=\frac{\langle U\rangle}{\langle E_{ph}\rangle}=\frac{CT}{k_B T}$$

$$\sigma_E=\Delta U_{rms}=\sqrt{N_{ph}}\langle E_{ph}\rangle=\sqrt{k_B T^2 C}$$

***E* → Δ*T*≈Δ*E*/*C* → Δ*X*(*T*)**

e.g: $R=R(T)$, $M=M(T)$

- detectors used for calorimetric neutrino mass experiments are more complex
- in metallic calorimeters energy is transferred to electronic system with T_e
- thermodinamycs and statistical mechanics still provide for TFN $\sigma_E=\sqrt{k_B T^2 C}$

200×200×2 μm³ (1.5 μg)

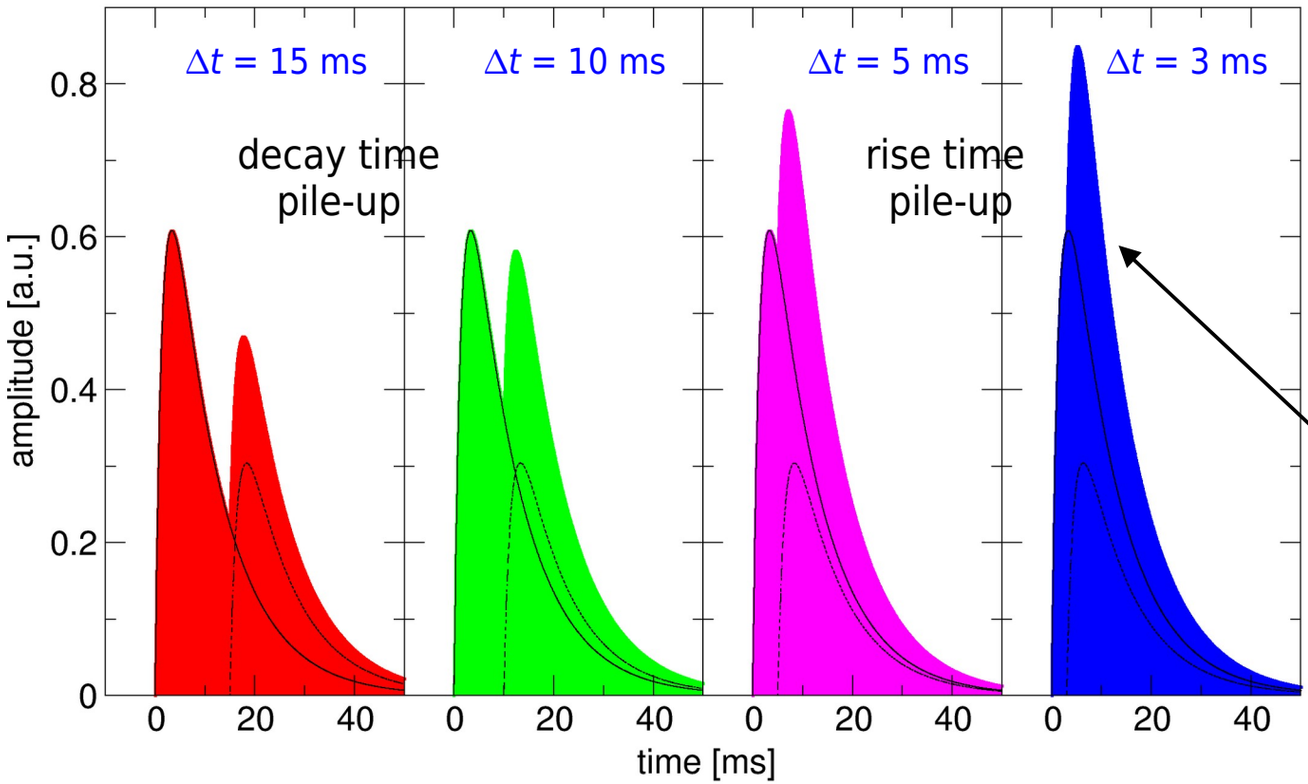
Au absorber @ 100 mK

$C\approx C_e\propto T_e \rightarrow C\approx 5\times10^{-13}$ J/K

σ_E≈3.4 eV (better estimate for TES detectors gives **σ_E≈0.4 eV**)

Pile-up in low temperature detectors

- calorimeters detect all β source decays
- low temperature detectors are *slow* devices
- simple pulse model $A(t)=A\left(e^{-t/\tau_{decay}}-e^{-t/\tau_{rise}}\right)$
  - for microcalorimeters: $\tau_{rise} \approx$ 0.1-10 μs and $\tau_{decay} \approx$ 0.1-10 ms



2 pulses with:

- τ_{rise}= 1.5 ms
- τ_{decay}= 10 ms
- A_2/A_1 = 0.5
- time separation Δt

first approximation for rise time p-up
resolving time $\boldsymbol{\tau_R} \approx$ rise time $\boldsymbol{\tau_{rise}}$

for $\Delta t < \tau_R$

→ accidental coincidence

→ $E_{meas} = E_1 + E_2$

- $\Delta t \gg \tau_{rise}$ → pile-up on the decay time → dead time
- $\Delta t \lesssim \tau_{rise}$ → pile-up on the rise time → spectral distortions and background

Calorimetry of beta decays: statistical sensitivity

rise time pile-up resolving time τ_R

ROI ΔE

source activity A_β

pile-up fraction $f_{\text{pile-up}} = \tau_R A_\beta$

measuring time T_M

number of detectors N_{det}

exposure $t_M = T_M \times N_{det}$

^{187}Re example:

$$N_\beta(E,m_\nu) \approx \frac{3}{Q^3}(Q-E)^2\sqrt{1-\frac{m_\nu^2}{(Q-E)^2}}$$

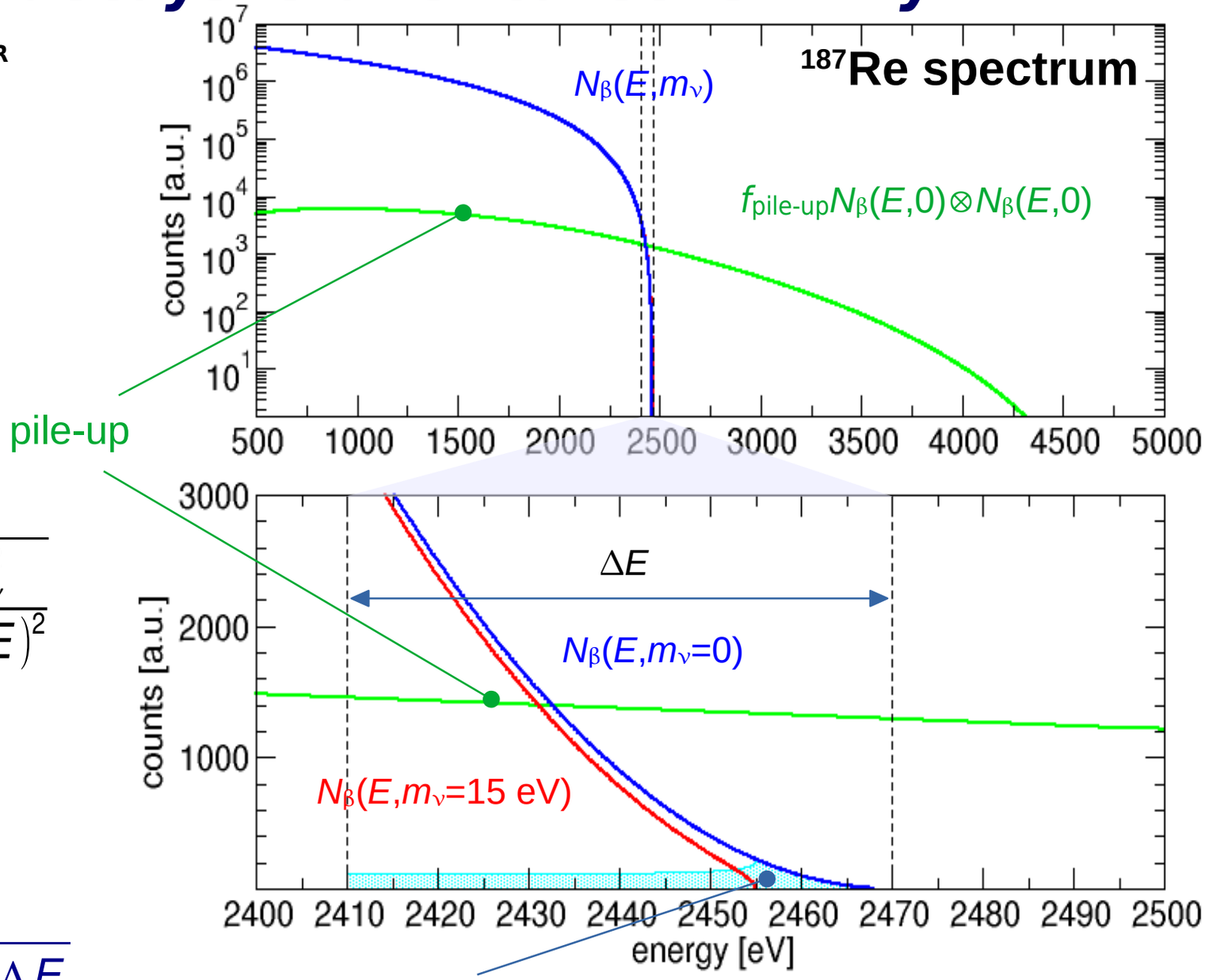
$\Delta E \approx \Delta E_{FWHM}$

pile-up is negligible when

$$f_{pile-up} = \tau_R A_\beta \ll \frac{\Delta E^2}{Q^2}$$

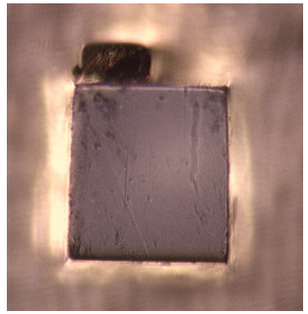
then

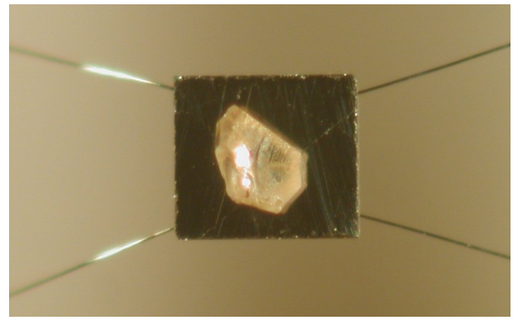
$$\sum_{90}(m_\nu) \approx \sqrt[4]{\frac{Q^3 \Delta E}{A_\beta t_M}} = \sqrt[4]{\frac{Q^3 \Delta E}{N_{decays}}}$$



A. Nucciotti et al., Astropart. Phys. 34, 80 (2010).

Rhenium-187 experiments





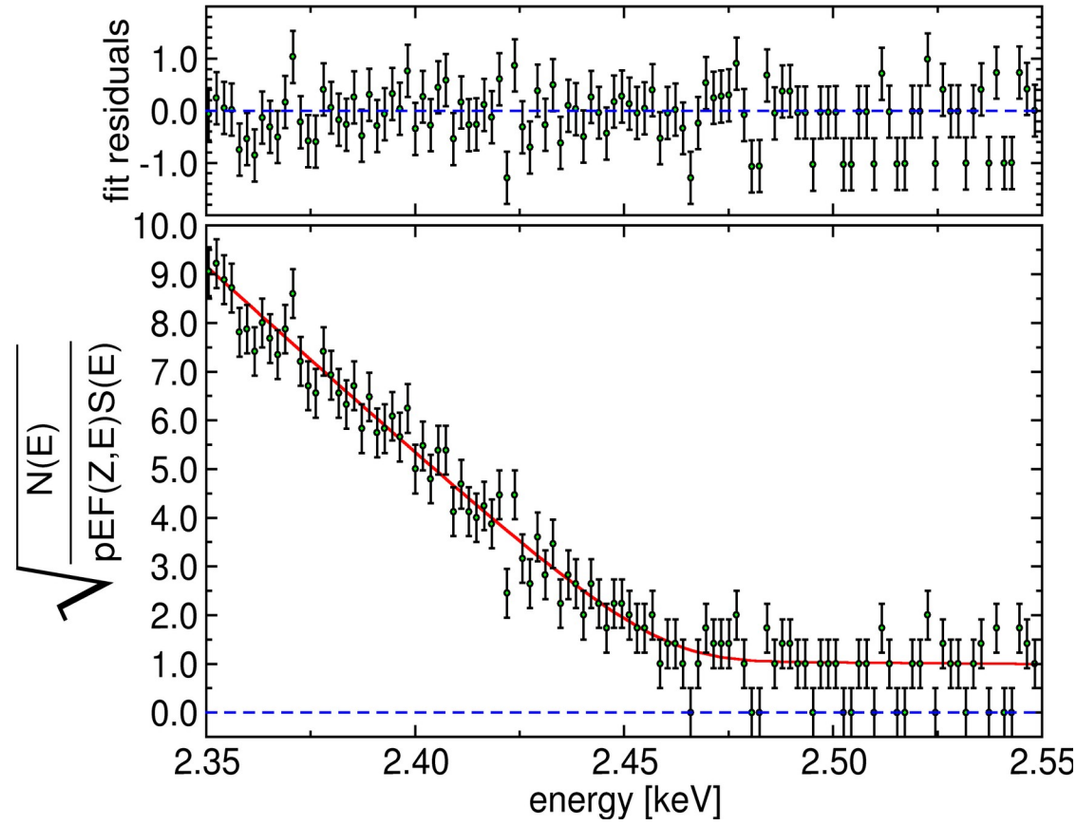
first ^{187}Re experiments: $N_{ev} \approx 10^7$ events

MIBETA @ Milano with $AgReO_4$

→ $m_\nu < 15$ eV 90% C.L. M.Sisti et al., NIM A 520 (2004) 125

- MANU @ Genova with metallic Re

→ $m_\nu < 26$ eV 95% C.L. F.Gatti et al., Nucl. Phys. B91 (2001) 293



1990 → 2006 MIBETA (Milano/MilanoBicocca) + MANU (Genova)

^{187}Re → m_ν<15 eV (+ BEFS...)

2006 MARE (Microcalorimeter Array for a Rhenium Experiment) int'l project for **m_ν<0.1 eV**

2007 → 2013 MARE R&D for phase 1→ Re+TES/MMC / $AgReO_4$+Si-Impl

2013 MARE project with ^{187}Re abandoned due to insurmountable technical obstacles

A. Nucciotti, Adv. High Energy Phys. 2016, 9153024 (2016)

Electron capture calorimetric experiments

^{163}Ho → ^{163}Dy[H] + ν_e

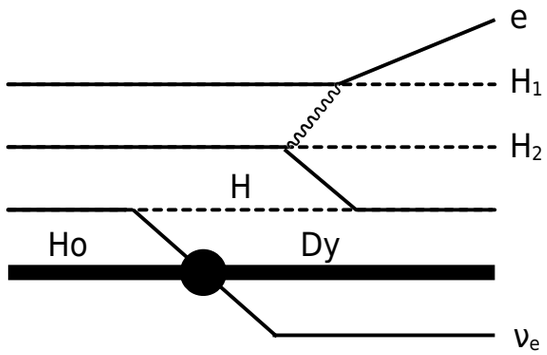
^{163}Dy[H] → ^{163}Dy + E_c

shell binding energy: E_b(M1)=2.05 keV

→ electron capture from shell ⩾M1

→ H=M1, M2, N1, N2, O1, O2, P1

Γ_{M1}≈13 eV



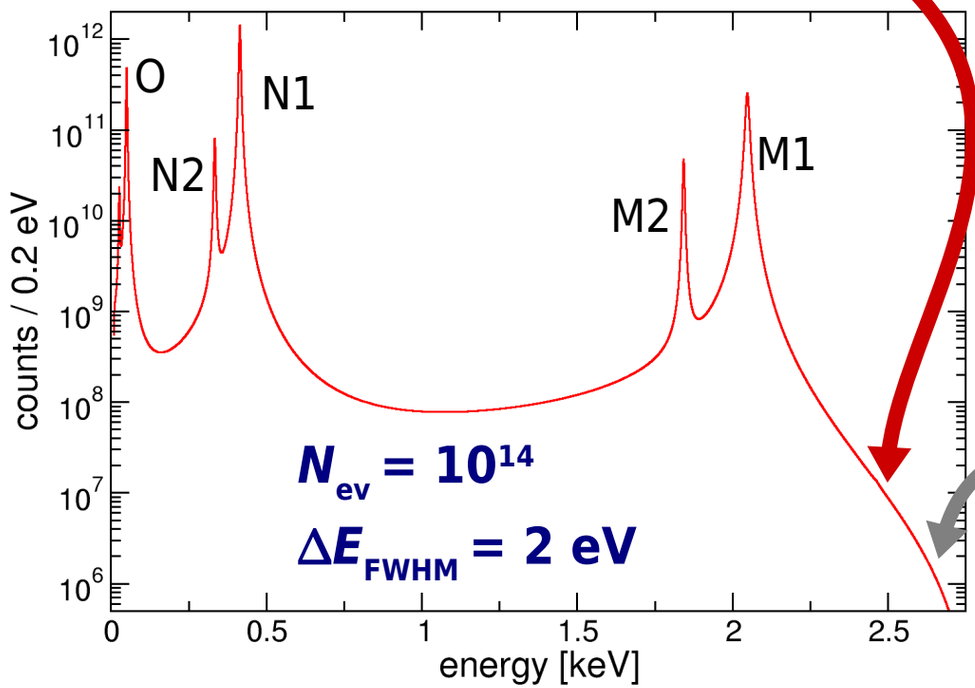
A. De Rújula and M. Lusignoli, Phys. Lett. B 118 (1982) 429

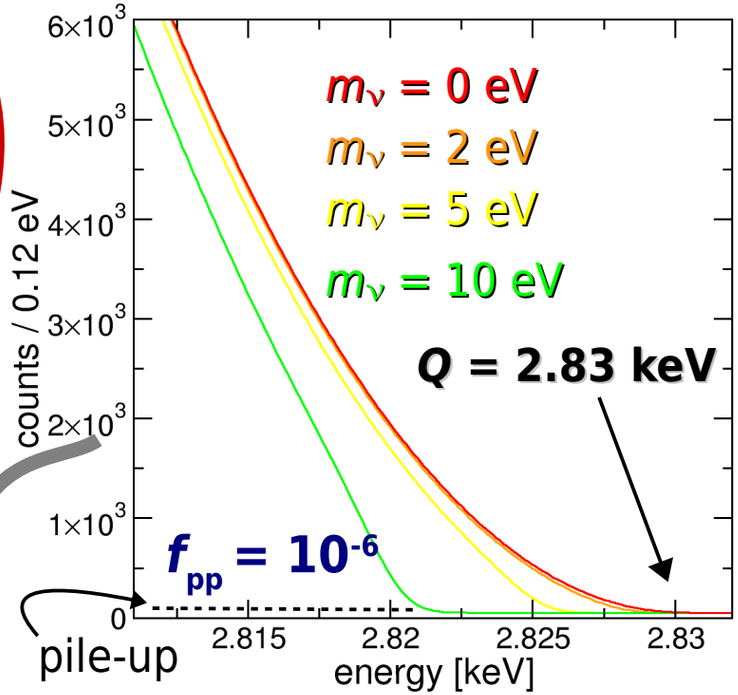
- calorimetric measurement of Dy atomic de-excitations (E_c)
  - ▷ mostly Auger and Coster-Kronig ($\omega_{M1,2}\approx10^{-3}$, $\omega_{N1,2}\approx10^{-5}$)
- Q=2.833±0.030^{stat}±0.015^{sys} keV

S. Eliseev et al., Phys. Rev. Lett. 115 (2015) 062501

  - ▸ end-point rate and ν mass sensitivity depend on $Q-E_{M1}$
- $\tau_{1/2}$≈4570 years → 2×10^{11} ^{163}Ho nuclei ↔ 1 Bq

$$N(E_c)=\frac{G_\beta^2}{4\pi^2}(Q-E_c)\sqrt{(Q-E_c)^2-m_\nu^2}\times\sum_i n_i C_i \beta_i^2 B_i \frac{\Gamma_i}{2\pi}\frac{1}{(E_c-E_i)^2+\Gamma_i^2/4}$$

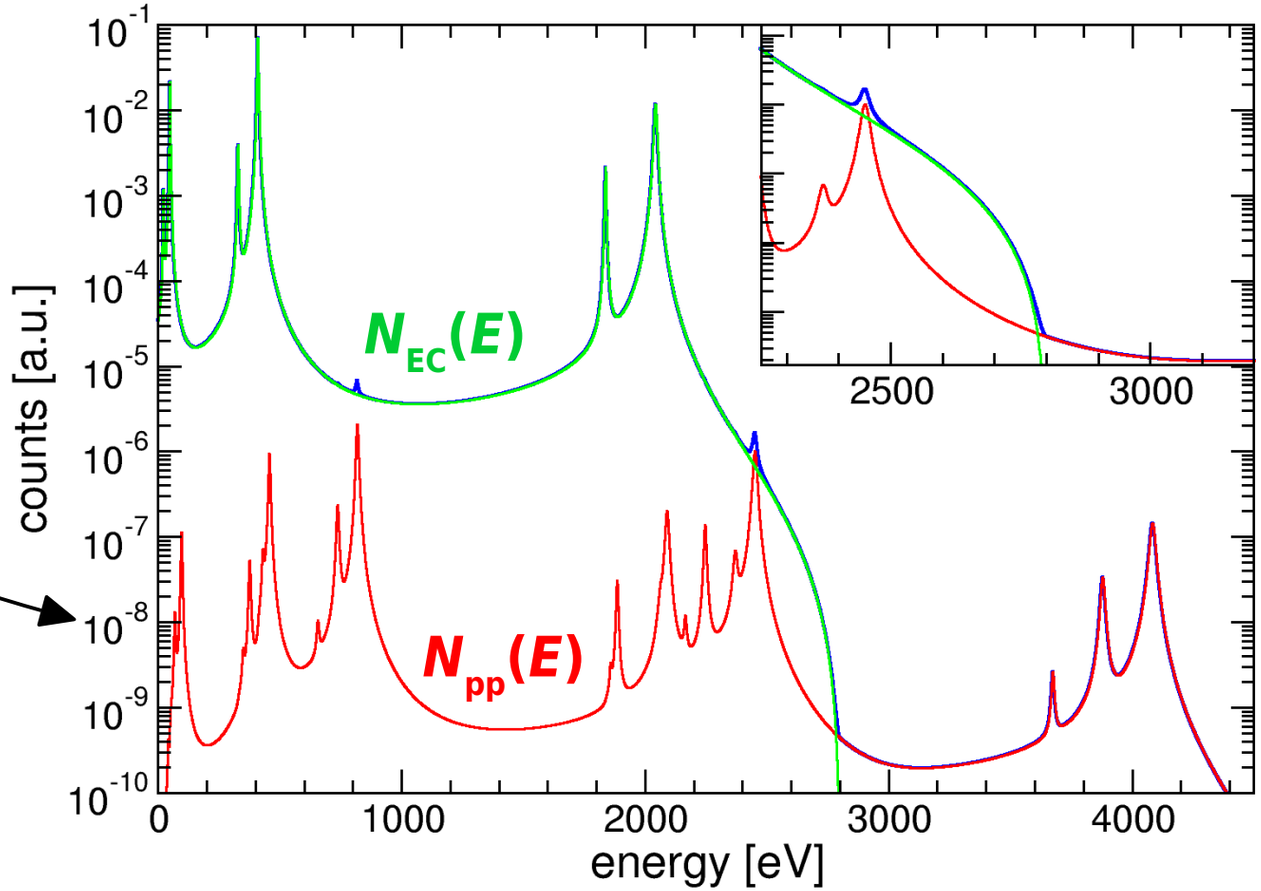




Pile-up in ^{163}Ho EC calorimetric experiments / 2

- accidental coincidences → complex pile-up spectrum
- calorimetric measurement → **detector speed is critical**
  - $N_{pp}(E) = f_{pp}\, N_{EC}(E) \otimes N_{EC}(E)$ with $f_{pp} \approx A_{EC}\, \tau_R$

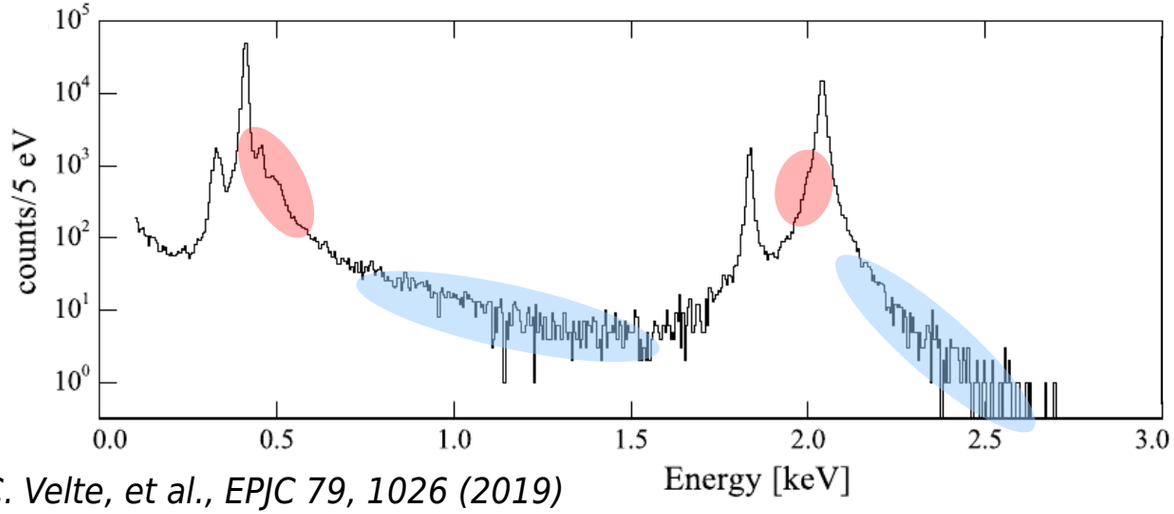
A_{EC} EC activity per detector
τ_R time resolution (≈rise time)



Q = 2800 eV

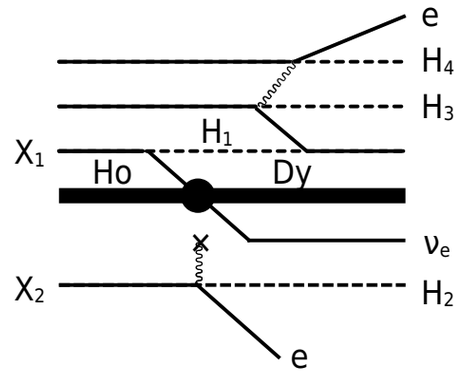
f_{pp} = 10^{-4}

Double hole processes / 1



C. Velte, et al., EPJC 79, 1026 (2019)

- high-statistic low background **ECHo** spectrum shows extra features
  - extra peaks close to N1 and M1
  - N/M peak asymmetries / high energy tails



Single hole: the Dy atom is left by EC with one hole in the shell (M, N, O…)

→ for H_1 in shell X_1 with binding energy $E_b(X_1)$ → resonance for $E_c=E_b(X_1)$

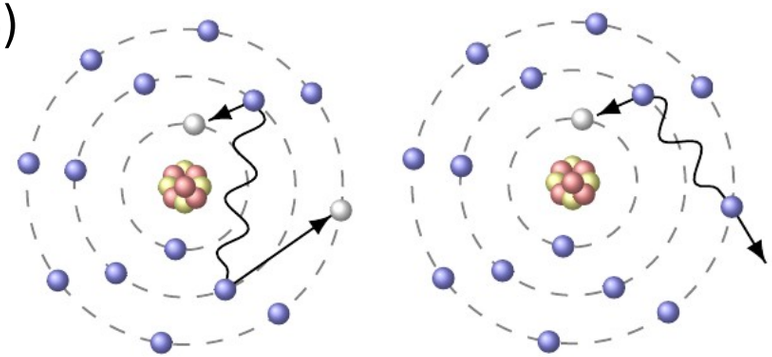
Double hole processes

the perturbation due to the nuclues charge change (Ho→Dy) “shakens” one or more additional atomic electron to an upper bound state (shake-up) or to the continuum (shake-off or Auger)

→ shake up: additional hole H_2 in X_2 → resonance for $E_c=E_b(X_1)+E_b(X_2)$

→ shake off: additional hole H_2 in X_2

→ tail to peaks from $E_c=E_b(X_1)+E_b(X_2)$ up to $E_c=Q$



Double hole processes / 2

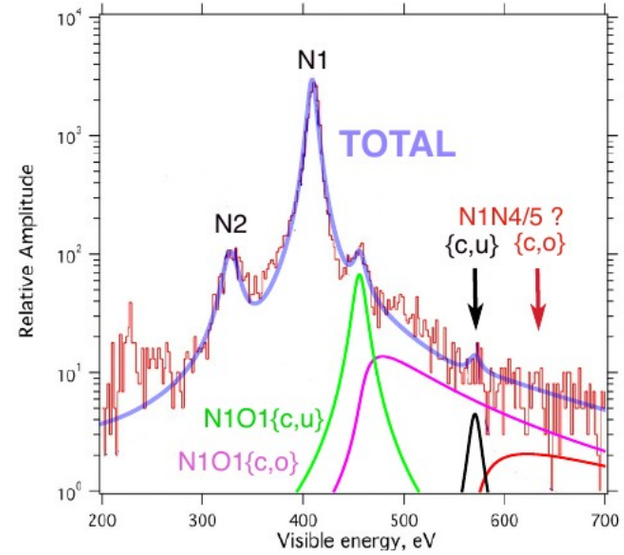
- several attempts to include double hole processes

  H. Robertson et al., A. Faessler et al., A. De Rújula and M. Lusignoli, ...

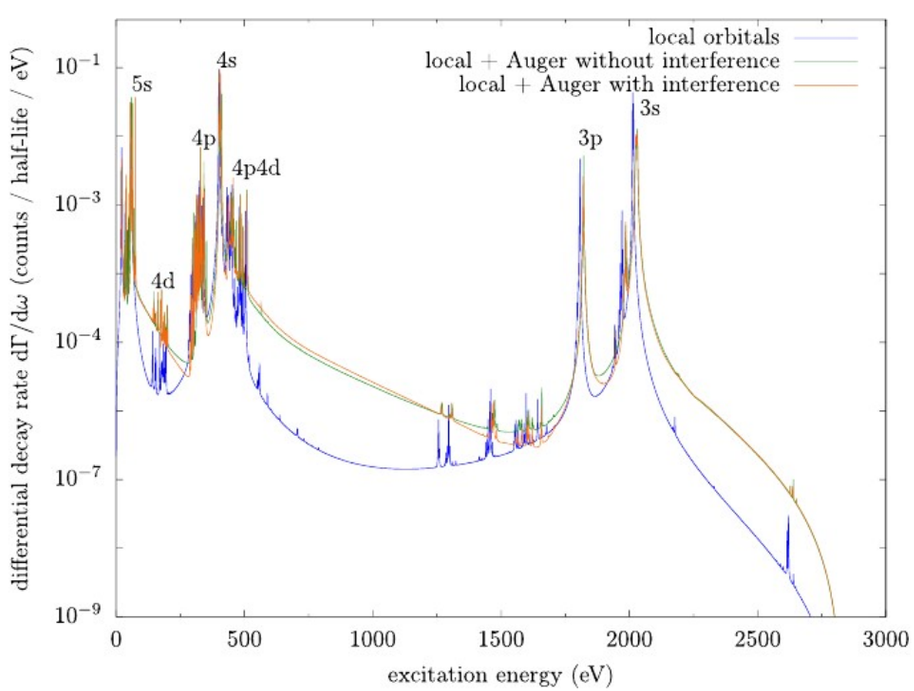
- recent work from M. Haverkort and collaborators:

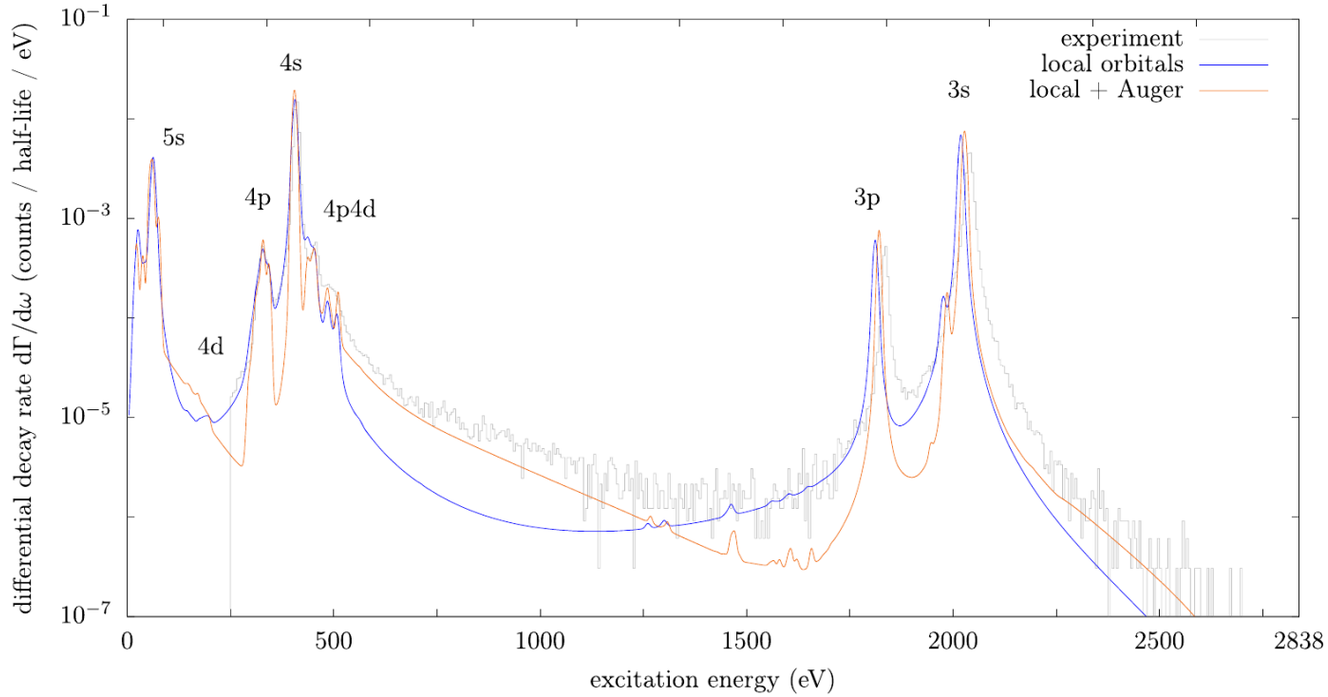
  ab-initio calculations including Coulomb interactions between multi core bound and unbound states (work in progress)

  - missing because of computational limits: intrinsic resonance linewidths, full shake-off contributions, radiative transitions, ...



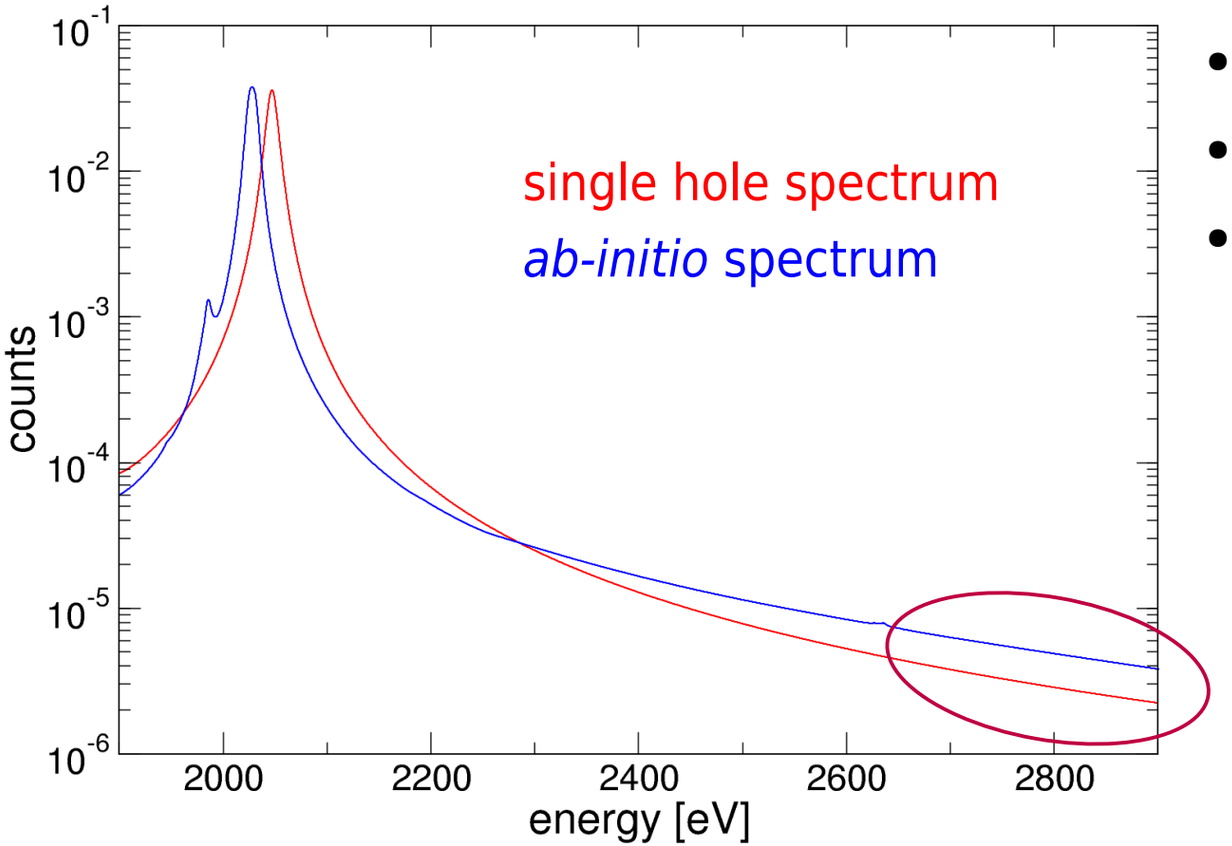
A.De Rújula & M. Lusignoli, J. High Energ. Phys. (2016) 2016: 15





M. Brass and M. W. Haverkort, New J. Phys. 22 (2020) 093018

End-point spectral shape



- "bare" spectra (without phase space)
- *ab-initio* with additional Lorentzian broadening
- spectra are normalized to unity
  - end-point region is smooth and featureless
  - phase space factor leaves unmistakable imprint

  → possibly small systematic uncertainties

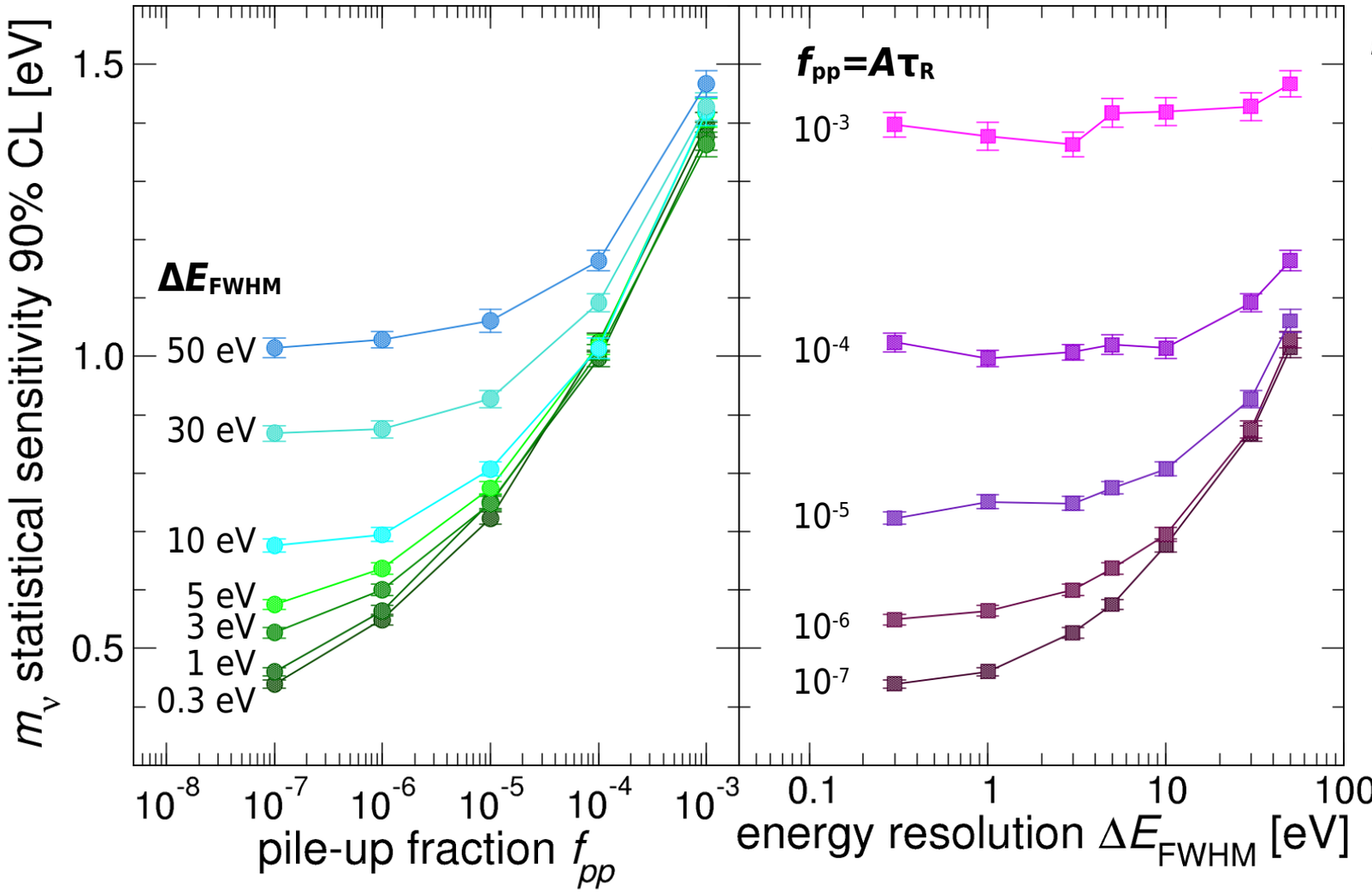
  → to be proved

- apparently *ab-initio* spectrum has higher rate at endpoint
- also pup spectrum is higher

  → small gain on statistical sensitivity

Statistical sensitivity: pile-up and energy resolution

- Montecarlo simulations for statistical sensitivity with **single-hole spectrum**
- simulations confirm that sensitivity Σ scales as $1/(N_{ev})^{0.25}$



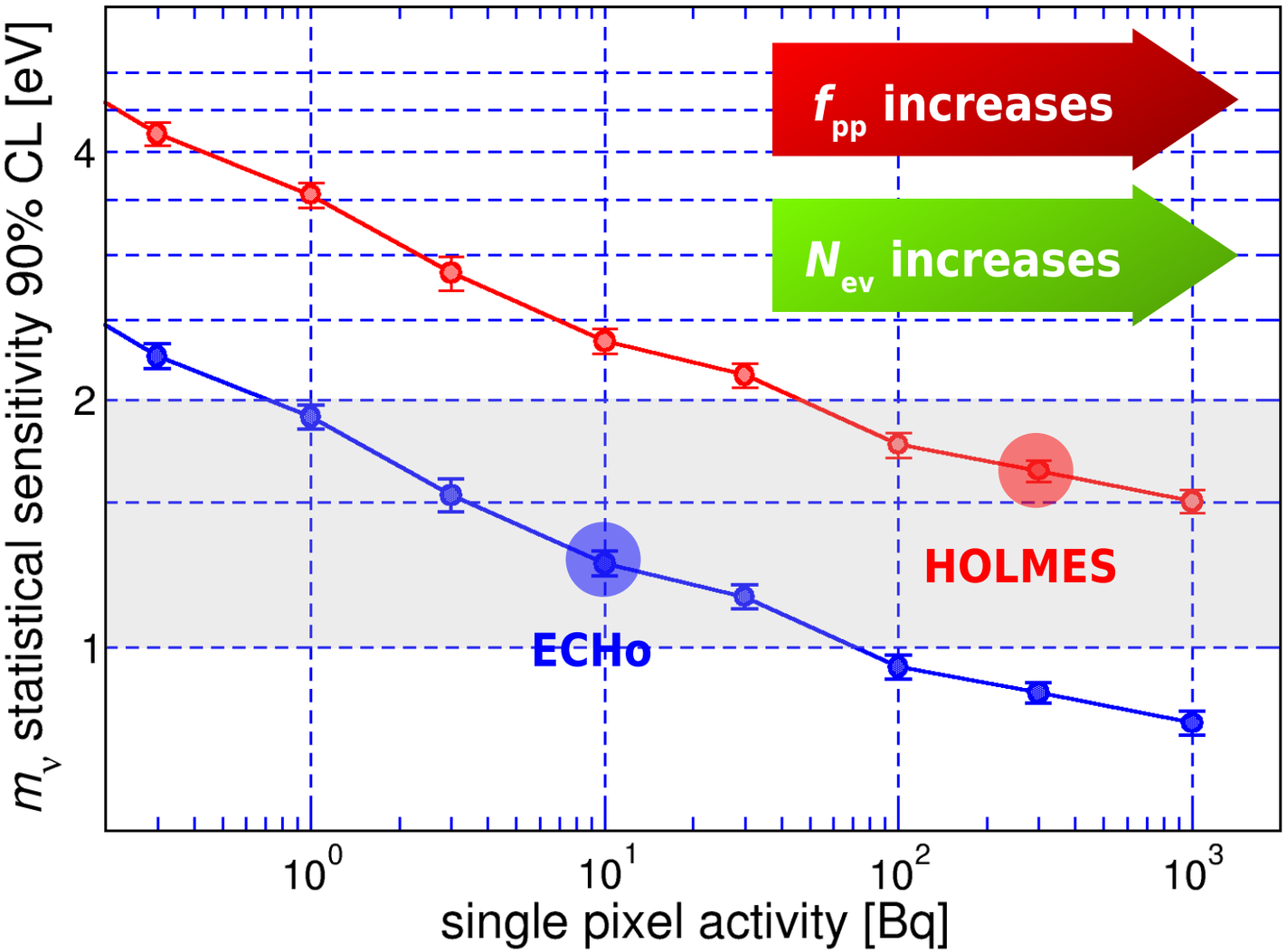
$N_{ev}=10^{14}$

$\Sigma(m_\nu) \propto \sqrt[4]{1/N_{ev}}$

$N_{ev} = A\ t_M\ N_{ch}$

A. Nucciotti, Eur. Phys. J. C 74.11 (2014)

Neutrino mass statistical sensitivity



MC simulations

A. Nucciotti, Eur. Phys. J. C 74.11 (2014)

single-hole spectrum

ΔE =1 eV

τ_R = 1 μs

$N_{det} t_M$ = 1000 det × 3 years

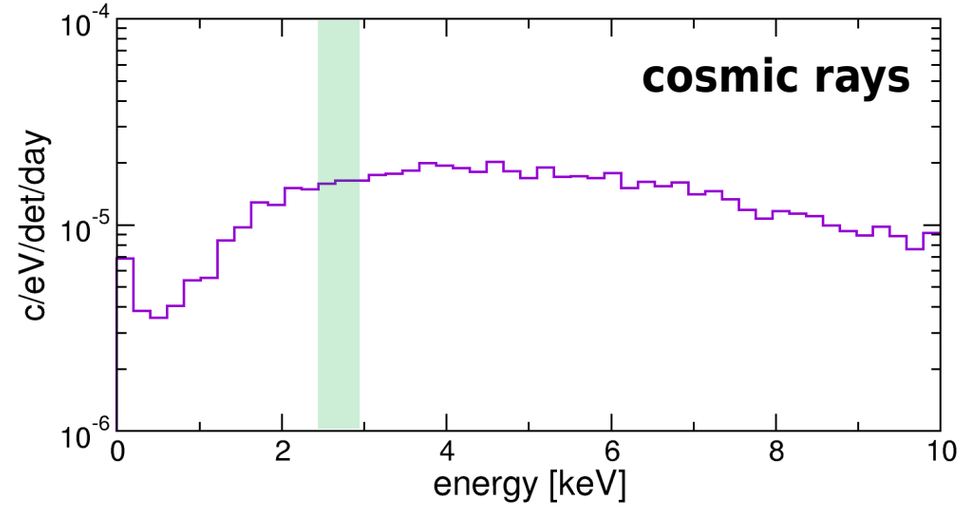
$N_{det} t_M$ = 12000 det × 3 years

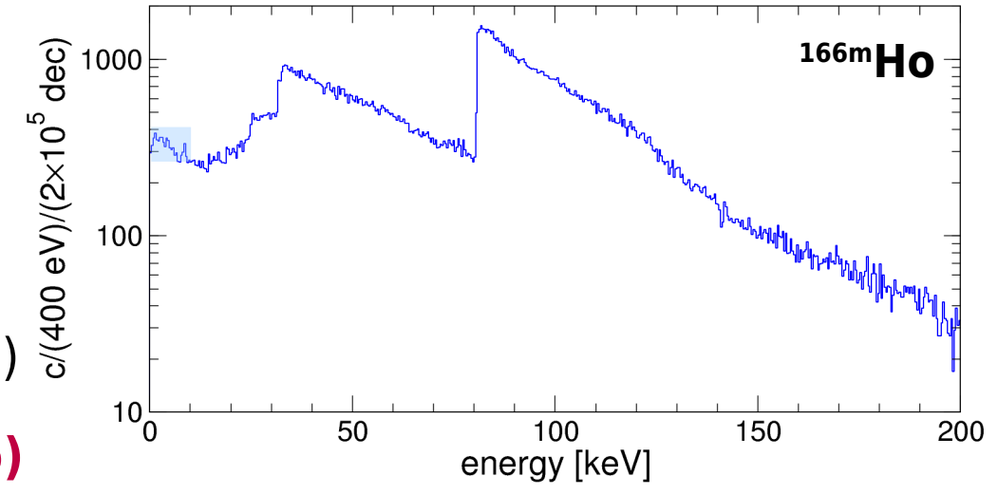
high activity → robustness against (flat) background

A_{EC}=300 Bq → b<≈0.1 counts/eV/day/det

Low energy background

- pile-up in ROI (single-hole) $b_{pp} \approx 0.35\, f_{pp}\, A$ **c/eV/day**
- environmental γ radiation
- γ, X and β from close surroundings
- cosmic rays
  - ▷ GEANT4 (HOLMES) → $b_{CR} \approx 10^{-5}$ **c/eV/day/det** (0 - 4 keV)
- internal radionuclides
  - ▷ **^{166m}Ho** (β^-, Q=1.8 MeV, $\tau_{1/2}$=1200 y)
  - ▷ co-produced with ^{163}Ho: $A(^{163}\text{Ho})/A(^{166m}\text{Ho})>500$ (HOLMES)
  - ▷ GEANT4 (HOLMES) → $b_{166m} \approx 0.3$ **c/eV/day/det/Bq(^{166m}Ho)**

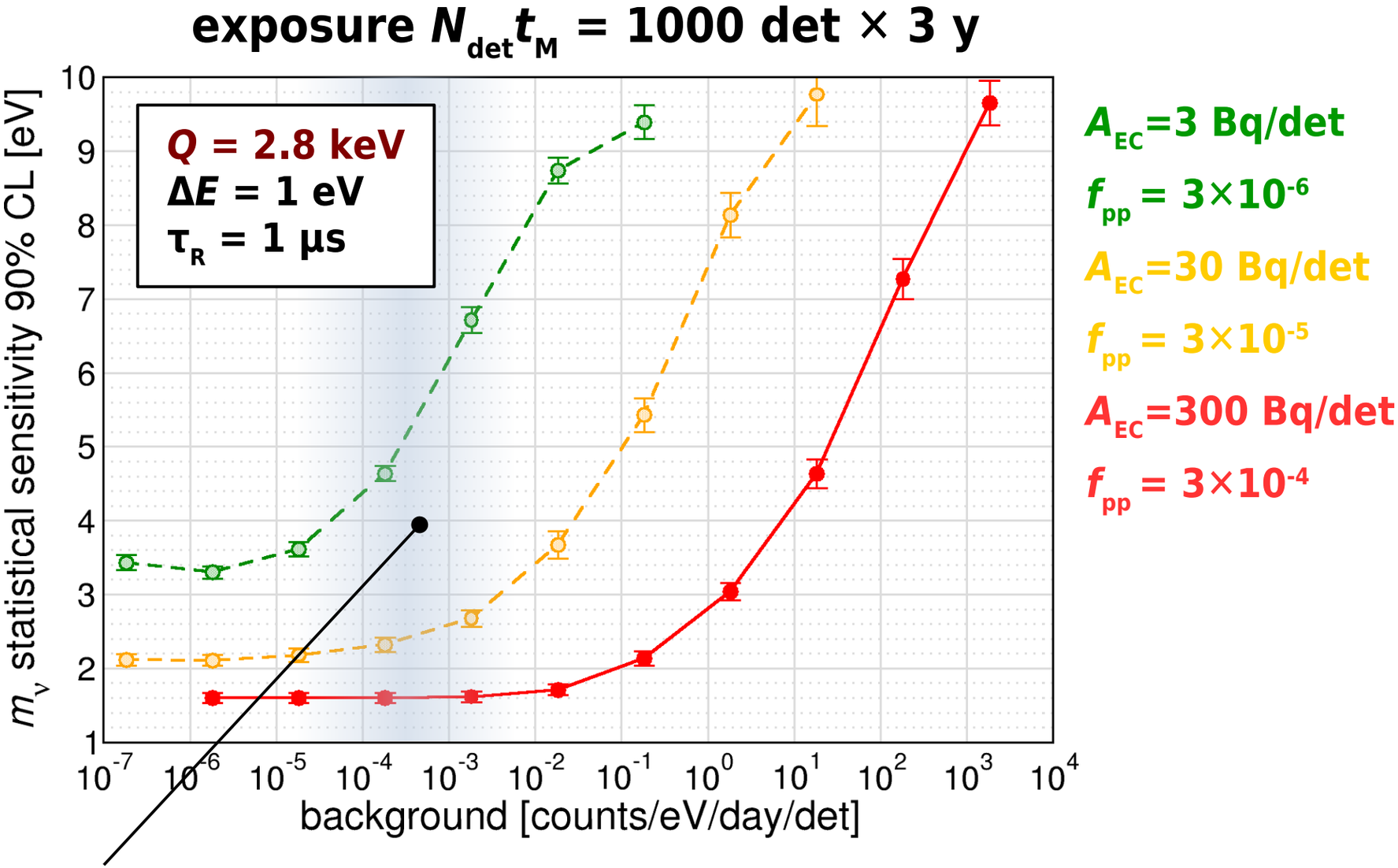




$A(^{163}\text{Ho})$ [Bq]	f_{pp}	b_{pp} [c/eV/day]	max b ‡ [c/eV/day]	max $A(^{166m}\text{Ho})$	$A(^{163}\text{Ho})/A(^{166m}\text{Ho})$	$N(^{163}\text{Ho})/N(^{166m}\text{Ho})$
3	3×10^{-6}	3.2×10^{-6}	10^{-5}	3×10^{-5}	10^5	**4×10^5**
300	3×10^{-4}	3.2×10^{-2}	10^{-1}	0.3	1000	**4000**

‡ from MC simulations

Effect of flat background on sensitivity



expected from simulations and preliminary HOLMES and ECHo measurements

A. Goeggelmann et al., EPJ.C 81 (2021) 363

A. Goeggelmann et al., EPJ.C 82 (2022) 139

The ECHo experiment

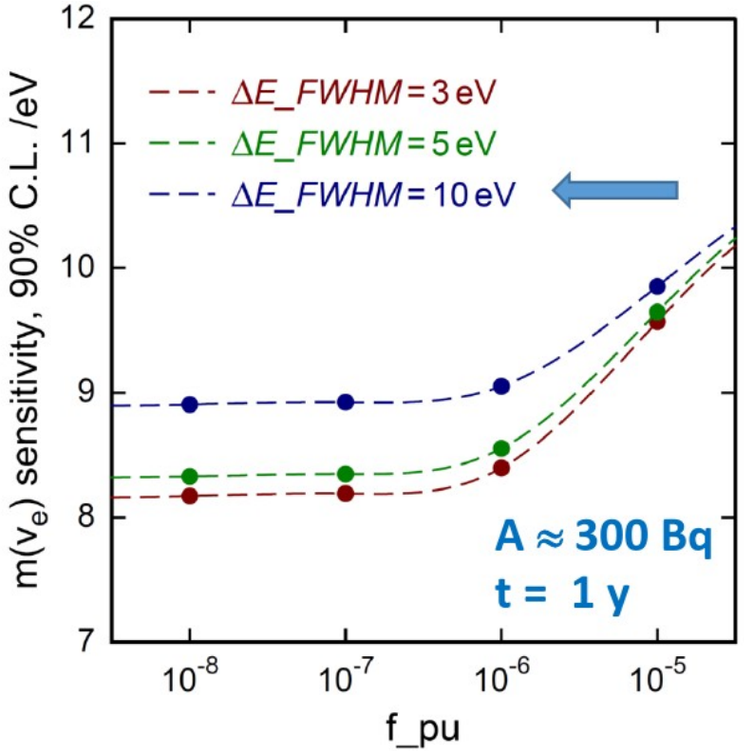
Arrays of Magnetic Metallic Calorimeters with ion-implanted ^{163}Ho

L. Gastaldo et al. Eur. Phys. J. Special Topics 226, 1623 (2017)

ECHo-1k

- number of detectors: 60~100 pixels
- activity: 1~5 Bq/pixel
- read-out: two-stage dc-SQUID
- energy resolution: ΔE_{FWHM}<10 eV

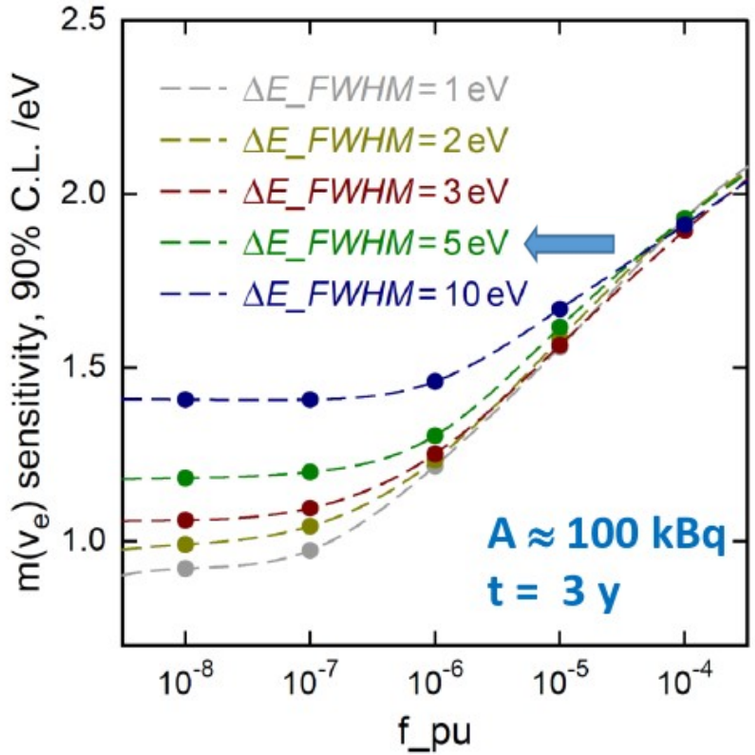
→ **m_ν statistical sensitivity <20 eV**



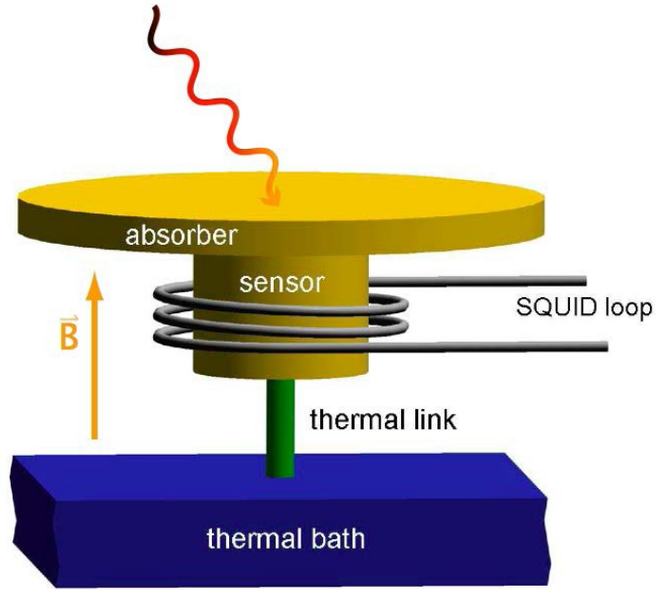
ECHo-100k

- number of detectors: 12000 pixels
- activity: 10 Bq/pixel
- read-out: microwave multiplexing
- energy resolution: ΔE_{FWHM}<5 eV

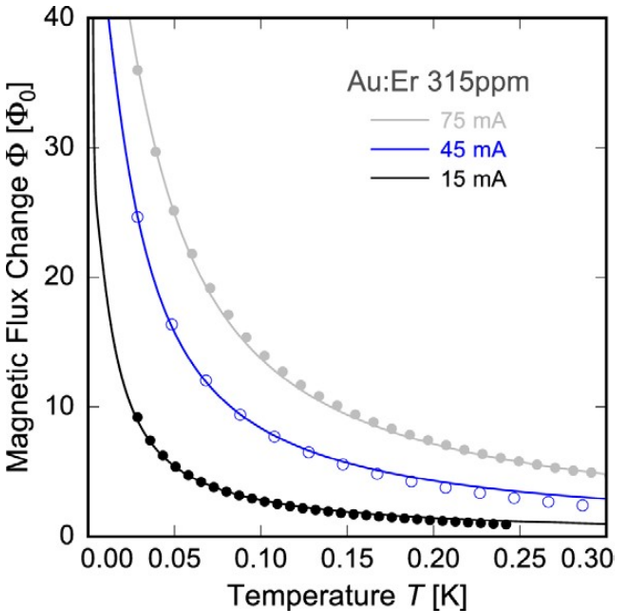
→ **m_ν statistical sensitivity <1.5 eV**



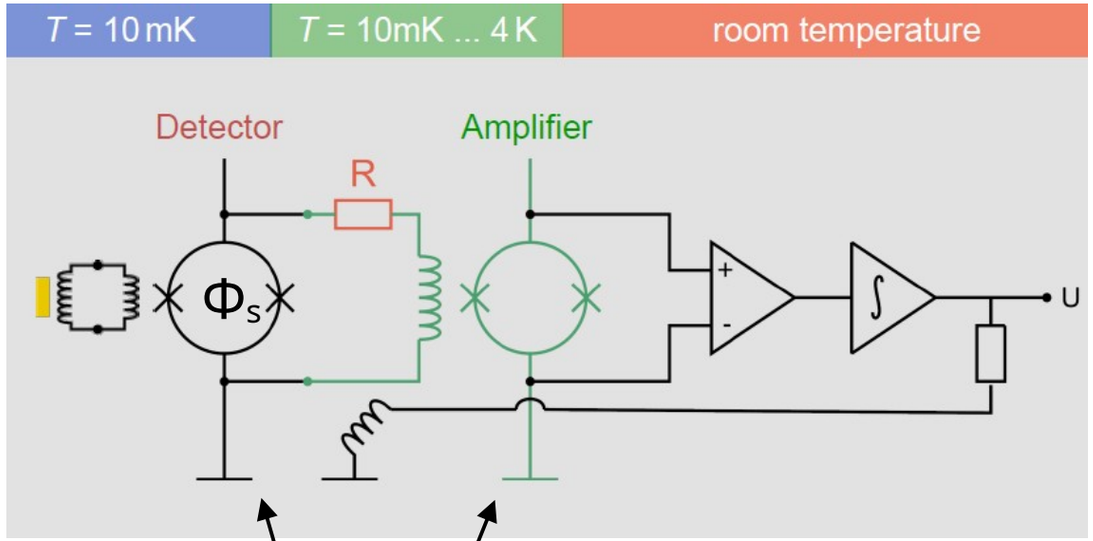
Metallic Magnetic sensors



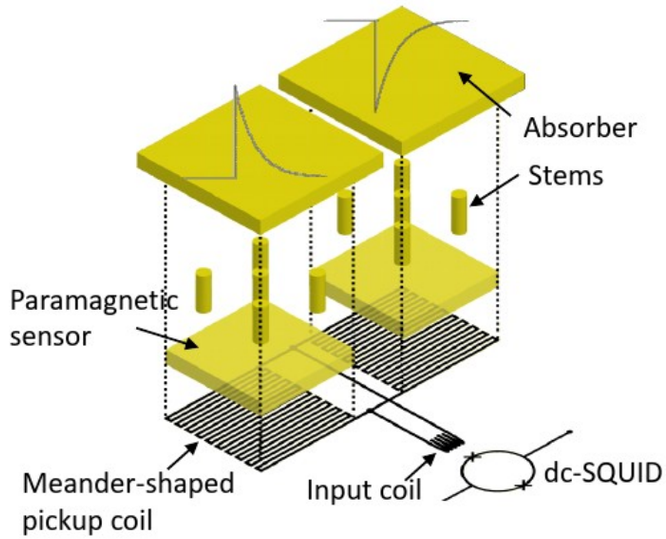
- paramagnetic temperature sensor
- **Ag:Er**, Au:Er ...
- no power dissipation in the sensor no Johnson noise
- SQUID read-out → multiplexing for arrays
- large operating temperature range 10 - 100 mK
  - ECHo operates at $T \approx 30$ mK
- strong spin-electron coupling → fast rise time (≈100 ns)



$\Delta T \rightarrow \Delta M \rightarrow \Delta \Phi_s$



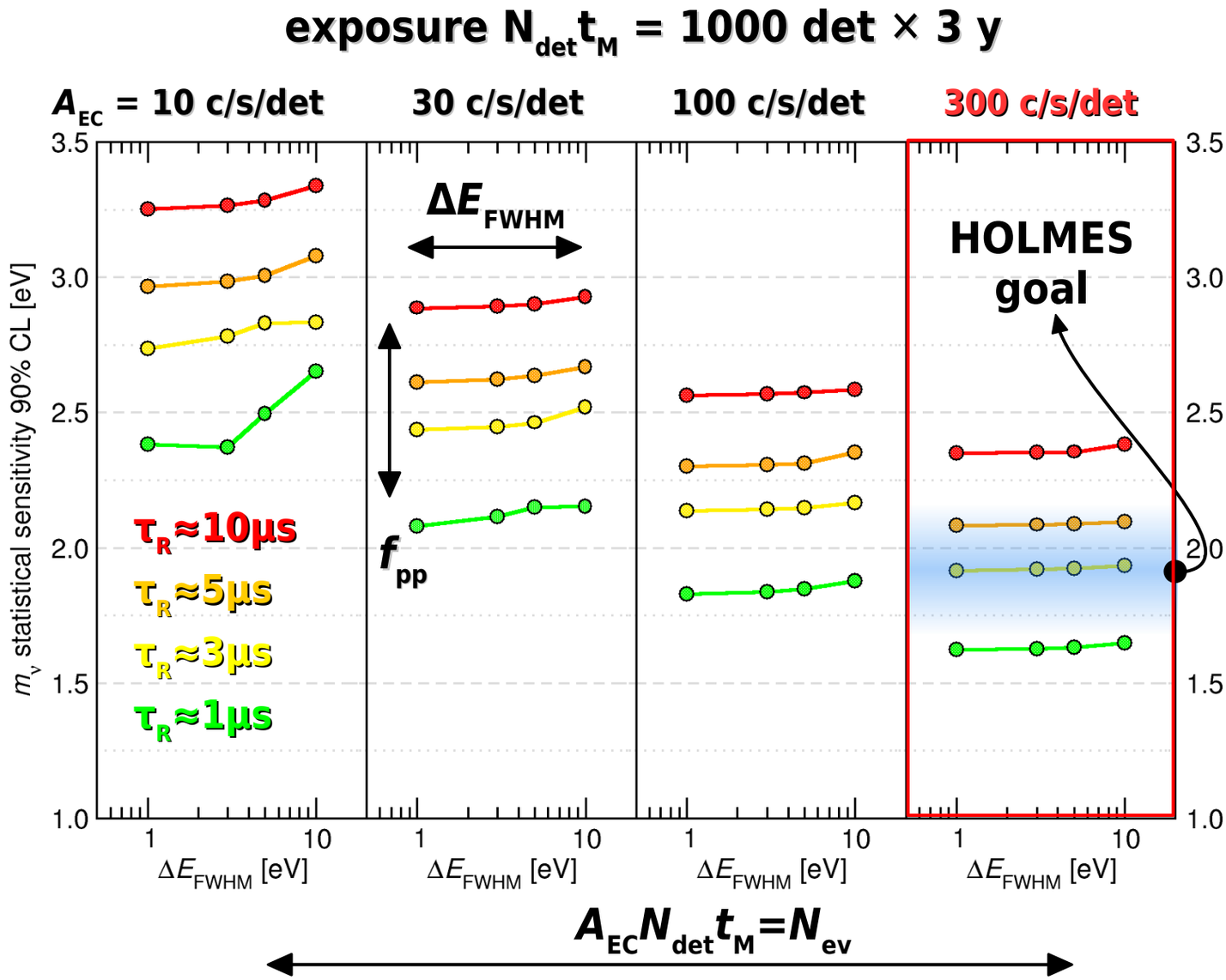
dc-SQUIDs



gradiometric configuration
2 MMC pixels on 1 SQUID channel

The HOLMES experiment

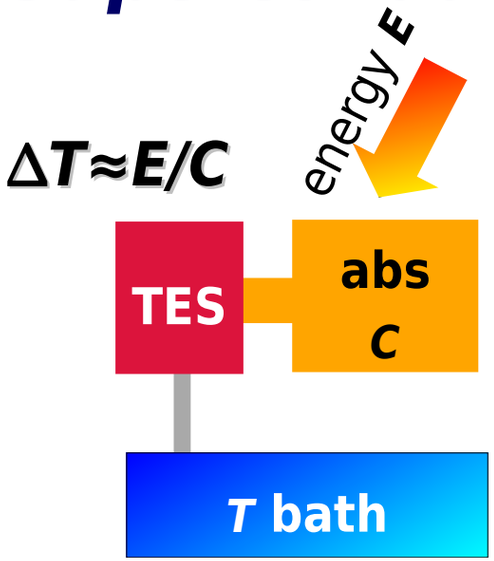
- Transitione Edge Sensors (TES) microcalorimeters with ion-implanted ^{163}Ho
- 6.5×10^{13} atom/det → A_{EC}=300 c/s/det
- $\Delta E \approx 1$ eV and $\tau_R \approx 1$ μs
- 1000 TES microcalorimeters
- 16 × 64-pixel arrays with microwave multiplexed read-out
- 6.5×10^{16} ^{163}Ho nuclei → ≈18 μg
  → 3×10^{13} events in 3 years
  → **m_ν statistical sensitivity ≈1 eV**



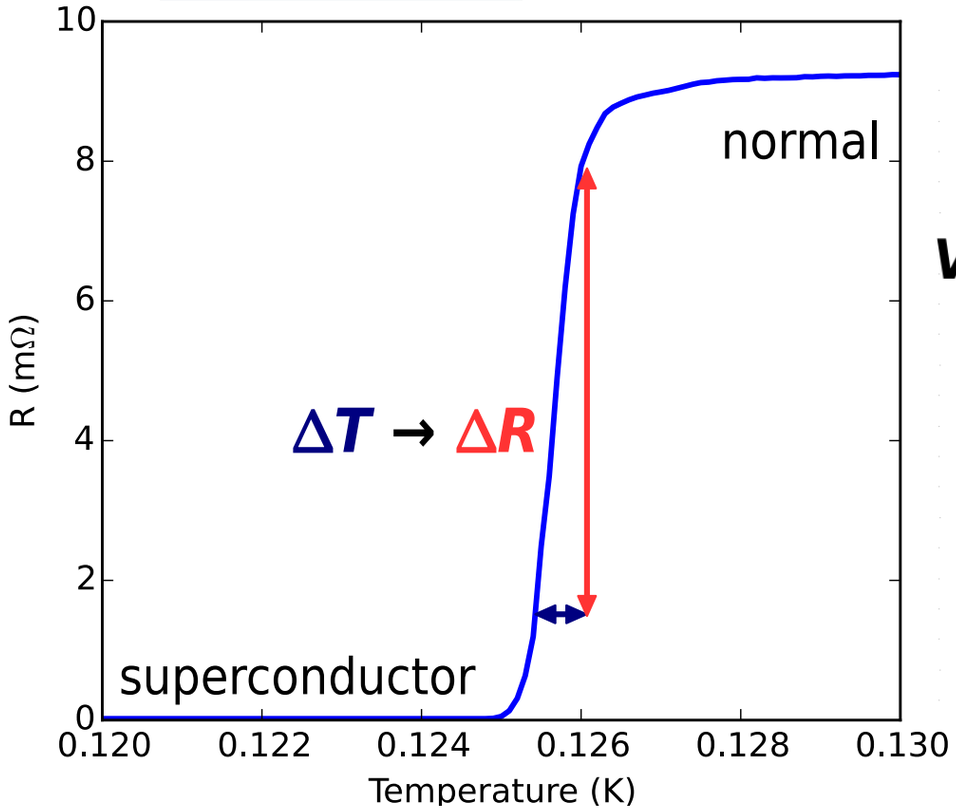
HOLMES

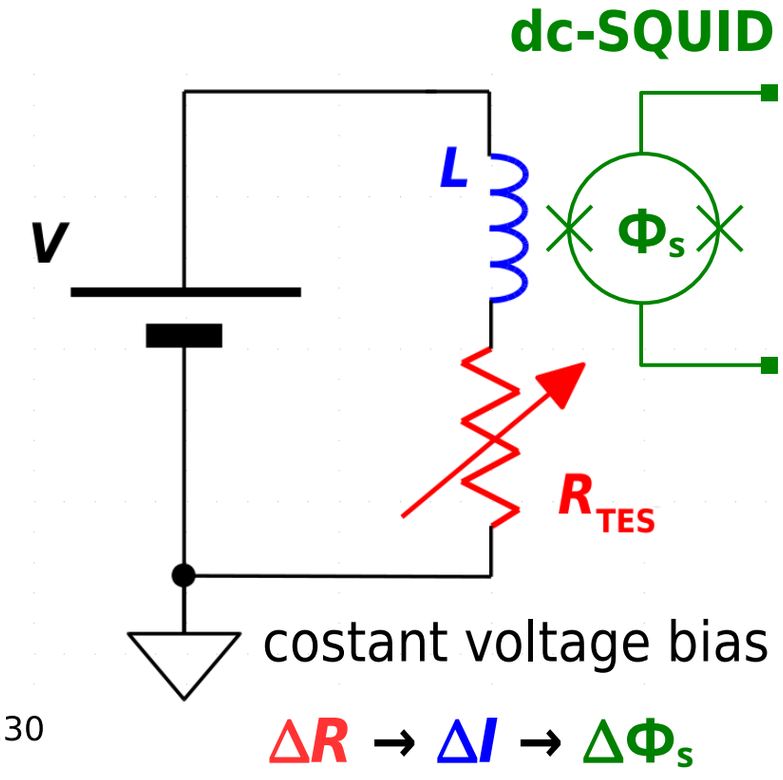
B. Alpert et al., Eur. Phys. J. C, (2015) 75:112

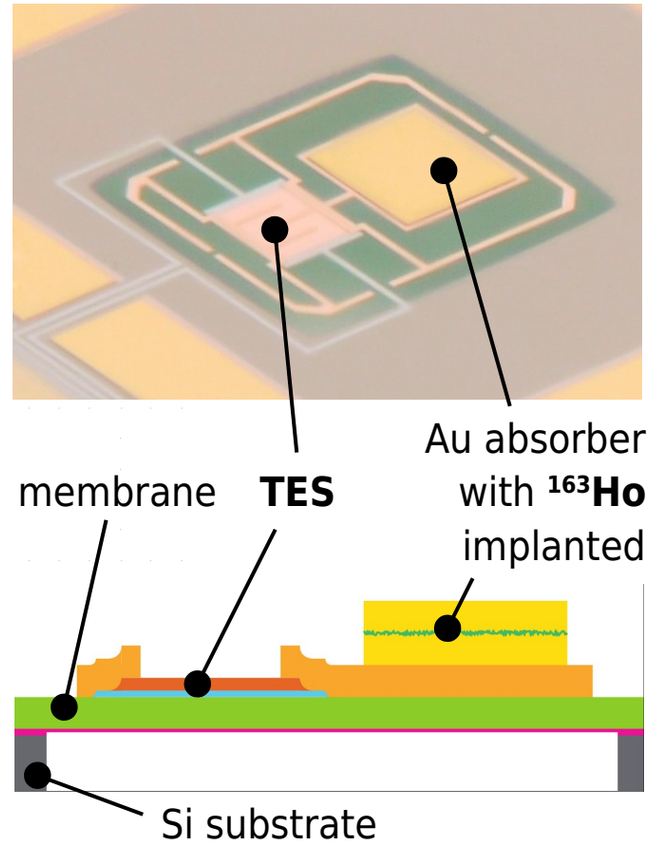
Superconducting transition edge sensors (TES)



- superconducting thin films operated inside the phase transition at T_c
  - HOLMES: Mo/Cu bilayer tuned for $T_c \approx 100$ mK
- high sensitivity $T\mathrm{d}R/(R\mathrm{d}T) \approx 100$ → high energy resolution $\sigma_E^2 \approx \xi^2 k_B T^2 C$
- strong internal coupling → high intrinsic speed
- low impedance (mΩ → Ω) → SQUID read-out → multiplexing for arrays

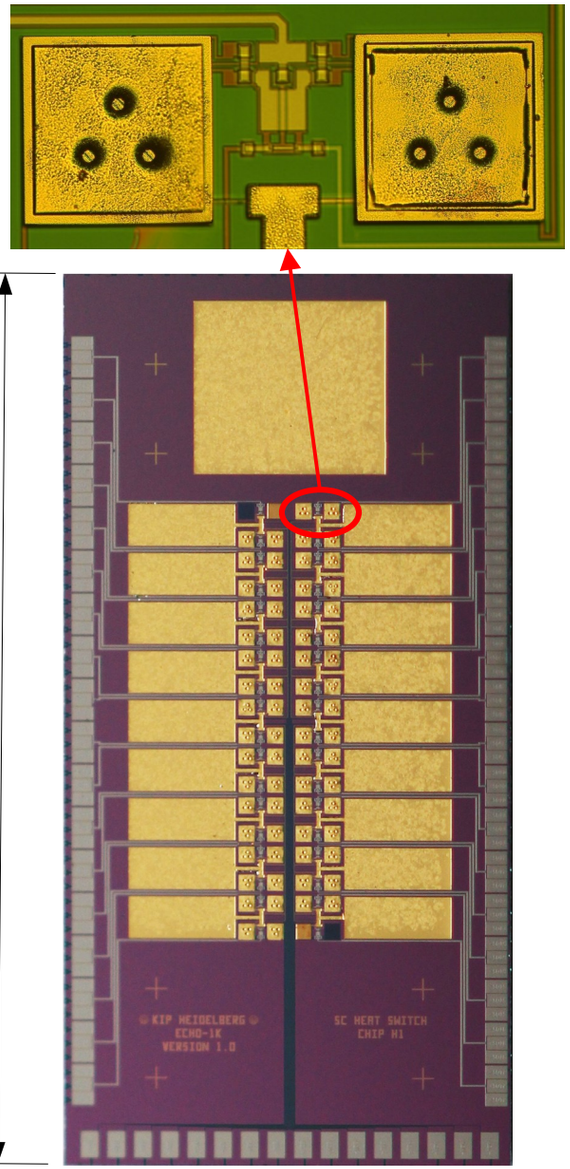






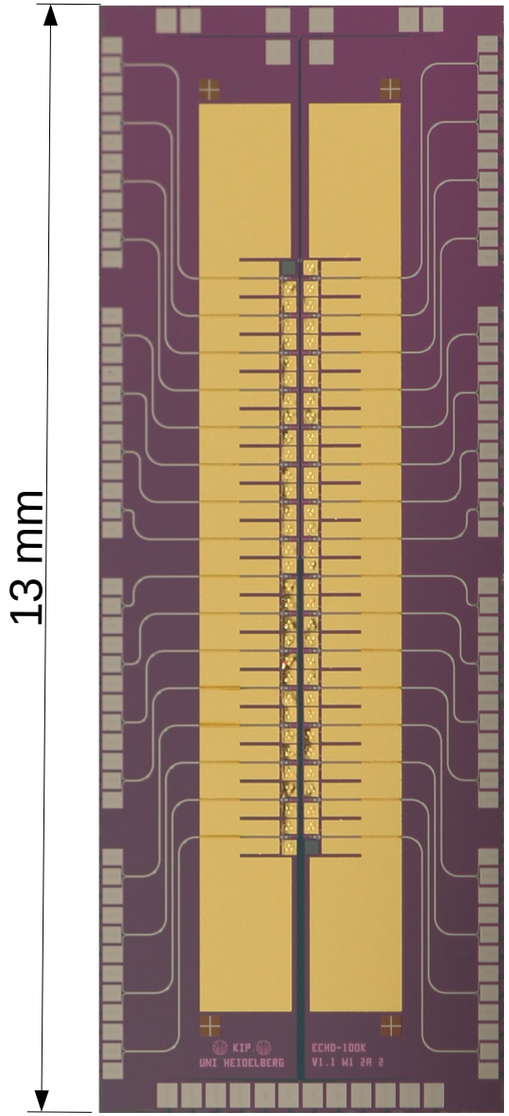
Microcalorimeters arrays

ECHo-1k



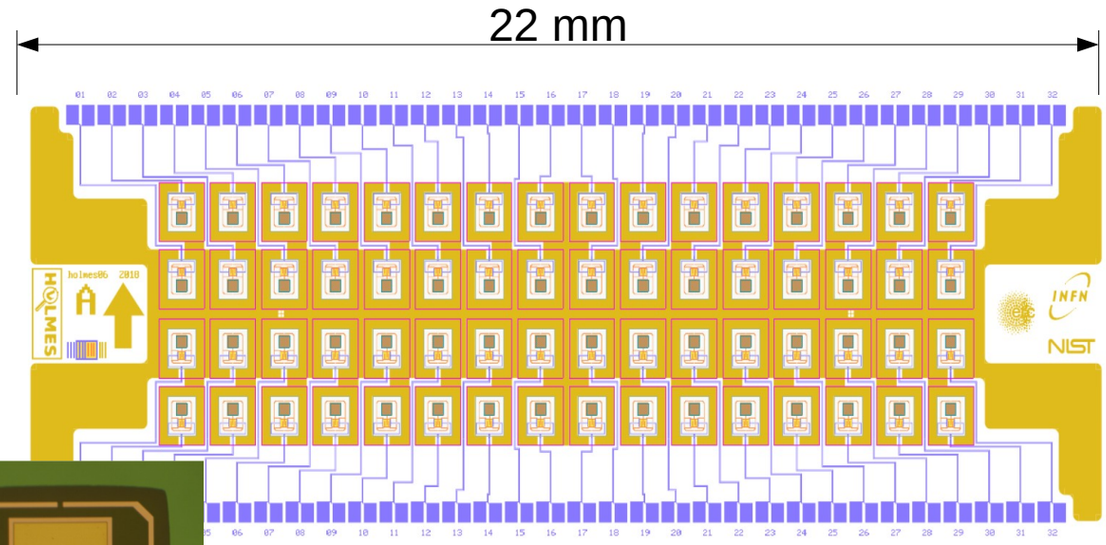
72 pixels

ECHo-100k

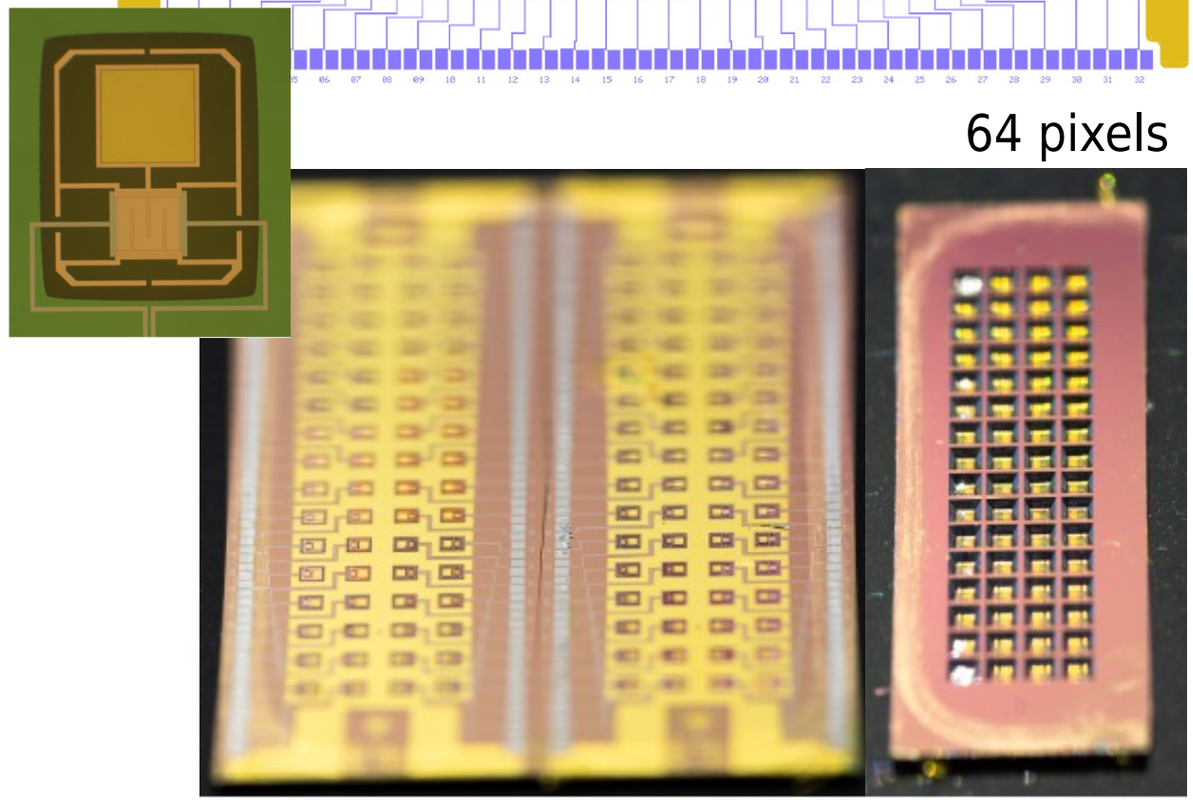


64 pixels

HOLMES



64 pixels



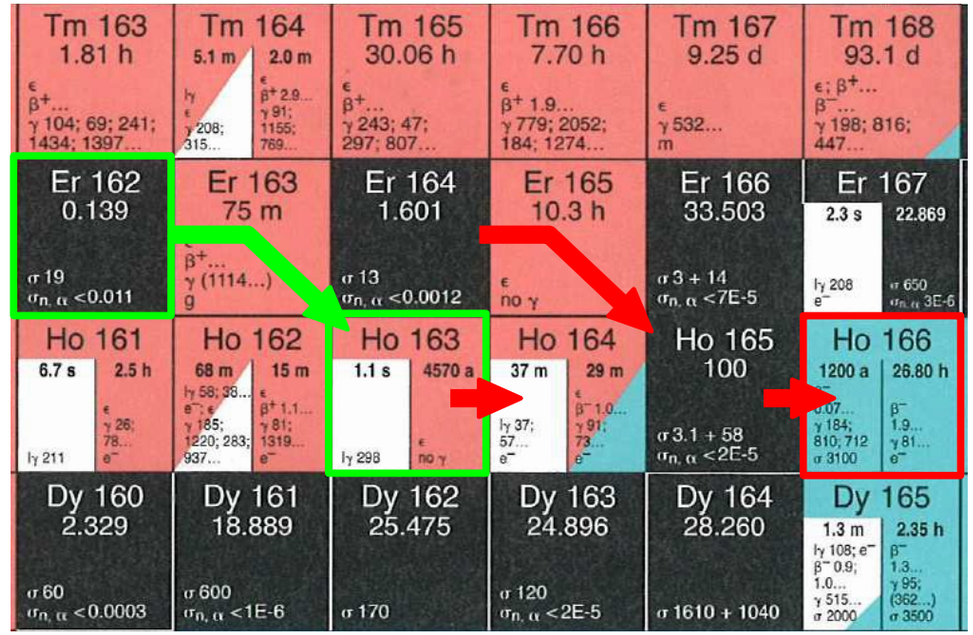
KOH micromachining

Isotope production

^{162}**Er (n,γ)** ^{163}**Er** $\quad \sigma_{thermal} \approx 20$ b

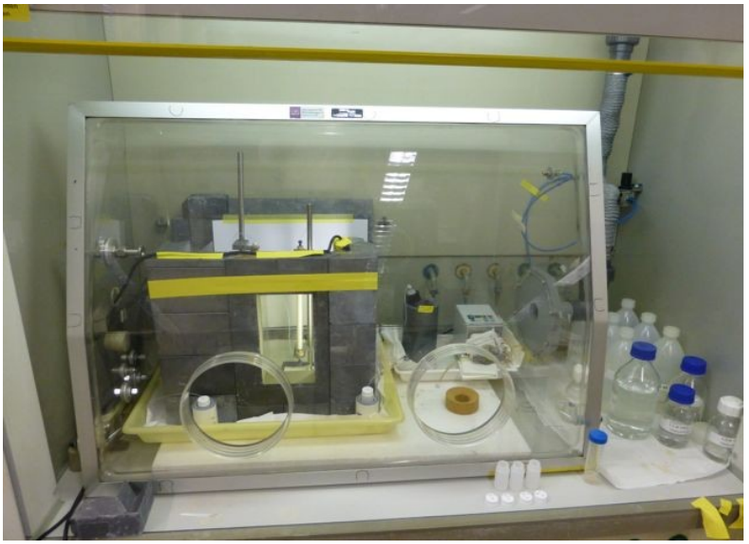
^{163}**Er** → ^{163}**Ho** + ν_e $\quad \tau_{1/2}^{EC} \approx 75$ min

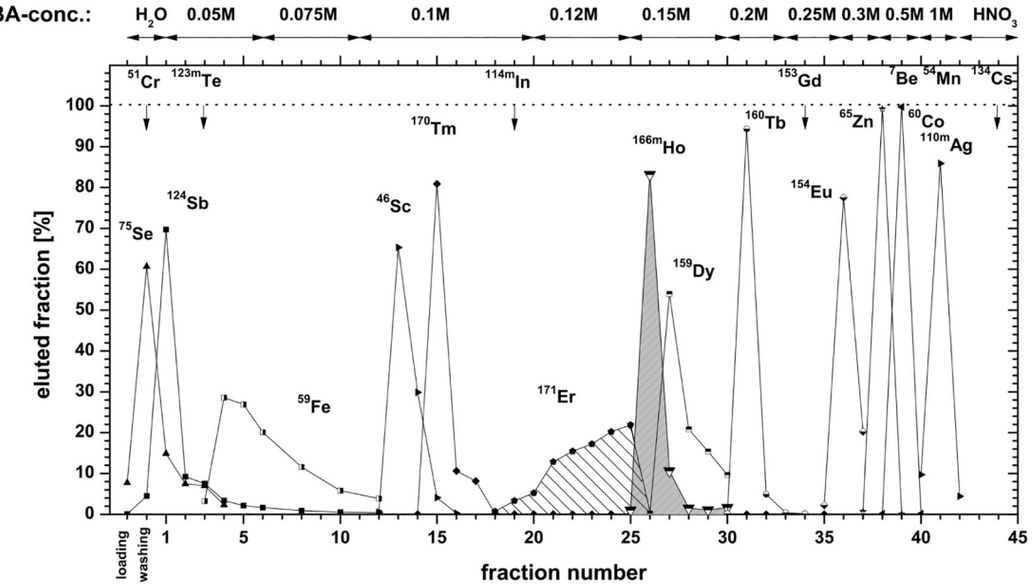
- ^{162}Er irradiation at **ILL nuclear reactor** (Grenoble, France)
  - ▶ thermal neutron flux 1.3×10^{15} n/cm^2/s
- Ho chemical separation with ion-exchange resins in hot-cell to remove Er matrix and and radioactive products
- separation efficiency >90 %
- HOLMES has collected ≈200 MBq of ^{163}Ho (+ ≈400 kBq of ^{166m}Ho)

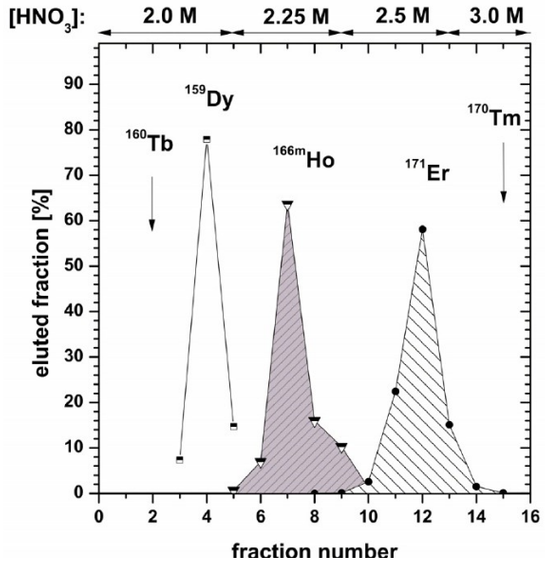


H. Dorrer et al., Radiochim. Acta, 106 (2018) 535

S. Heinitz et al., PLoS ONE 13(8): e0200910



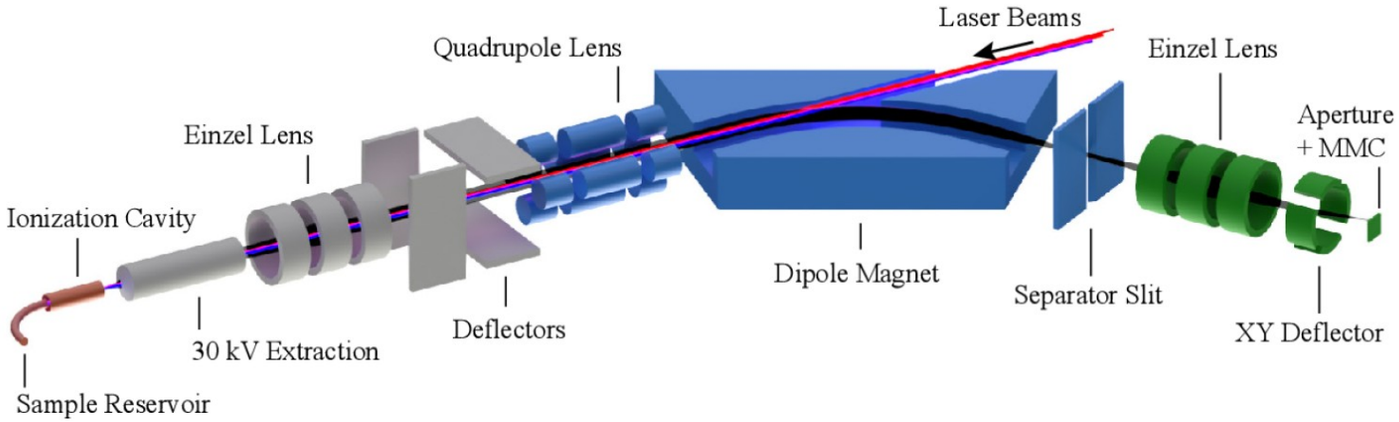




Mass separation and isotope embedding

^{166m}Ho must be separated by magnetic mass spectrometer

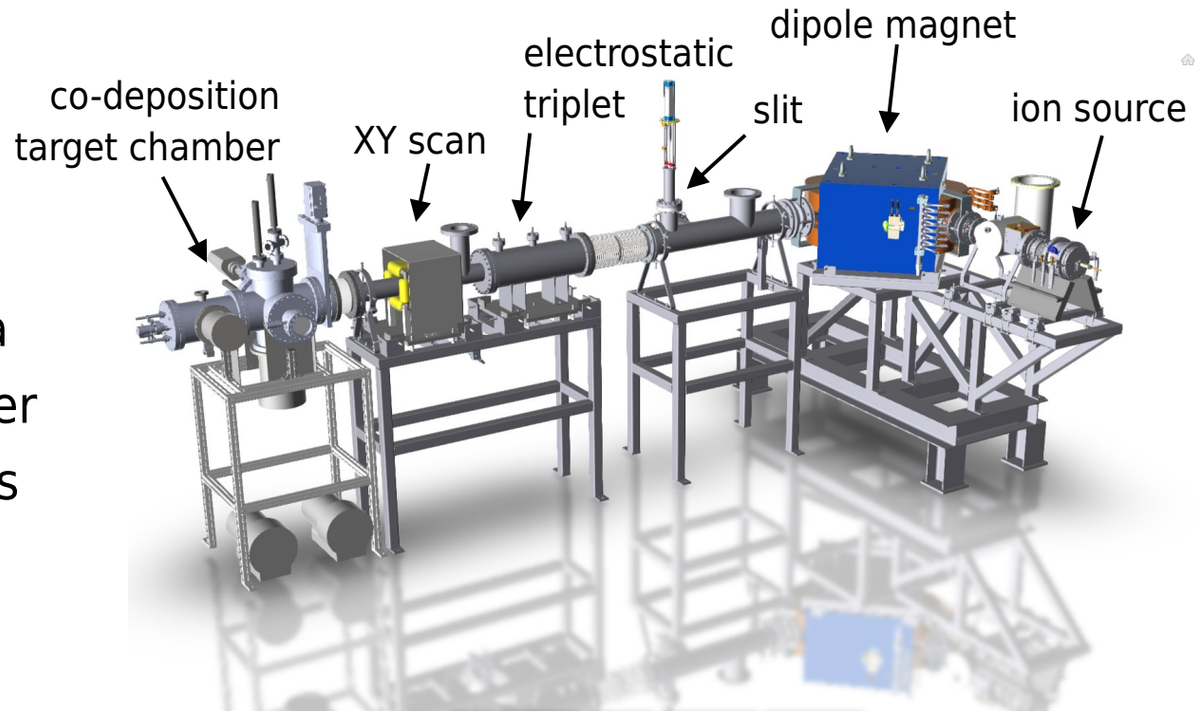
requires high current, high source and geometrical efficiency



ECHo: resonant ionization laser ion source (RILIS)

- RISIKO at Mainz University
- efficiency: (69 ± 5stat ± 4syst) %
- ^{166m}Ho/^{163}Ho < 4(2)10^{-9}

T. Kieck et al., NIM A, 945, 2019, 162602.

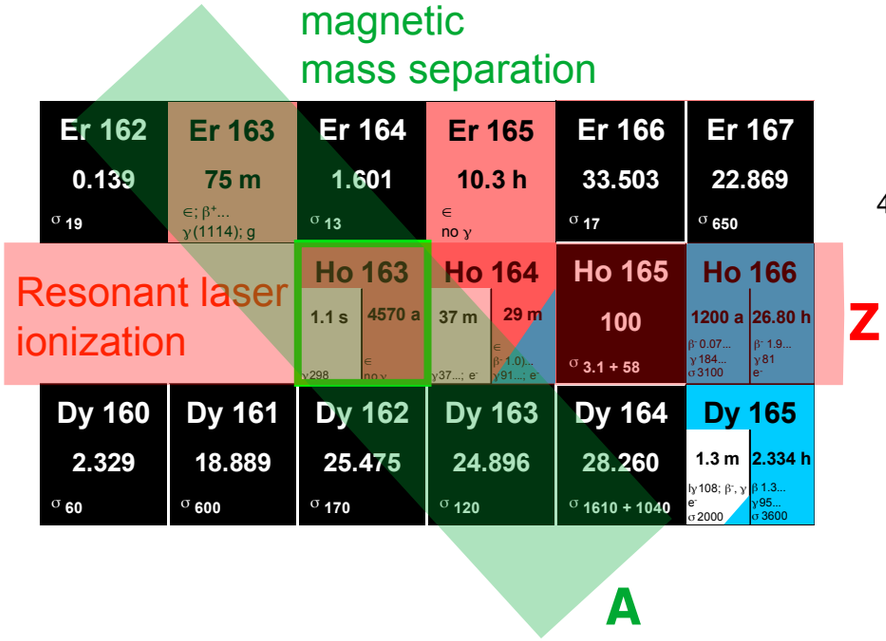


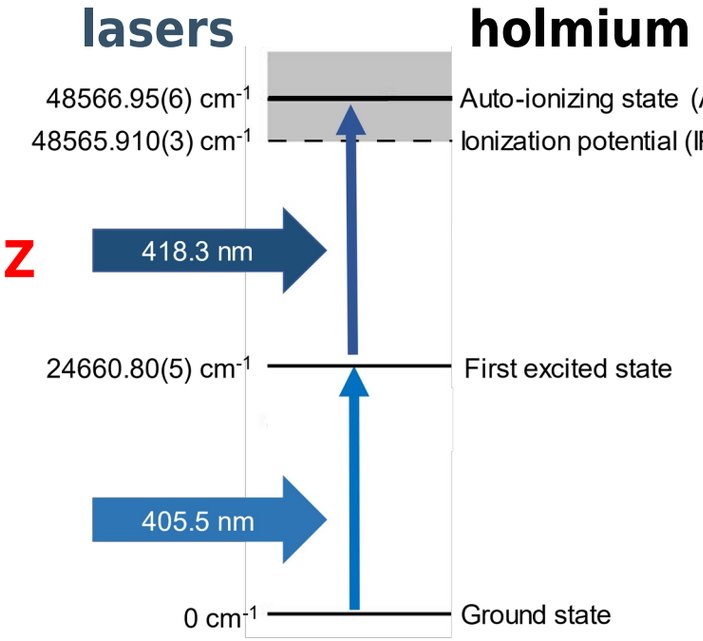
HOLMES: Ar plasma sputter ion source

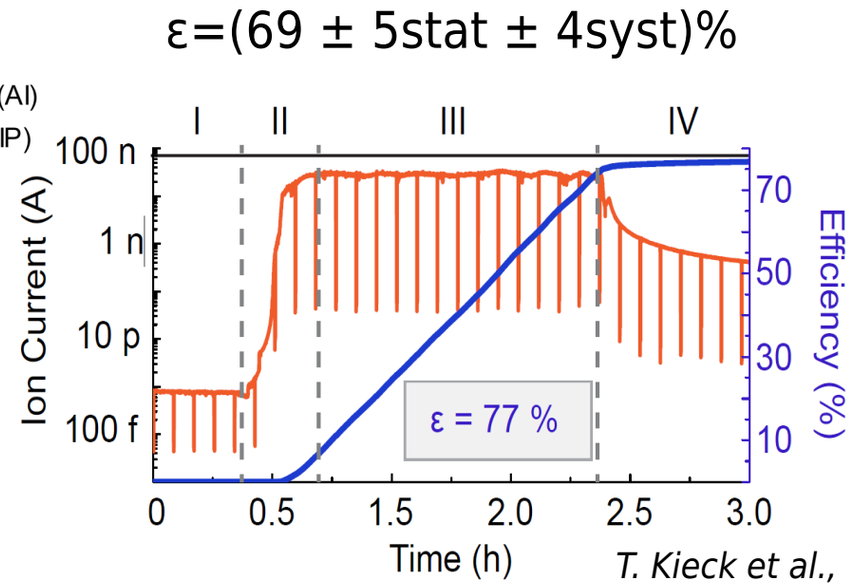
- all components ready for installation in Genova
- now testing without triplet/XY-scan and chamber
- high current ion source optimization in progress

RISIKO for ECHo

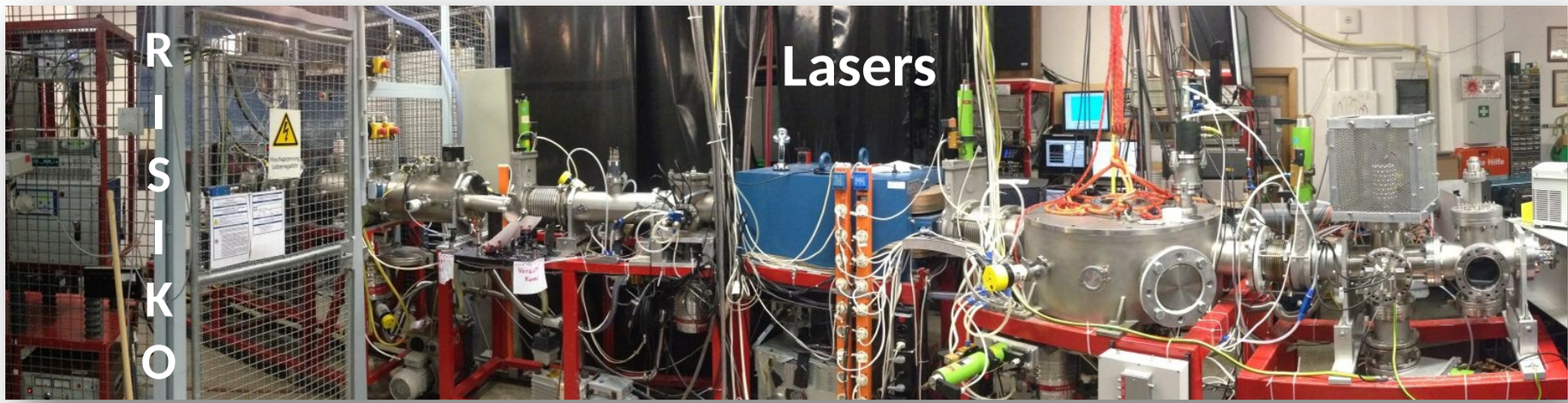
- Resonant laser ionization
  → 2 lasers tuned to selectively ionize Ho
  → high ionization efficiency ion source







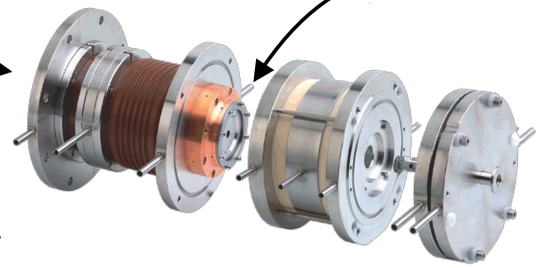
T. Kieck et al., NIM A, 945, 2019, 162602.

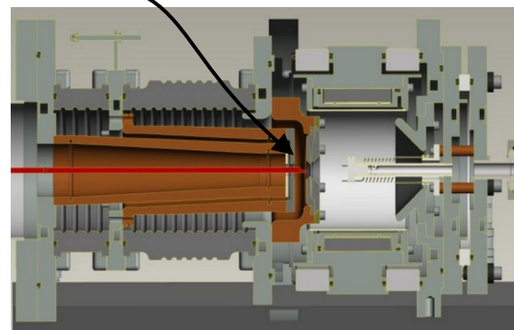


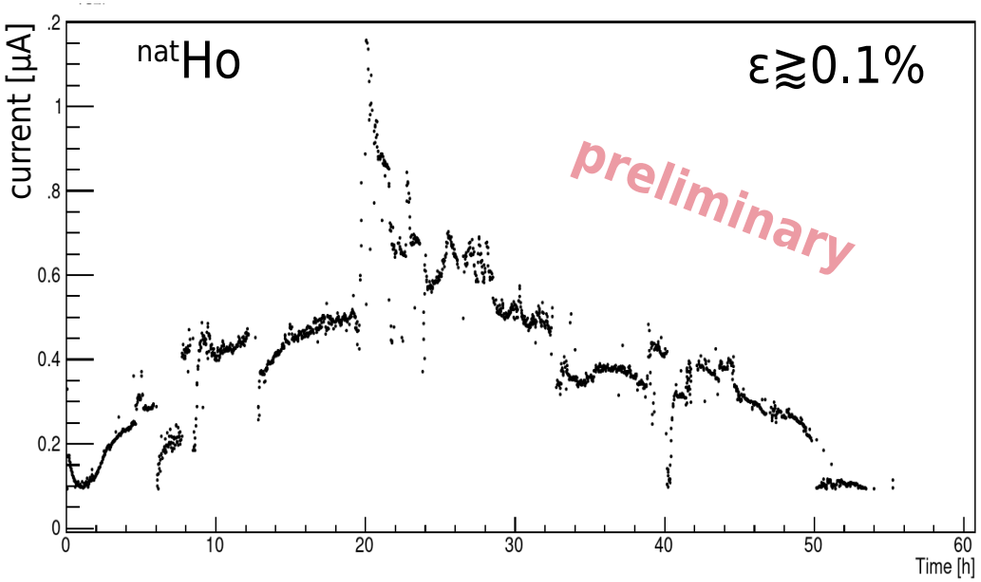
HOLMES ion implanter progresses

- high current sputter source
- several sputter targets tested
- sintered Zr-Y-Al-Ho$(NO_3)_3$

  → highest effieciency, current and stability

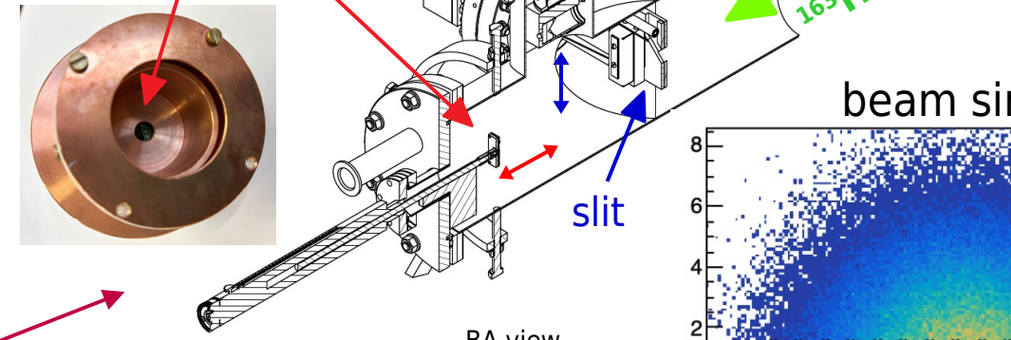


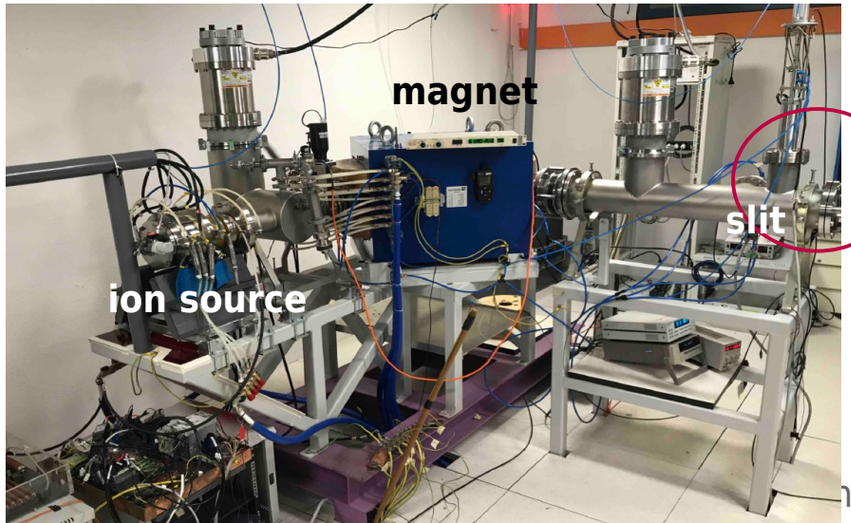




configuration without focussing

for first low activity implants





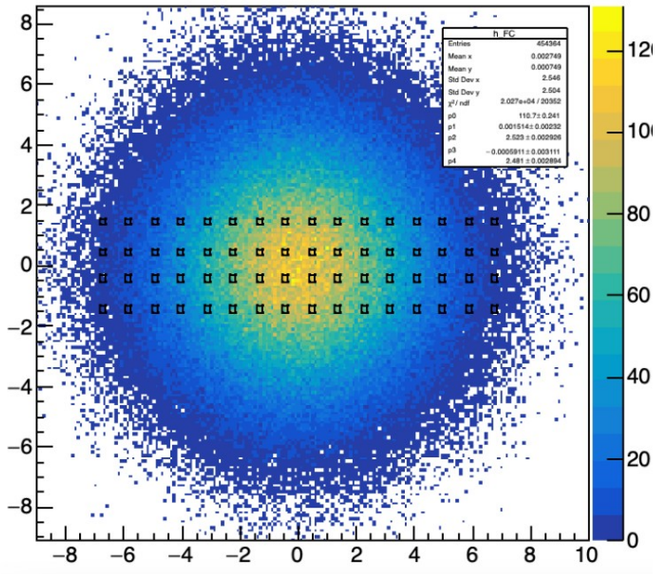
8 MBq of ^{163}Ho in ^{nat}Ho (0.35%)

$I(^{nat}Ho)$=0.2 μA for 24 h

→ peak pixel activity ≈1 Bq

uniform activity with 3 shots

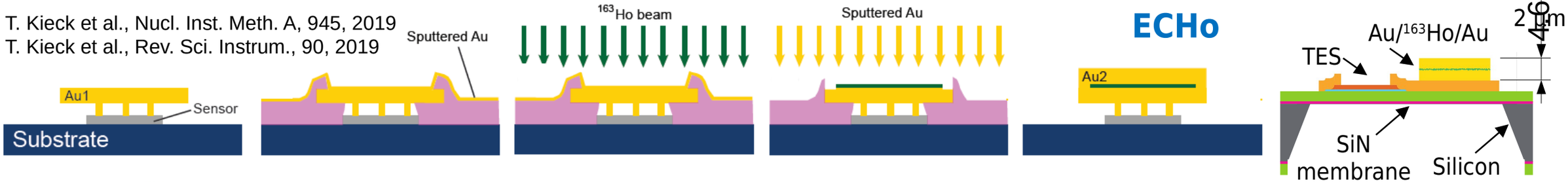
beam simulation

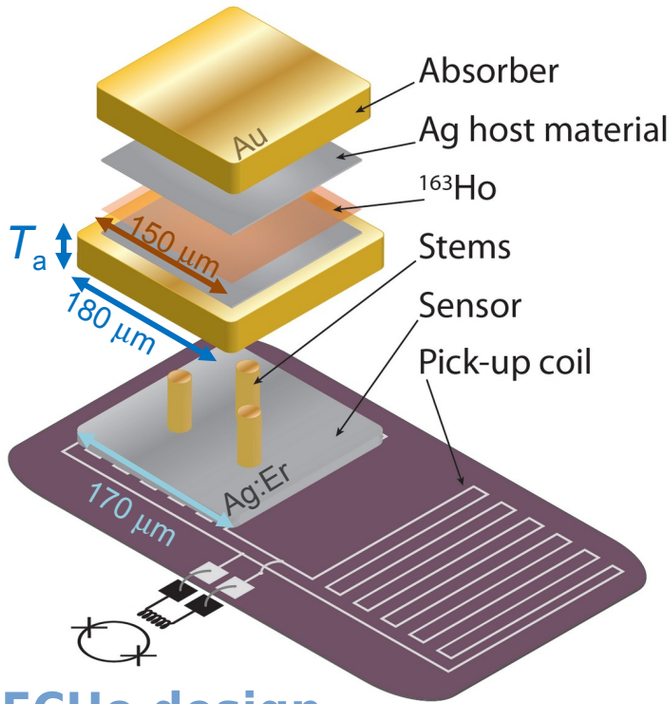


Detector absorbers for calorimetry

- Au absorber must stop all radiation from atomic de-excitations with $E_c \approx Q$
  - for H=M1 → 3-4 Auger/C-K electrons carry most of E_c (the most energetic with $\langle E_e \rangle \approx 2$ keV)
  - for H=M1 → rarely ($\omega_M \approx 10^{-3}$) one X with $\langle E_X \rangle \approx 2.5$ keV and low energy electrons
  - shake-off electrons have energies mostly $\lesssim 800$ eV

T. Kieck et al., Nucl. Inst. Meth. A, 945, 2019
T. Kieck et al., Rev. Sci. Instrum., 90, 2019

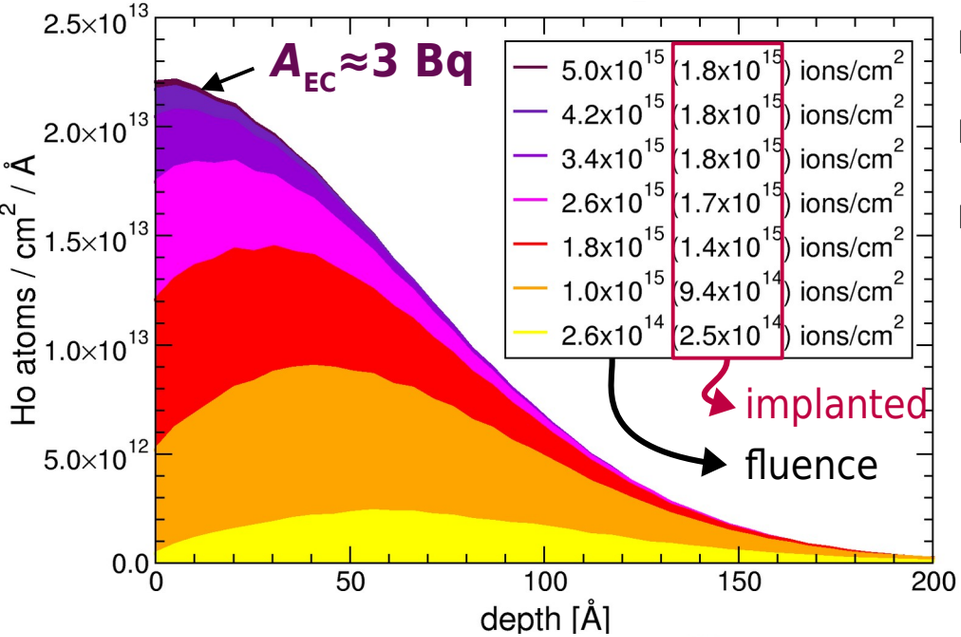




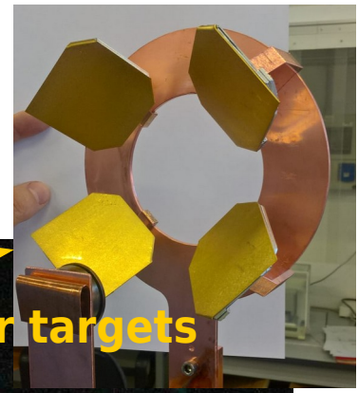
ECHo design

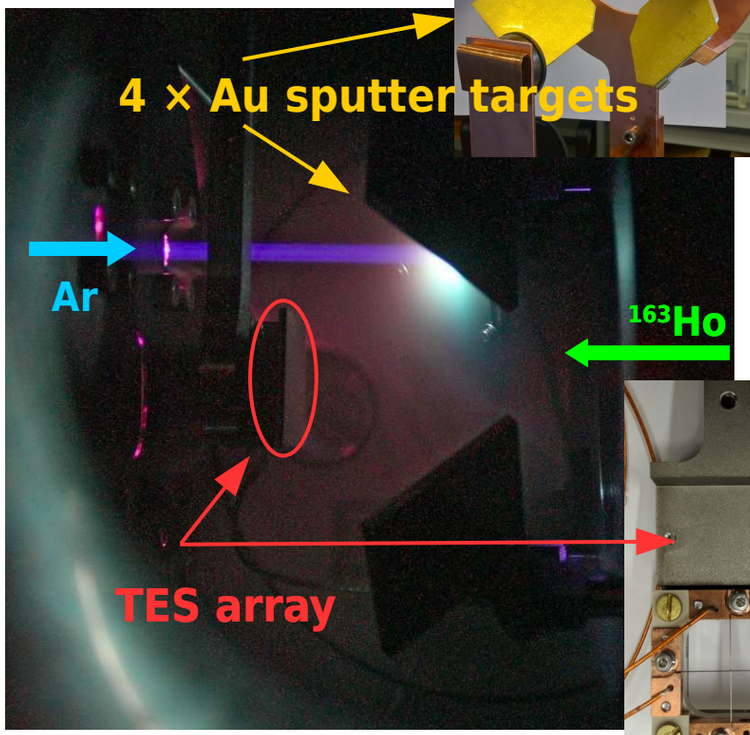
- **ECHo**: T_a=3 µm
  - electrons: fully contained (no simulation)
  - X-rays (2.6 keV<E_X<Q): 1.5×10^{-5} escaping ("raytracing" sim)
- **HOLMES**: T_a=1 µm (Geant4 MC simulations for $E_{X,e}=Q$)
  - fully implanted surface (no containement border)
  - electrons: 1.5×10^{-4} escaping (→ tail with same intensity)
  - X-rays: 4×10^{-3} escaping (→ tail with intensity 1.3×10^{-3})

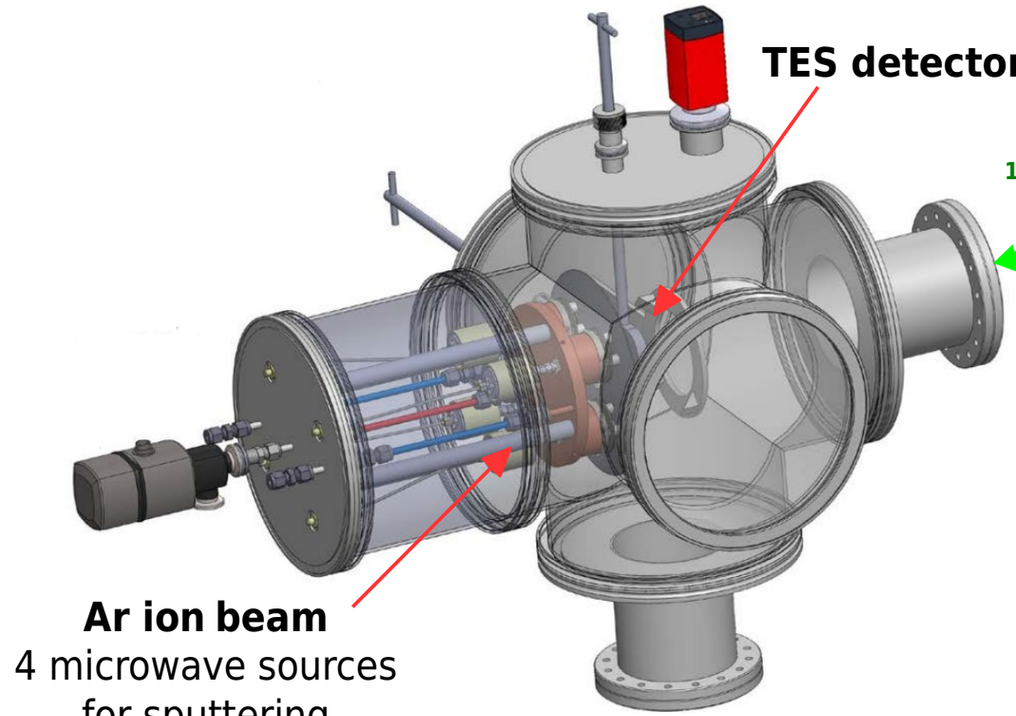
HOLMES co-deposition system

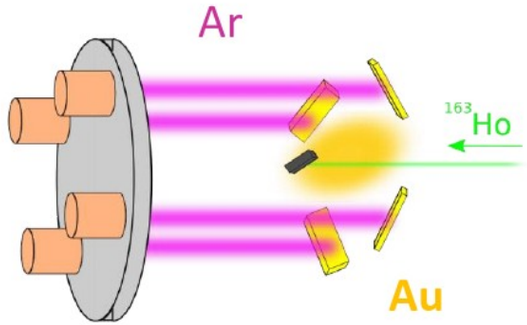


- ion implant simulation with SRIM2013
- **^{163}Ho ions on Au (E_{ion} = 50 keV)**
- ^{163}Ho ion beam sputters off Au from absorber (≈26 Au/Ho)
  - implanted ^{163}Ho saturates at A_{EC}≈3 Bq (HOLMES design)
  - compensate by Au co-evaporation
  - in situ upper 1 µm Au layer deposition
    → with 4 ion sources >100 nm/h









Implanted Ho heat capacity

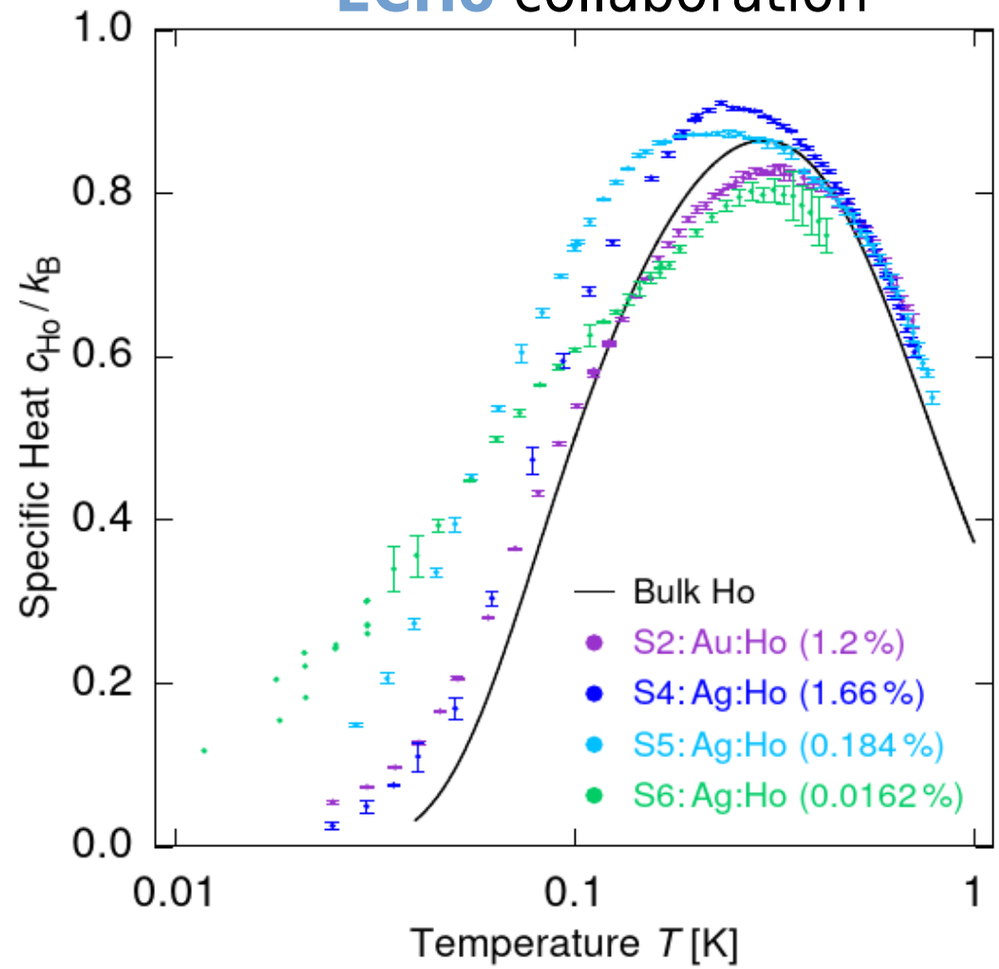
- optimal ΔE depends on C and T

$$\Delta E \propto T\sqrt{C}$$

$$C = C_a + C_{\mathrm{Ho}}$$

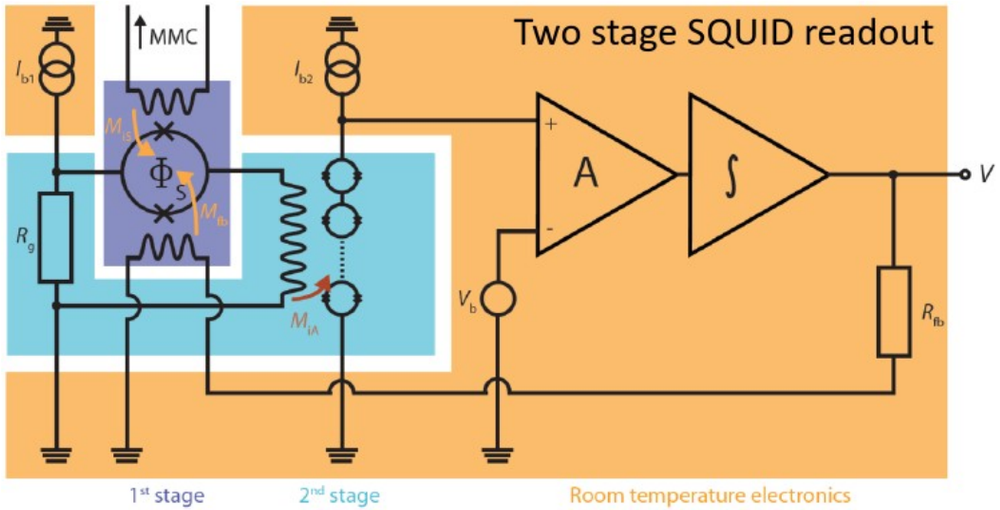
- Ho heat capacity C_{Ho} dominated by a Schottky anomaly at ≈300 mK
  - $\boldsymbol{J}$=8 and $\boldsymbol{I}$=7/2 → hyperfine and crystal field splittings
- contradictory C measurements for Ag:Ho
  - still under investigation
- high activities could be manageable
  - operating at 30 mK or below
  - A=300 Bq → x_{Ho}>10 % → closer to bulk C
  - to be explored by HOLMES

ECHo collaboration



Herbst, M. et al., J Low Temp Phys 202, 106–120 (2021)

ECHo-1k read-out



parallel SQUID read-out

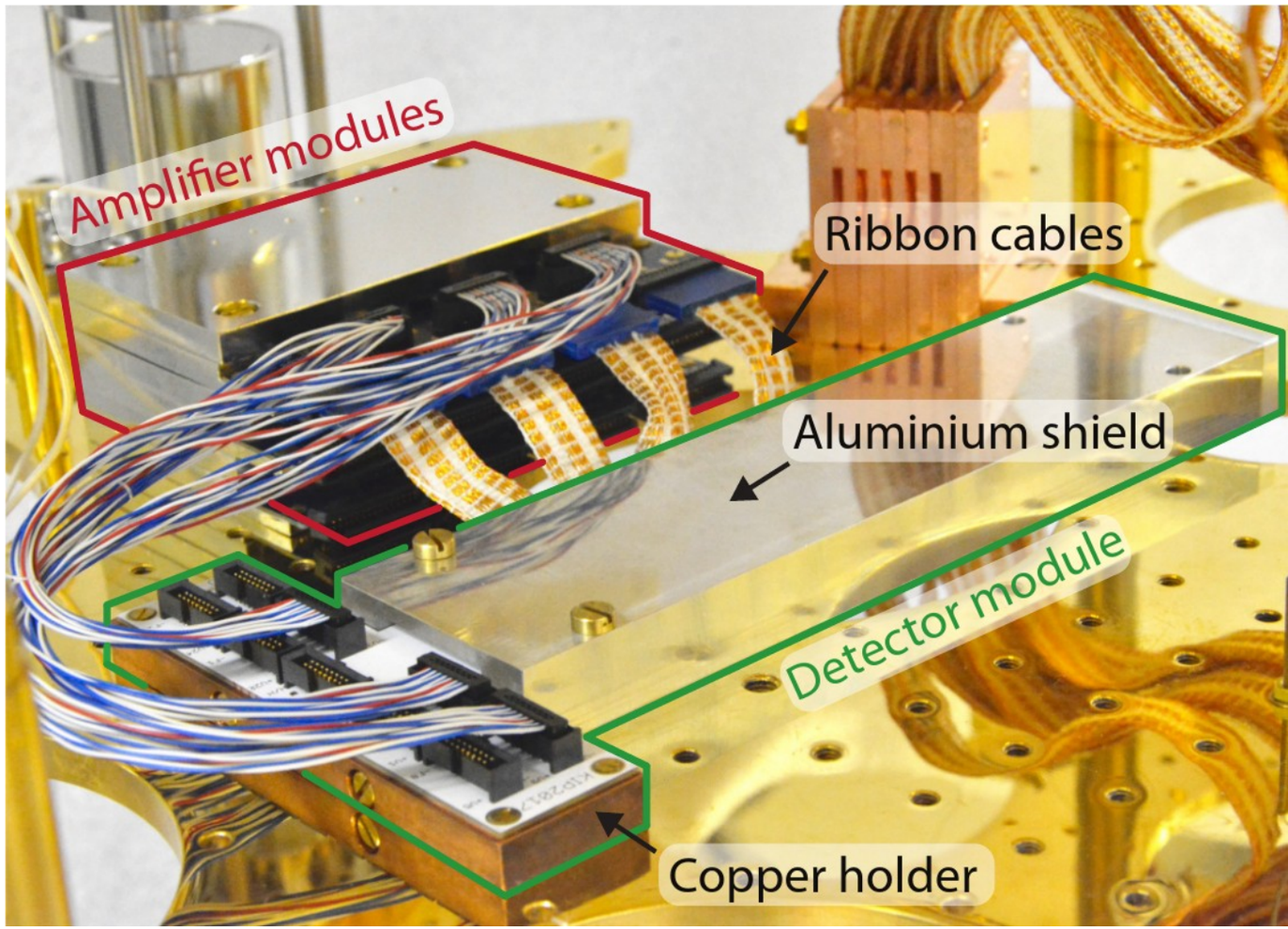
1 SQUID channel for 2 pixels

72 read-out channels for ECHo-1k

10 wires/channel to room temperature

→ **720** wires from low to room temperature

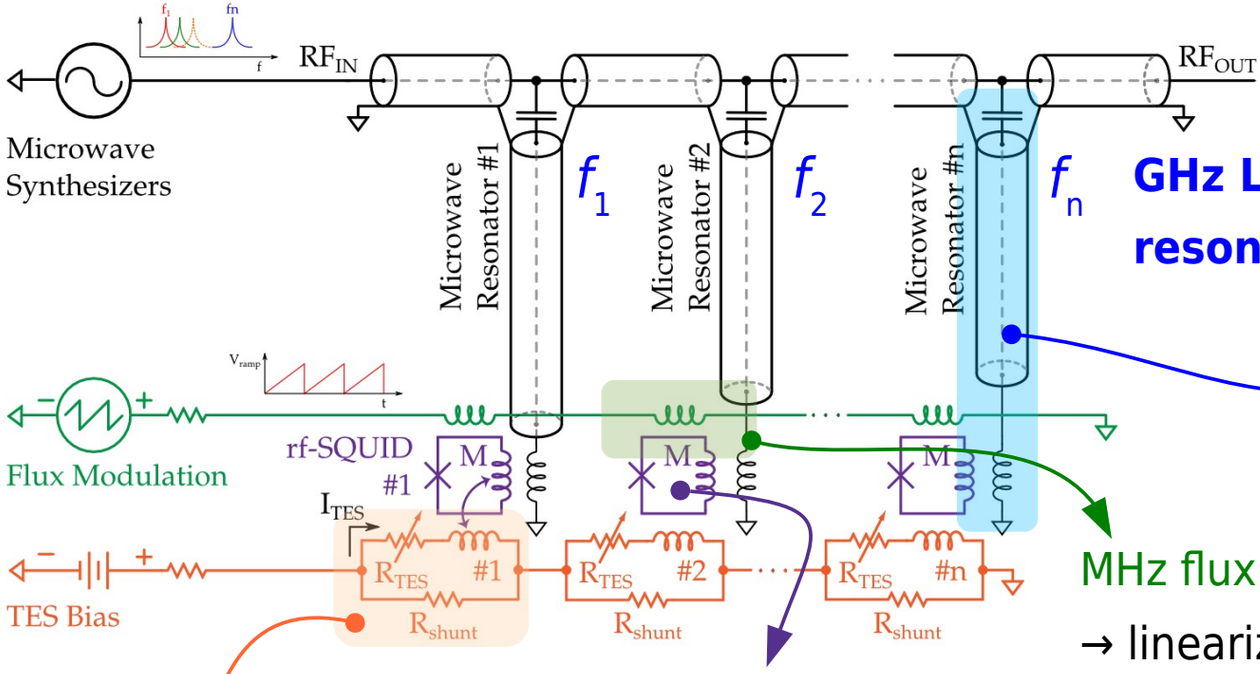
→ not viable for ⩾1000 channels ...

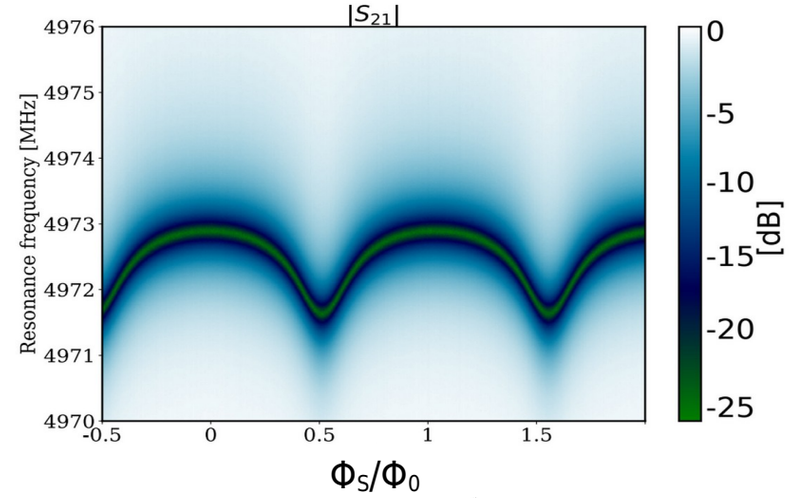


F. Mantegazzini et al. JINST16 (2021) P08003

Microwave multiplexing for array read-out

microwave multiplexing to read-out many detectors with one single RF line and HEMT amplifier





MHz flux ramp modulation $f_{ramp} \rightarrow f_{sampl}$

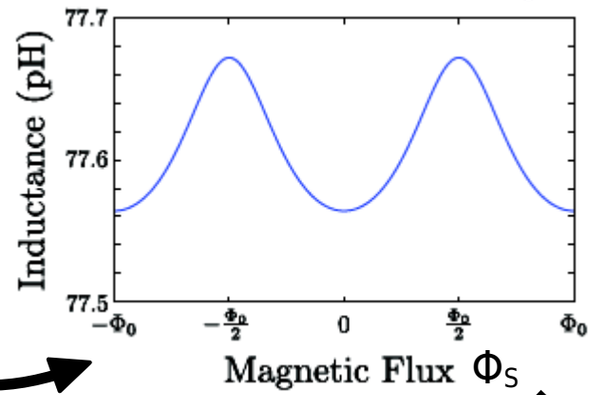
→ linearizes flux-inductance relation

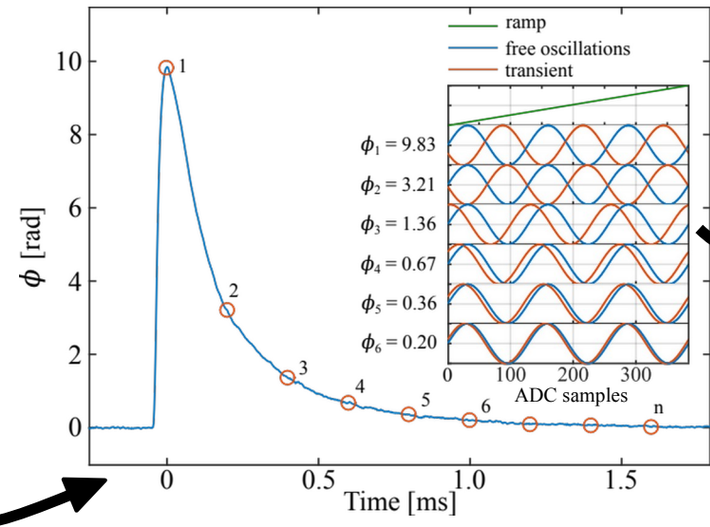
→ converts flux $\Delta\Phi_S$ to phase shift ϕ

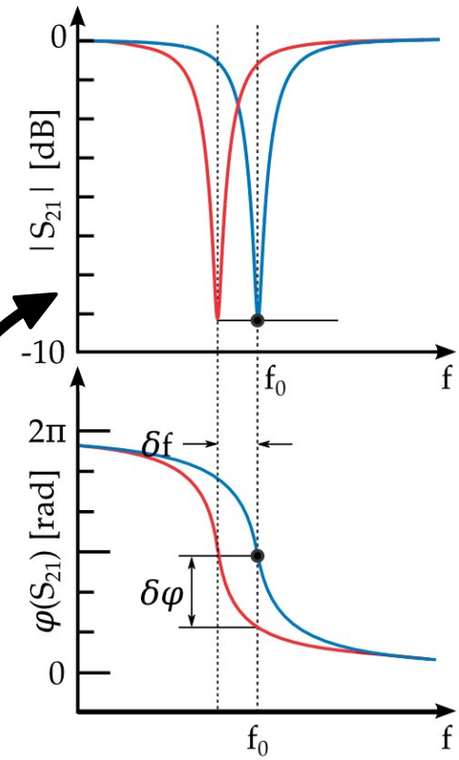
dc-biased TES

$\Delta T \rightarrow \Delta I \rightarrow \Delta\Phi_S$

non-hysteretic rf-SQUID inductance vs. flux Φ_S

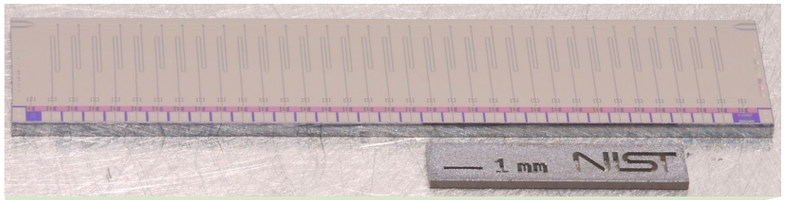






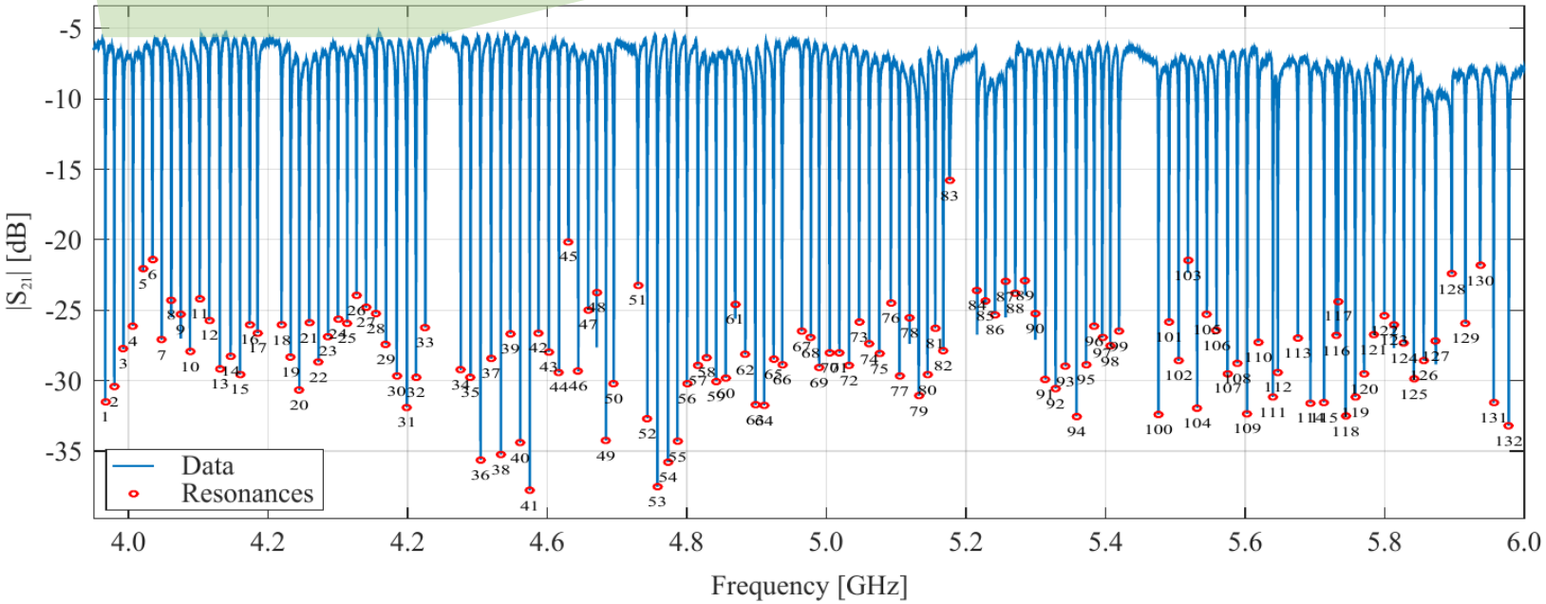
D.T. Becker et al 2019 JINST 14 P10035

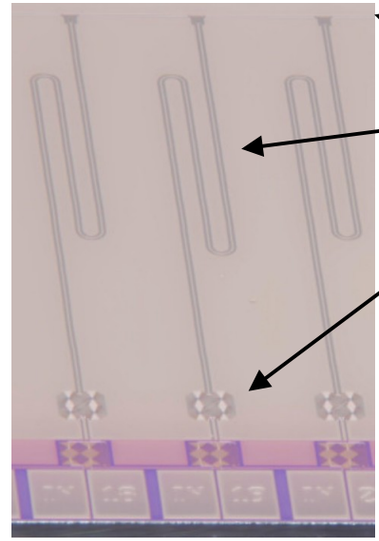
Microwave mux results: HOLMES

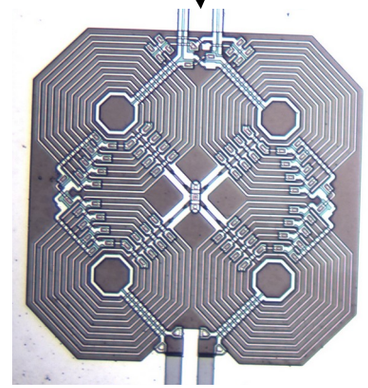


4 μmux chips in series

→ 132 resonances in 2 GHz



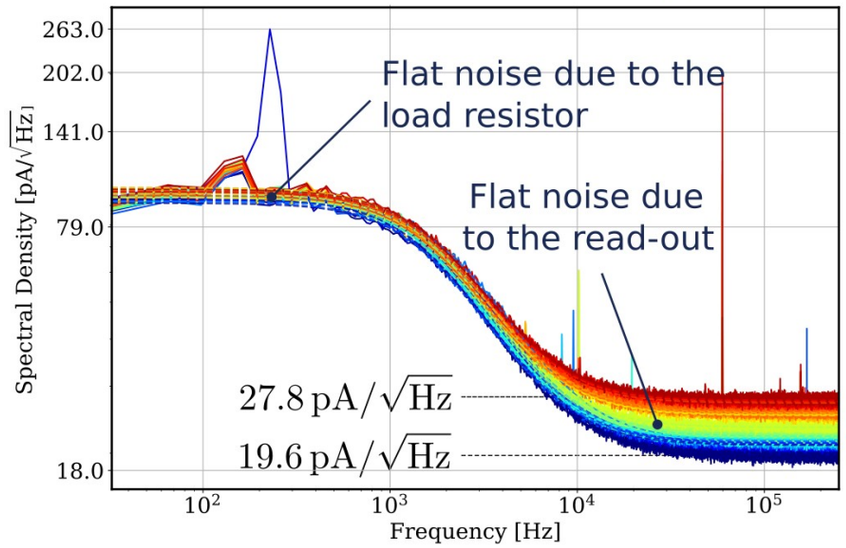




		required	measured
Resonator bandwidth	Δf_{BW} [MHz]	2	2±1
Resonator spacing	Δf [MHz]	14	14±1
Resonator depth	ΔS [dB]	>10	29±6

Read out noise

32 Channels, no TES bias applied



multiplexing factor n_{TES}

$$\frac{n_{\mathrm{TES}}}{f_{\mathrm{ADC}}} \approx \frac{\tau_{\mathrm{rise}}[\mu\mathrm{s}]}{140} \frac{1}{\mathrm{MS/s}}$$

for f_{ADC}=512 MS/s and τ_{rise}≈10 μs

→ n_{TES}≈32

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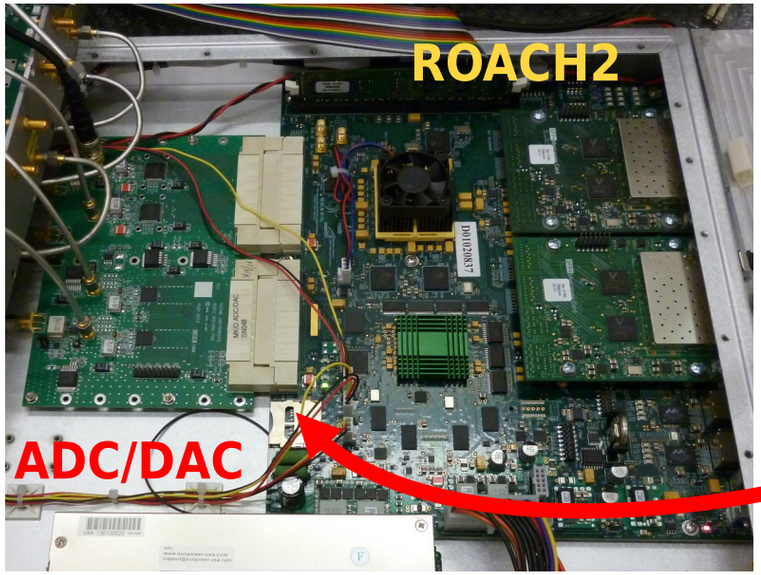
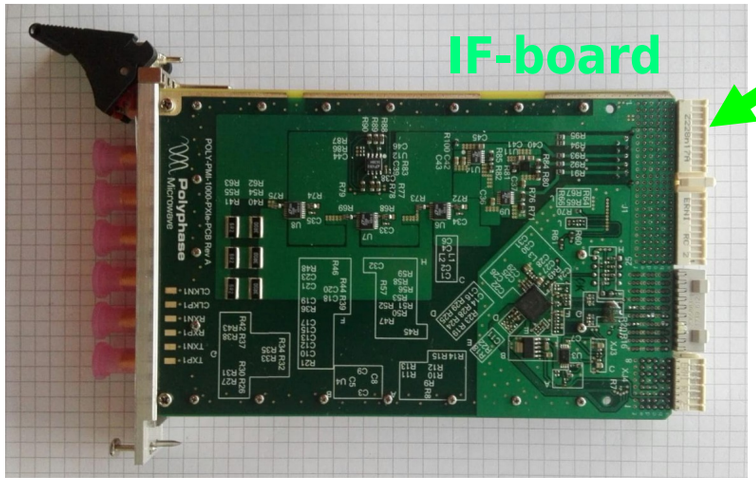
HOLMES heterodyne readout

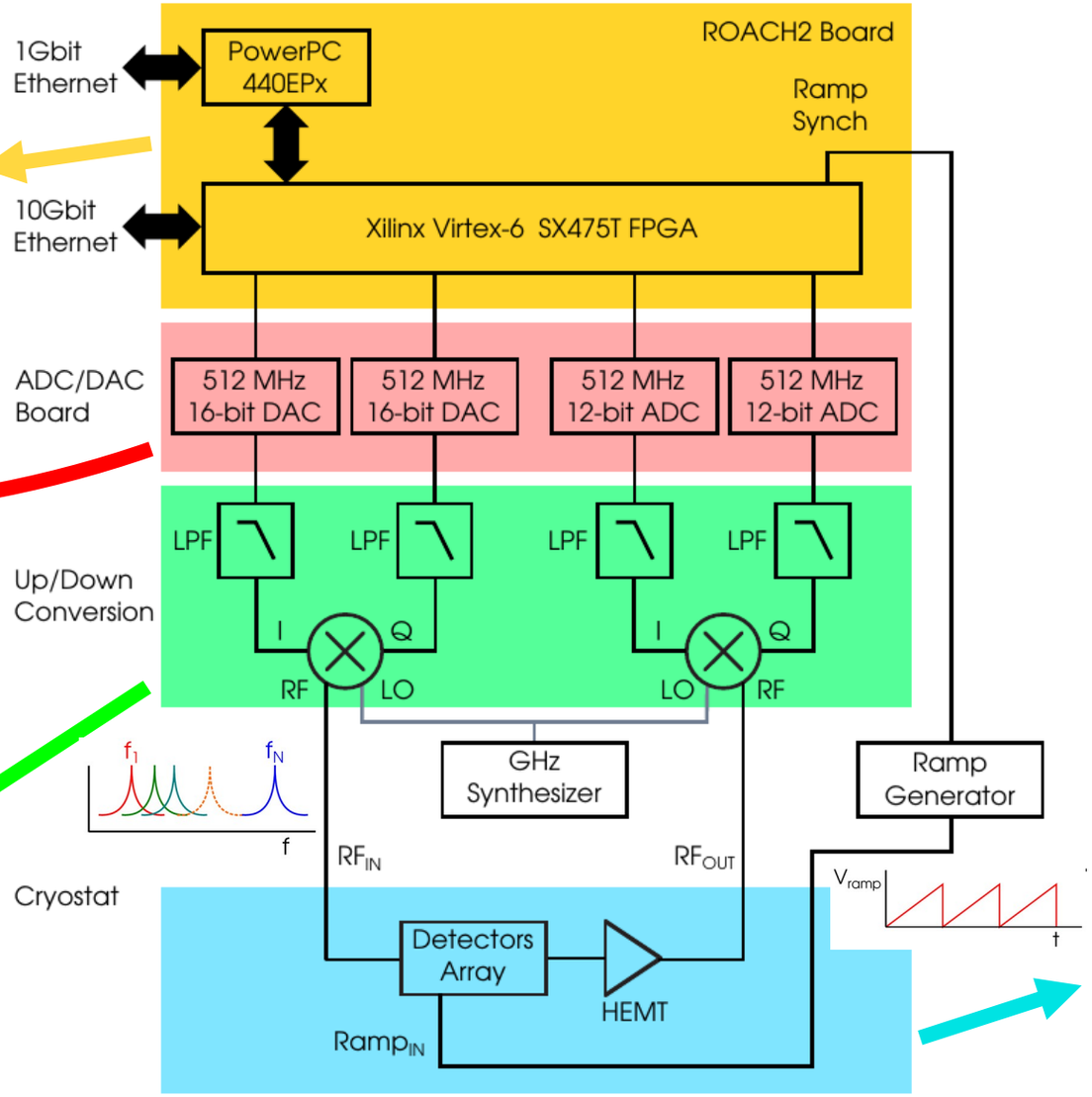
Software Defined Radio generates RF tones and demodulates output RFsignals



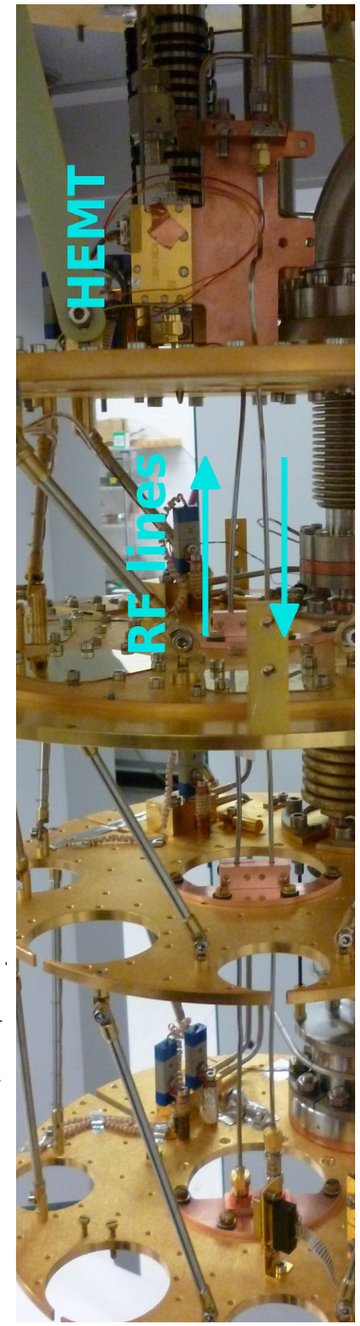
1 ROACH2 (f_{ADC}=512 MS/s)
for 32 detectors



1 IF-board for 32 detectors

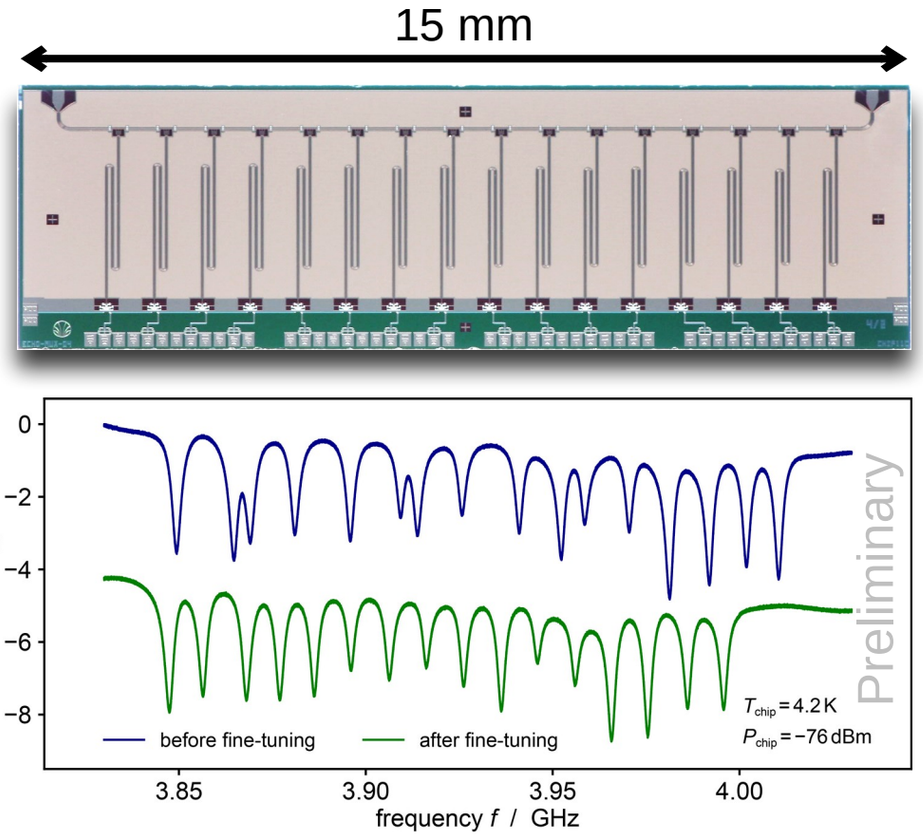


1 HEMT (BW 4-8 GHz)
for 256 detectors

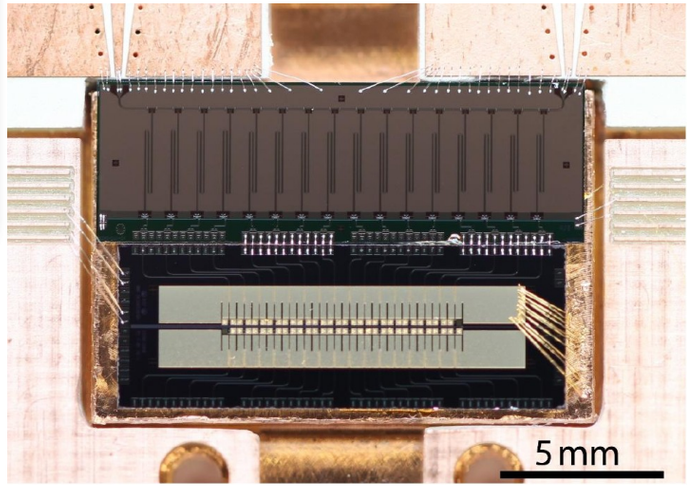


ECHo-100k

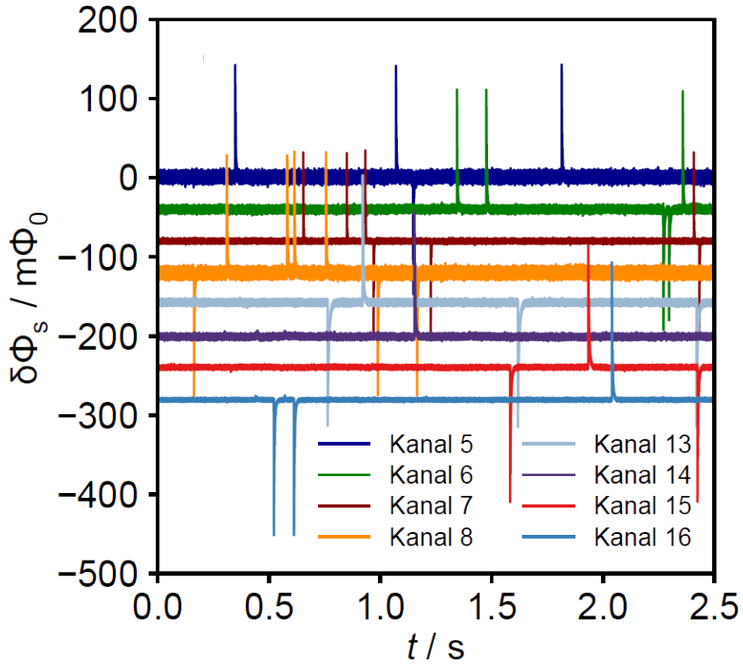
ECHo-100k will implement microwave multiplexing to readout 12000 MMCs

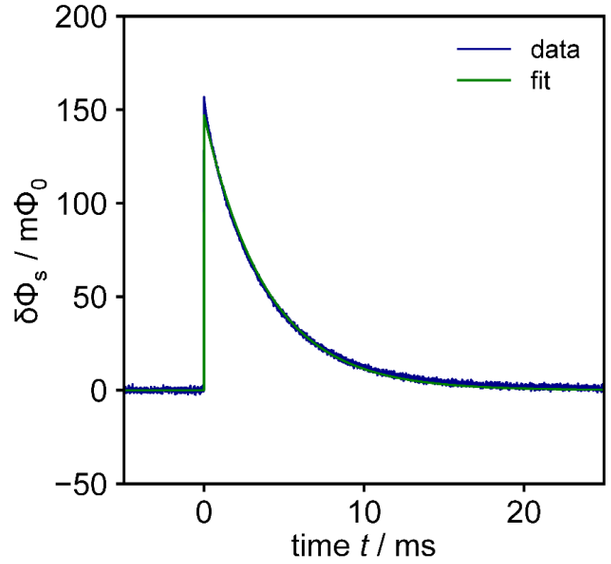


16 channel µmux chip



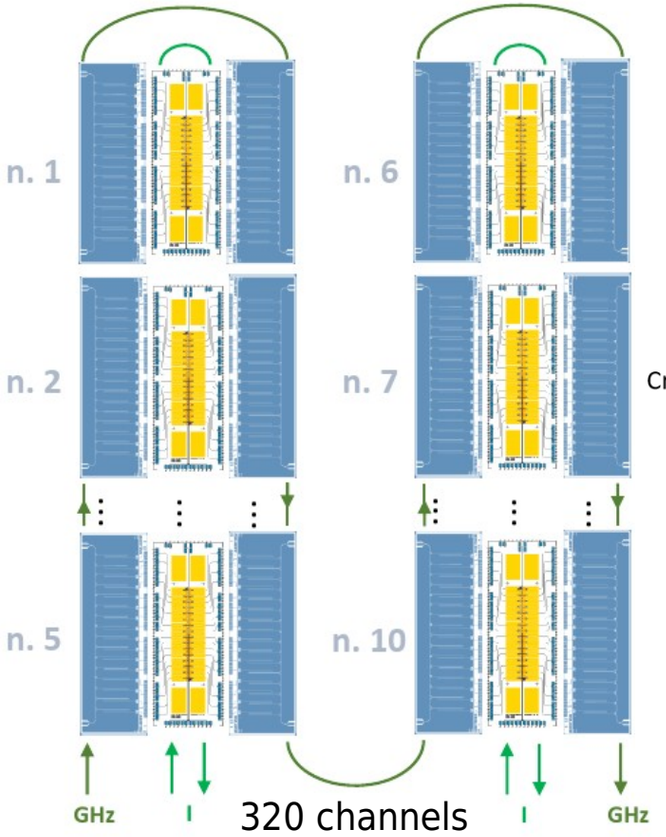
ECHo-100k
16+16 channels array
(32+32 pixels)



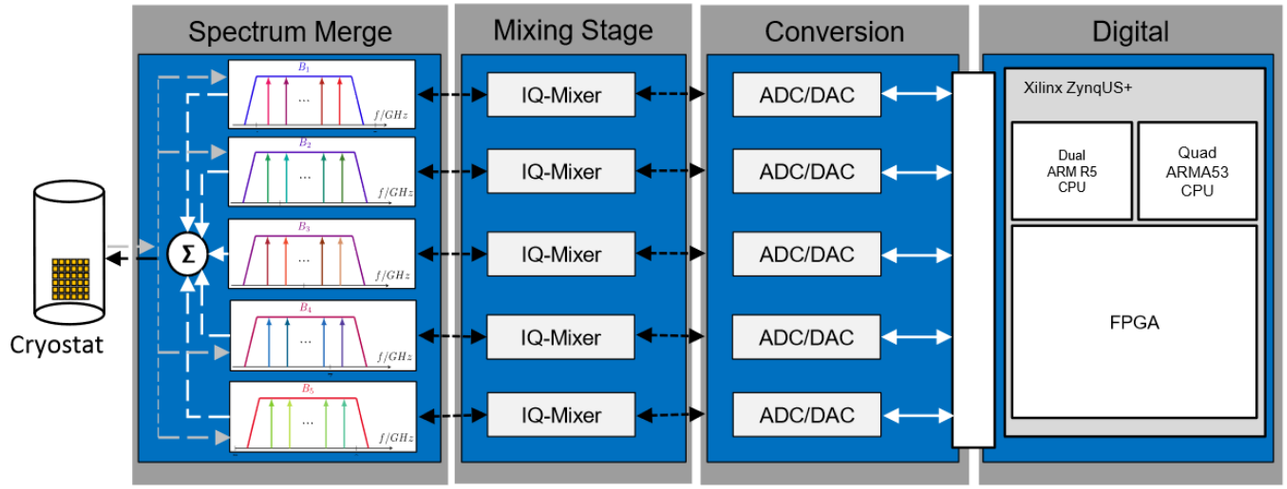


- successful read-out of 8 channels (16 pixels)
- noise limited by low available RF power
  → $\Delta E_{FWHM} \approx$ 10-30 eV (low power) depending on Q_l
- with high power read-out of 2 channels
  → $\Delta E_{FWHM} \approx$ 9 eV

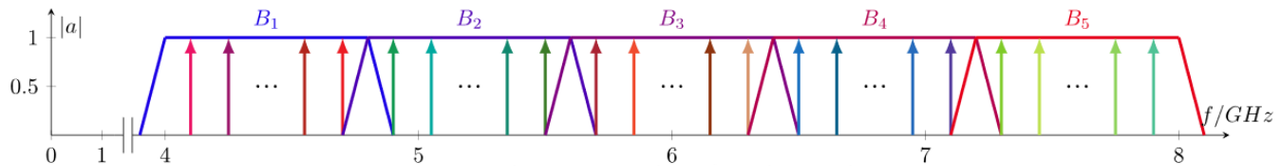
ECHO-100k: SDR for heterodyne readout development



O. Sander *et al.*, *IEEE Transactions on Nuclear Science* 66.7 (2019)



Software-defined Radio (SDR) system architecture



Spectrum of one single transmitter/receiver-pair

ECHo-100k

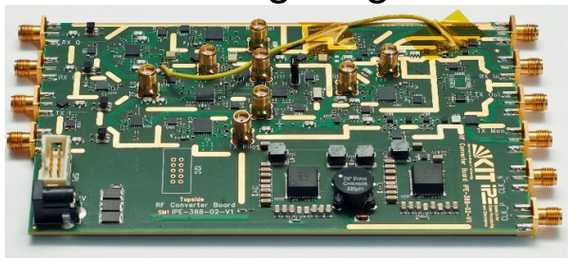
12000 MMC pixels =

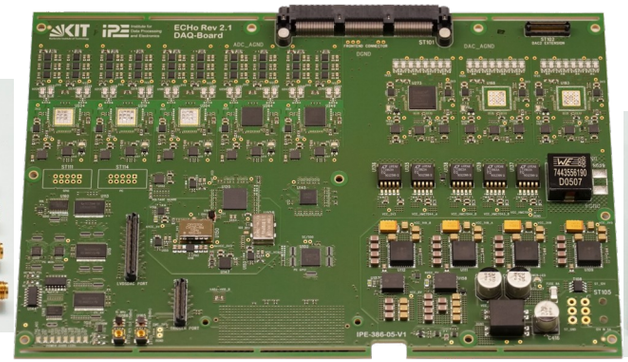
400 MMC channels ×

15 SDR systems / HEMT

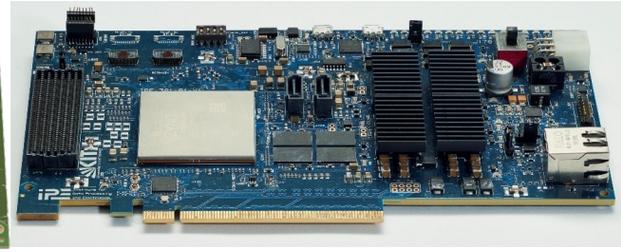
1GS/s ADC / 500MS/s DAC Board

Mixing stage





FPGA-Board



HEMT 4-8 GHz band covered by 5×800 MHz bands → 80 channel/band → 400 total channels

Data analysis

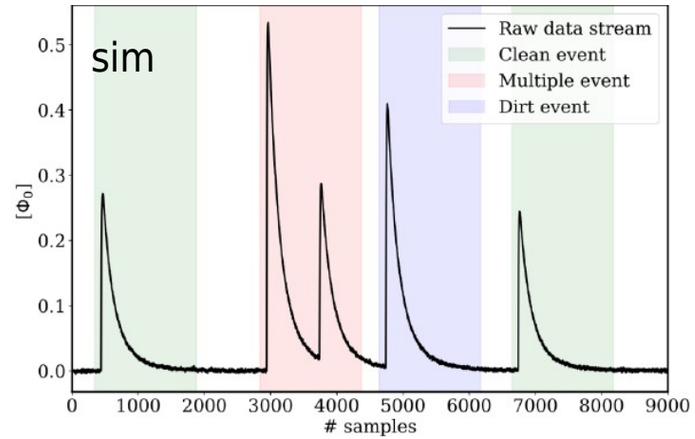
First level data reduction

Evaluation of pulse information

Energy estimation

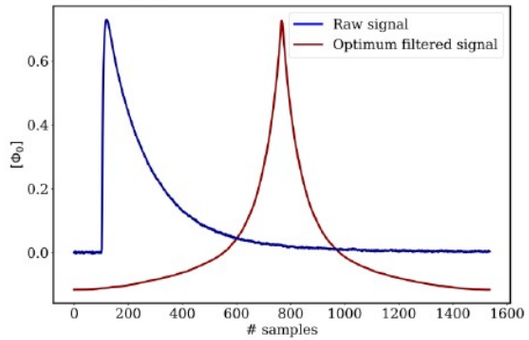
- Optimum filter
- Arrival time correction
- Gain drift correction
- Energy calibration

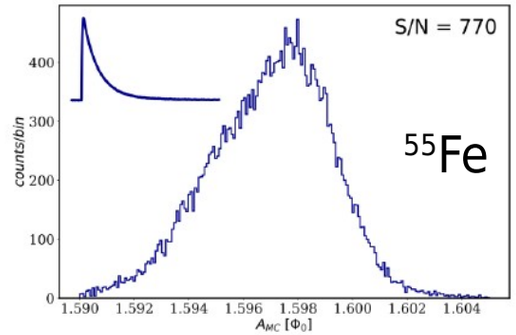
Second level data reduction

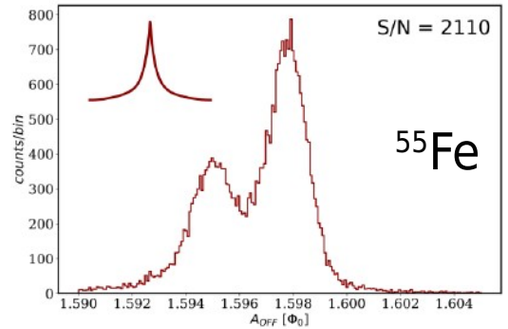


- classify events
- discard bad ones
- estimate time contants, arrival time, pretrigger level (baseline), and other parameters

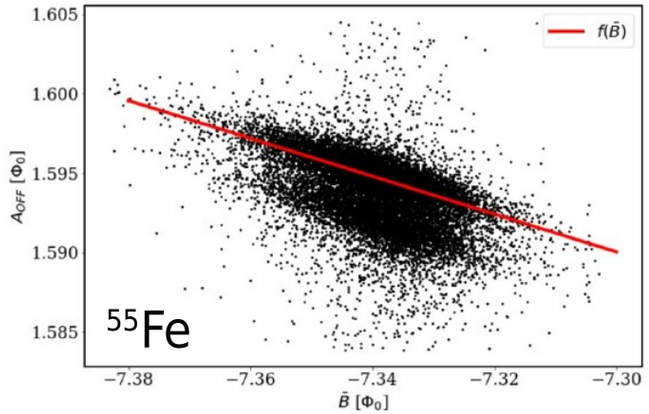
energy estimator based on optimal filter in various implementations





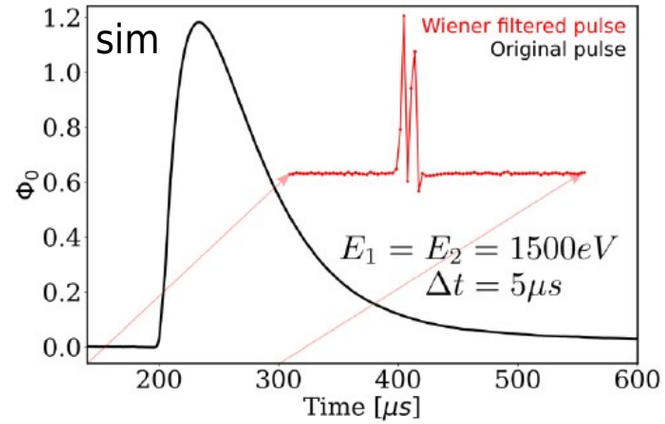


gain drift correction



amplitude vs. baseline ($\propto T_{op}$) correlation

rise time pile-up detection



M. Borghesi, PhD, U. Milano-Bicocca, 2022

Hammann, R. et al. EPJ C 81, 963 (2021)

Pile-up discrimination

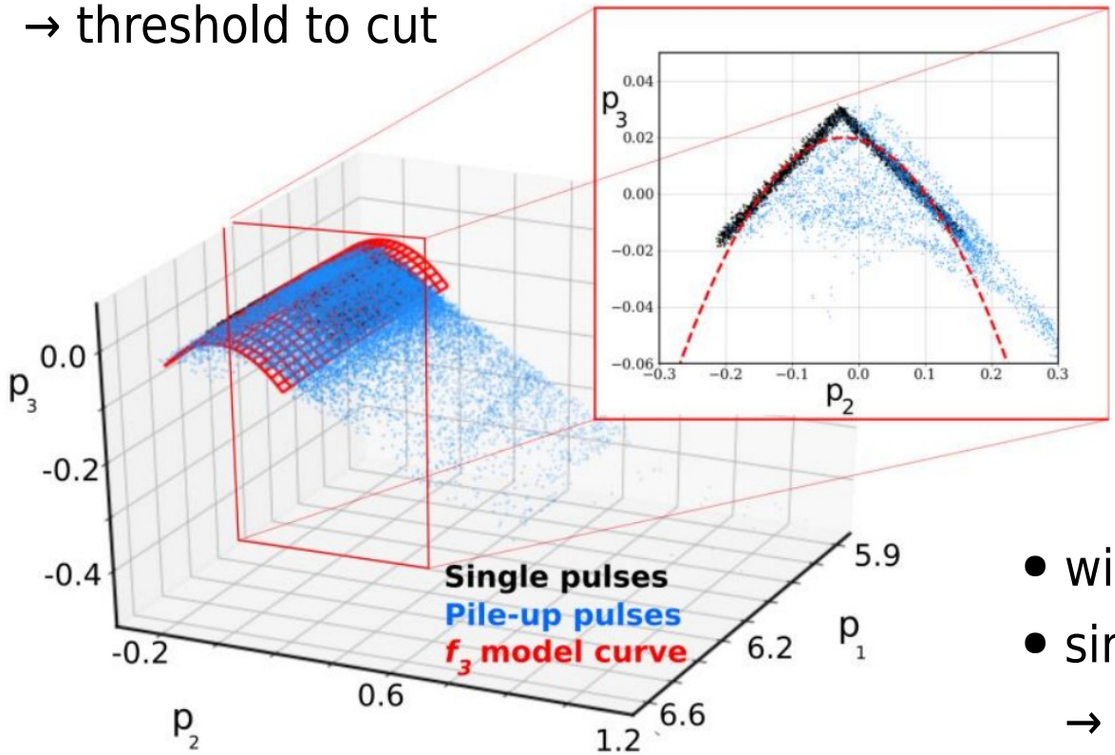
- advanced discrimination technique to identify pile-up events
  - Discrimination through Singular Vector Projections (DSVP)
  - “unsupervised learning” technique, based on singular value decomposition, PCA and multiple linear regression

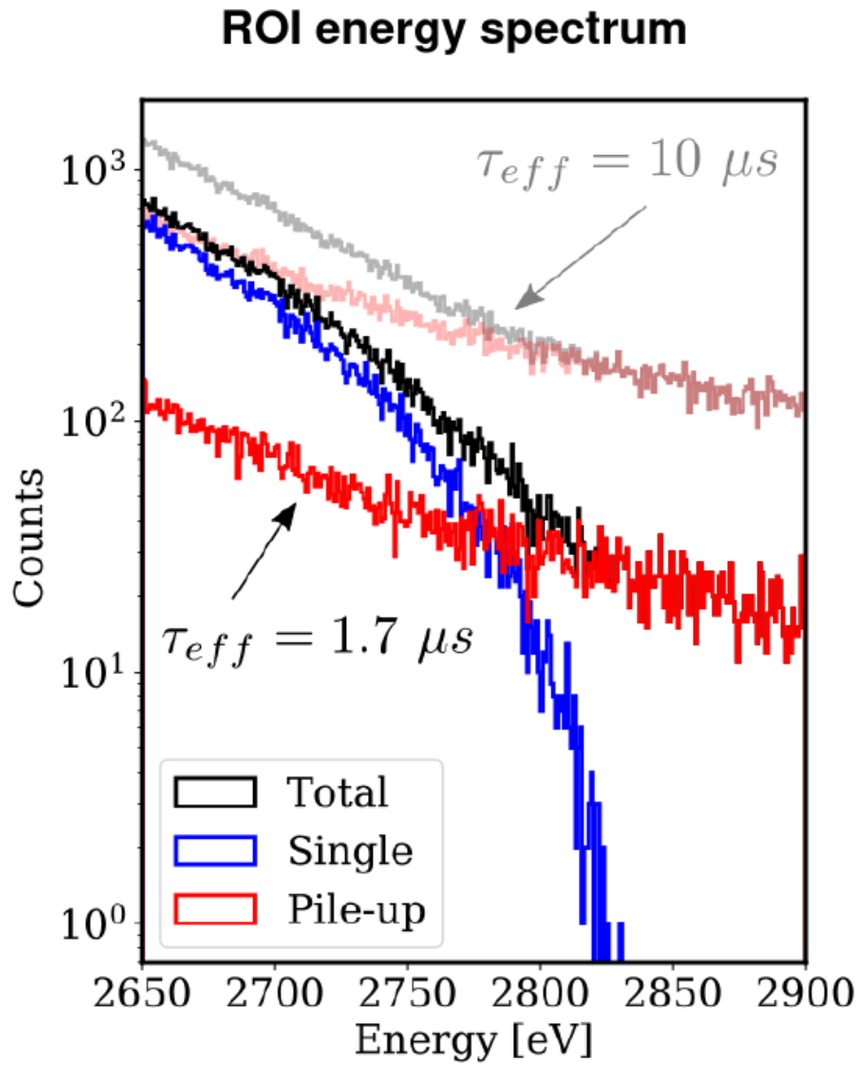
→ SVP decomposition

→ raw dataset cleaning (PCA)

→ model point distribution in singular vector projection space

→ threshold to cut





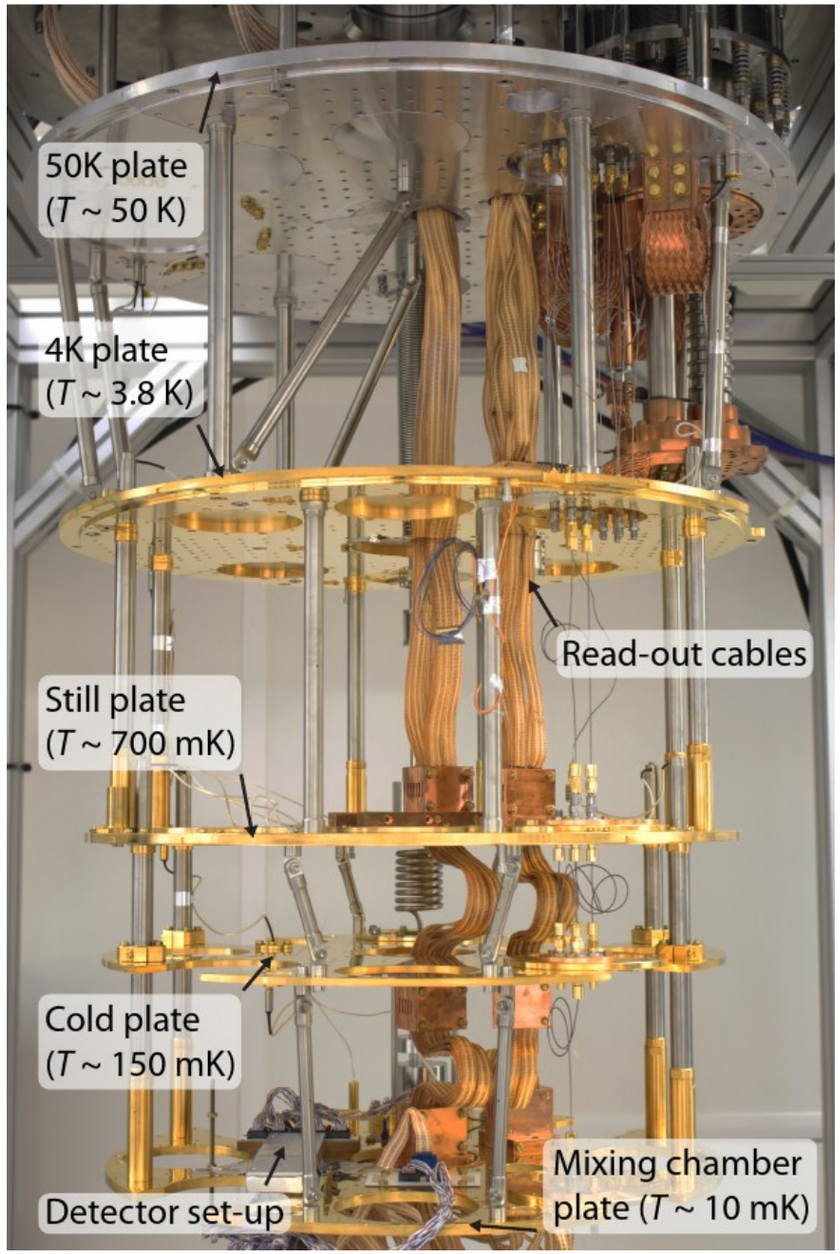
- with DSVP τ_R is limited by the pulse sampling time $1/f_{ramp}$
- simulations for HOLMES pulses

  → $\tau_R \approx$ 1-2 μs for $\tau_{rise} \approx$ 10-20 μs and f_{ramp}=500 kHz

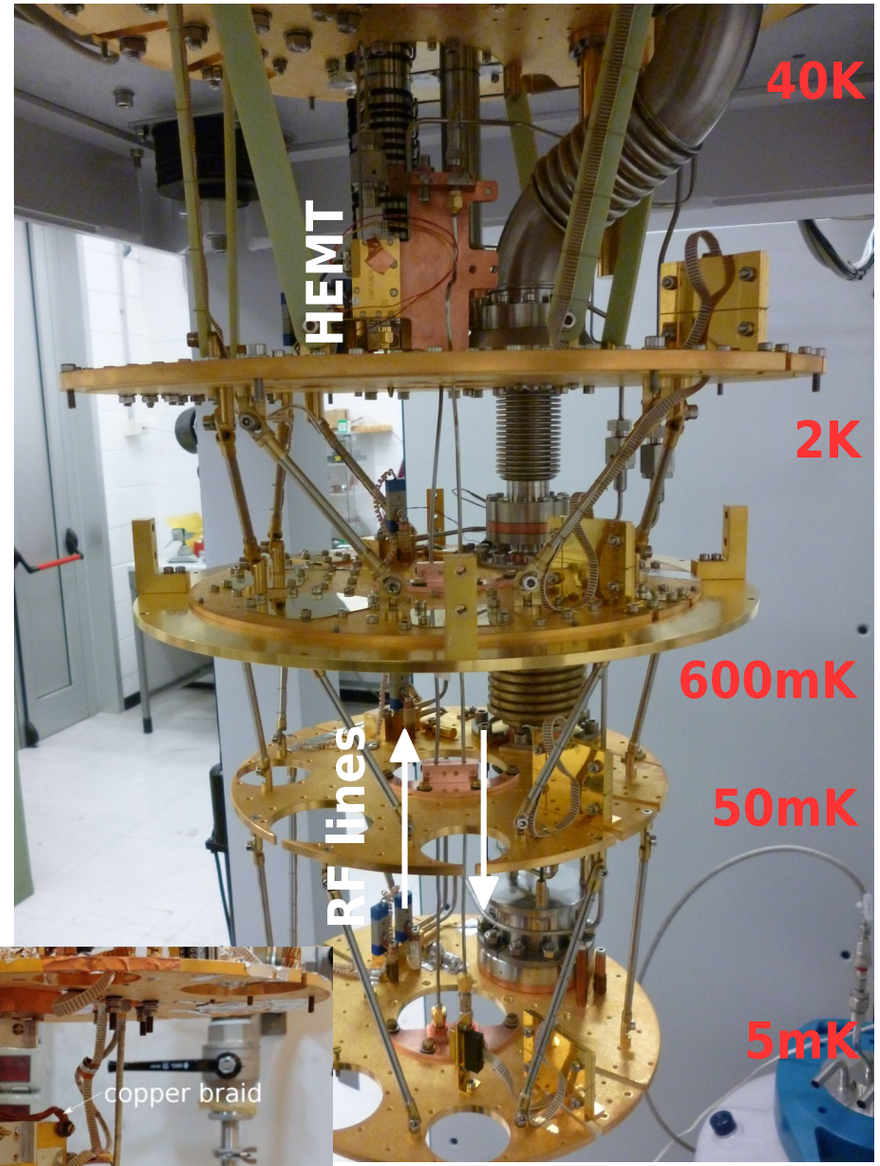
Borghesi, M. et al. EPJ C 81, 385 (2021)

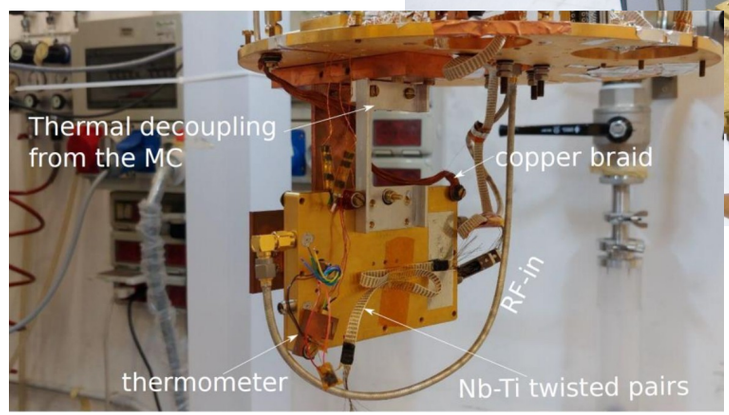
Cryogenic set-ups

ECHo-1k



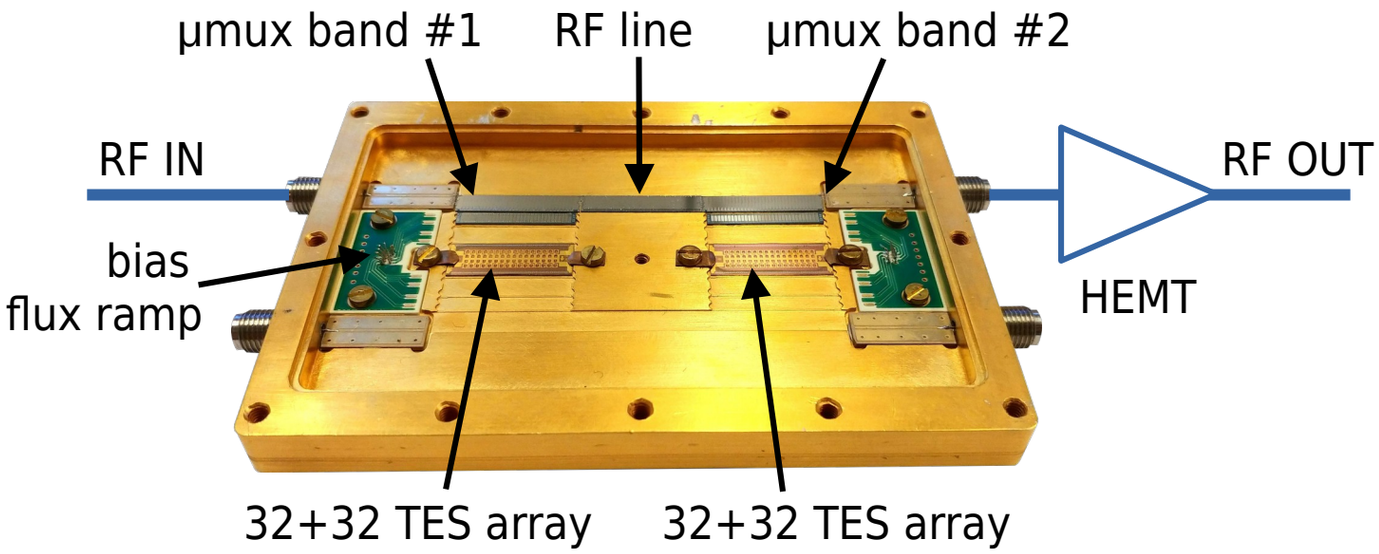
HOLMES (for 256 pixels)





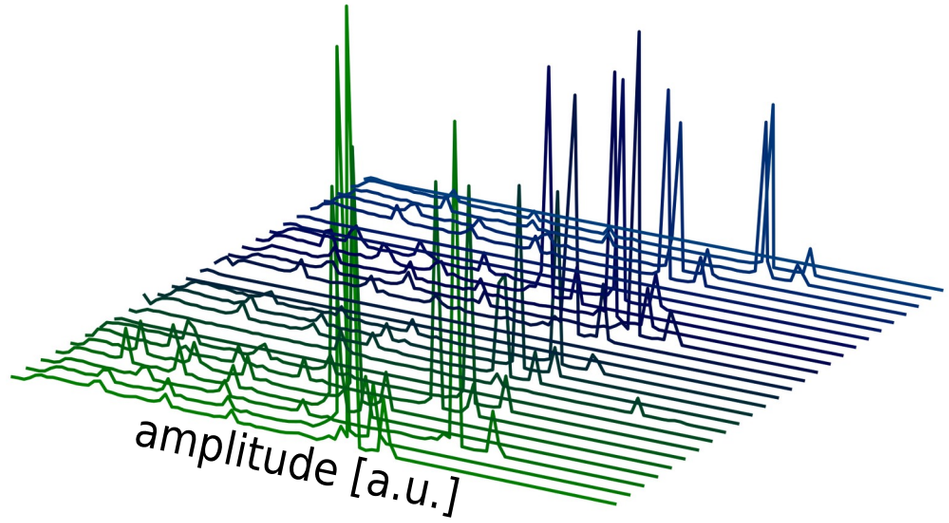
HOLMES status: detectors, readout, analysis

64 multiplexed detectors



raw data from 26 multiplexed detectors

^{55}Fe X-rays + Al, Cl, Ca X-ray fluorescence

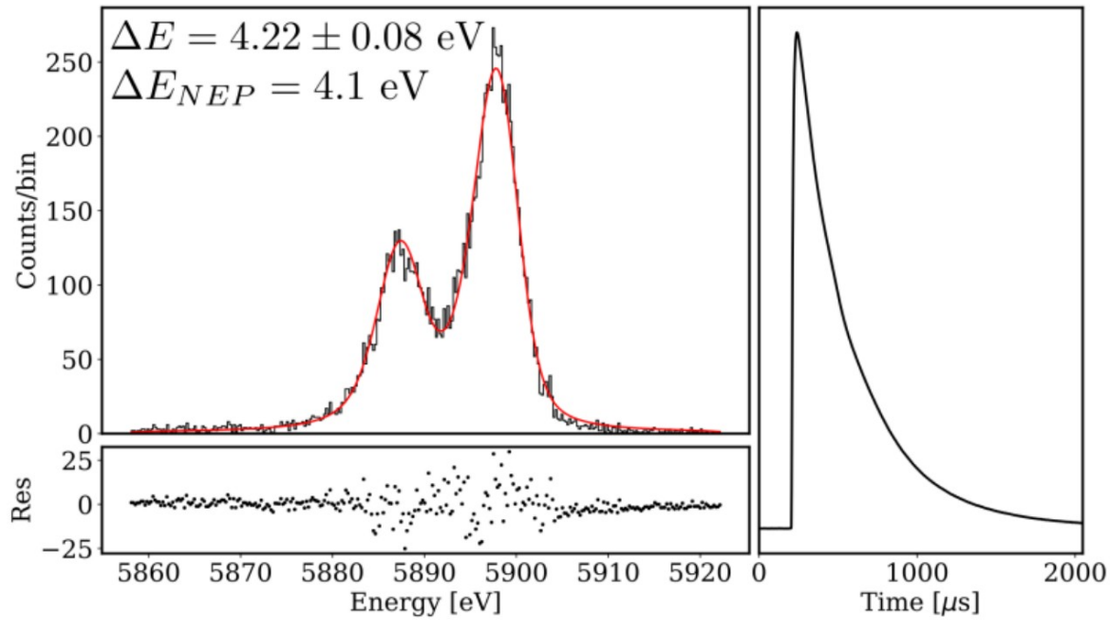


- fully processed TES arrays without ^{163}Ho implant
- set-up for 126 multiplexed pixels
- 2 μmux chips but only 32 bonded pixels
- 16 channel firmware version (stable) for ROACH2
- at 5.9 keV:

→ $\Delta E_{FWHM} \approx 4\text{-}6$ eV

→ $\tau_{rise} \approx 15$ μs (R/L limited to match DAQ) → $\tau_R \approx 1.5$ μs

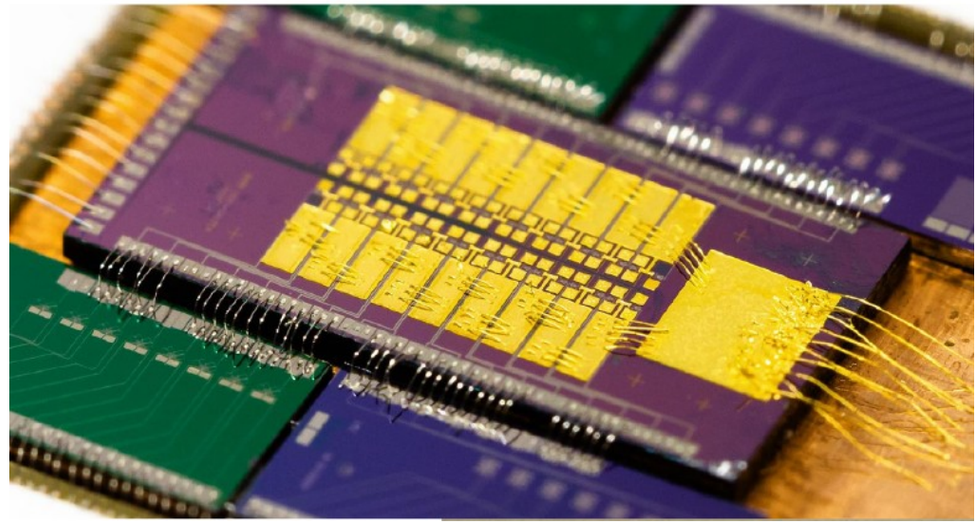
→ $\tau_{decay} \approx 300$ μs

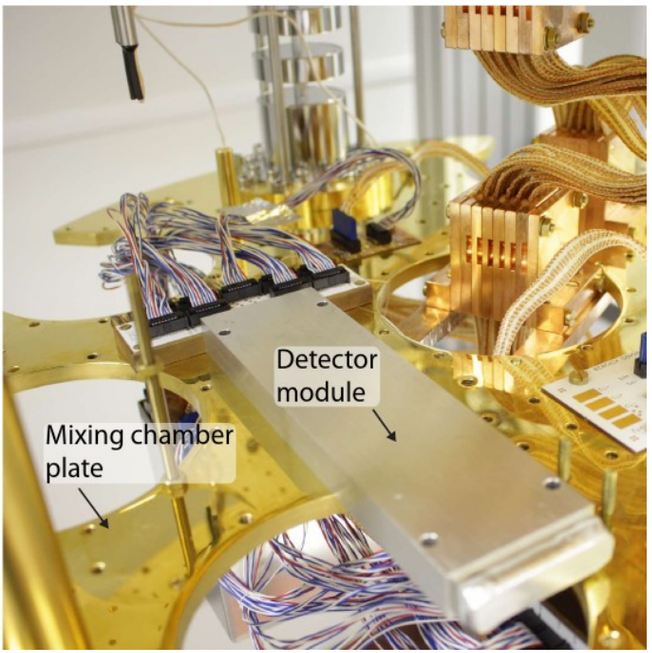


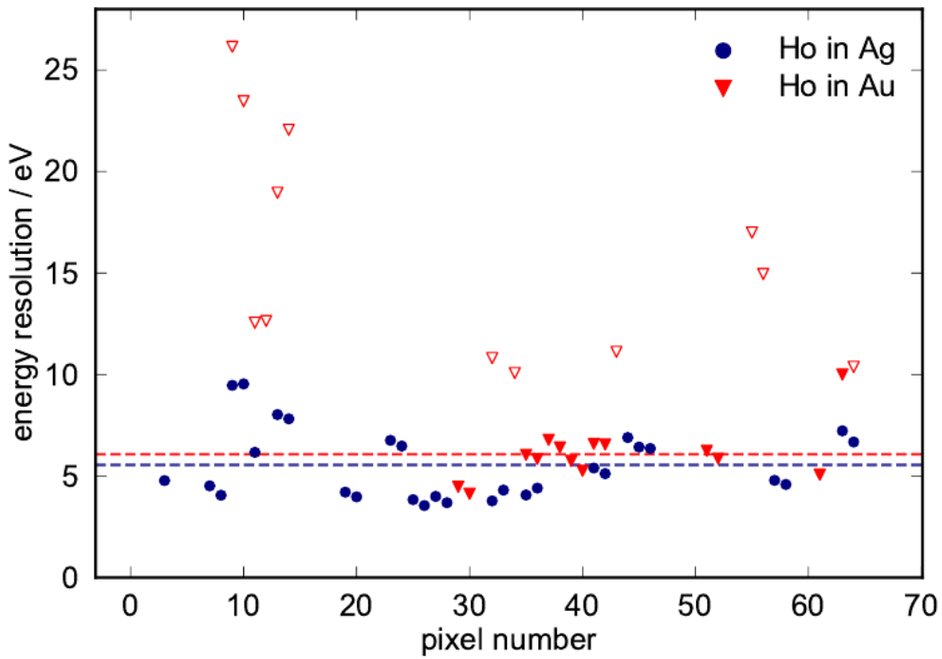
ECHo-1k status: detectors

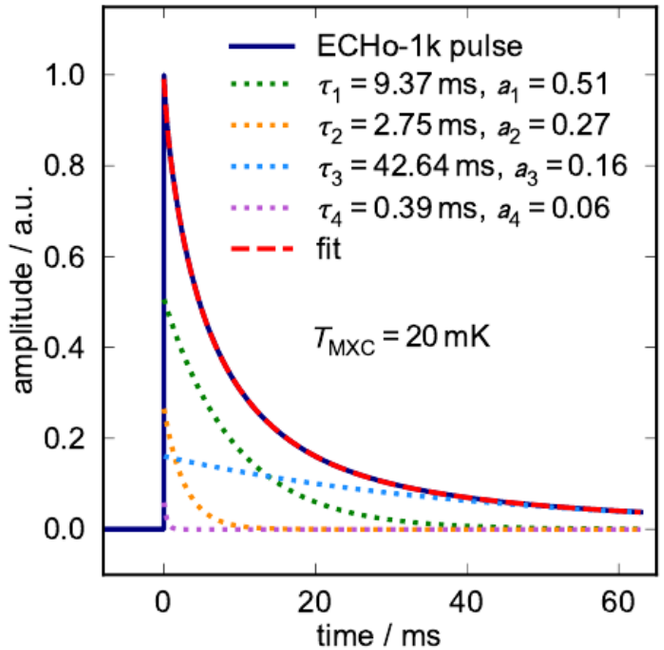
2 detector modules with ^{163}Ho in Au and Ag host material
parallel dc-SQUID readout

host	^{163}Ho pixels	bkg pixels	⟨A⟩ [Bq]	A_{tot} [Bq]
Au	23	3	0.94	28.1
Ag	34	6	0.71	25.9







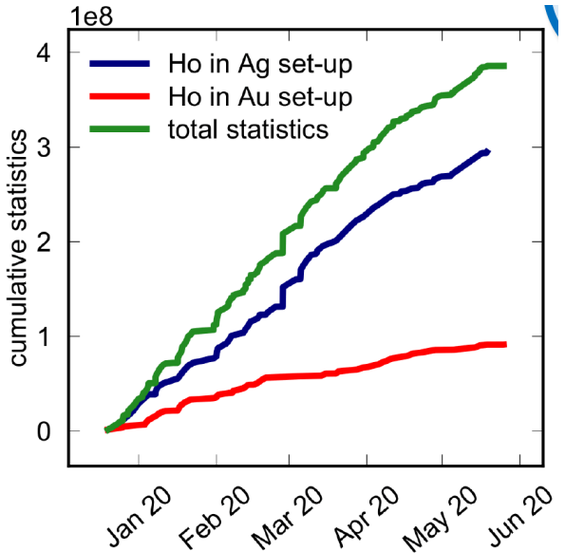


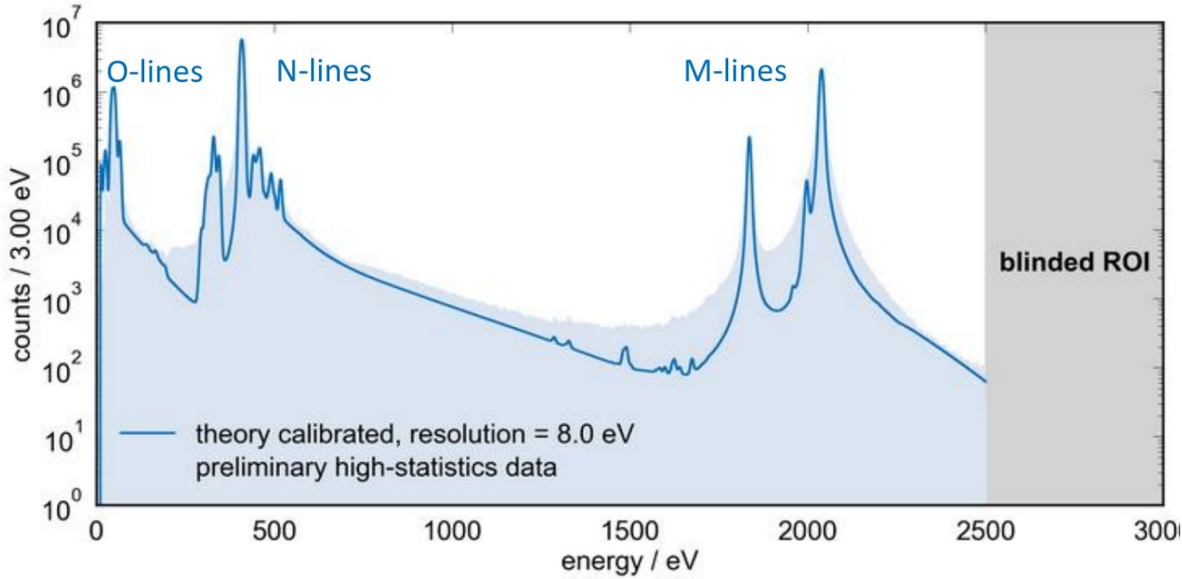
- $\tau_{rise} \approx 1$ µs limited by SQUID bandwidth
- complex decays time: mostly $\tau_1 \approx 10$ ms

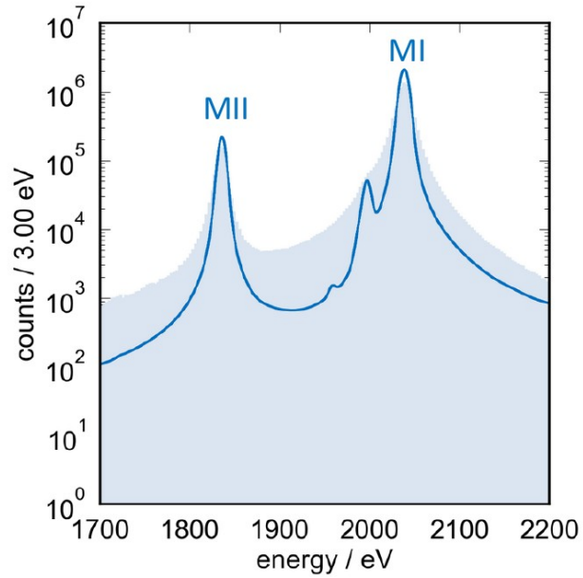
F. Mantegazzini et al, NIM A 1030 (2022) 166406

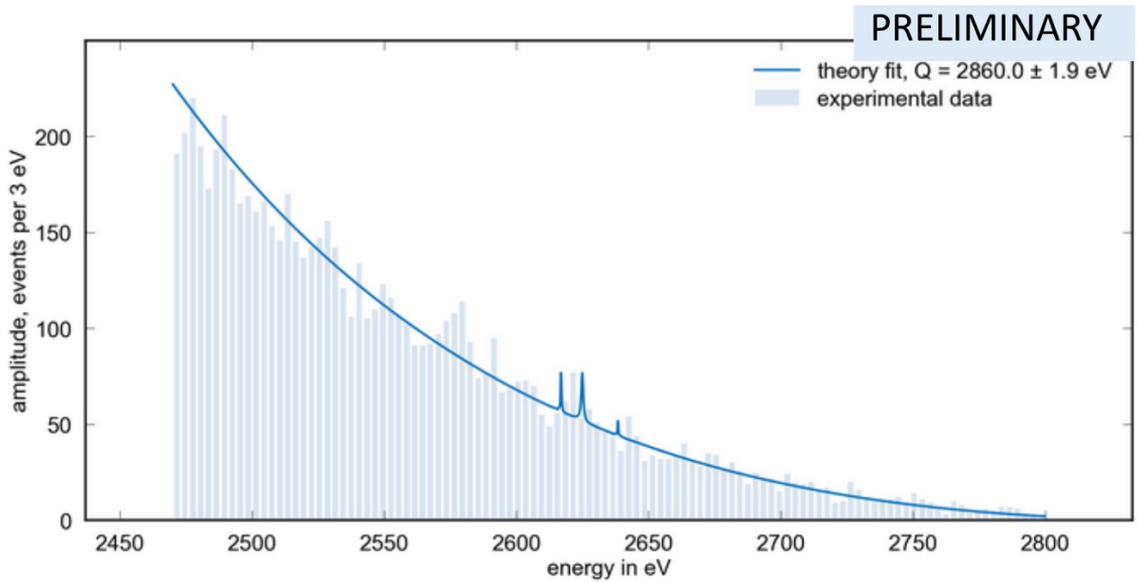
ECHo-1k status: 10^7 events spectrum

6×10^7 events acquired with detectors having ^{163}Ho in Ag host material







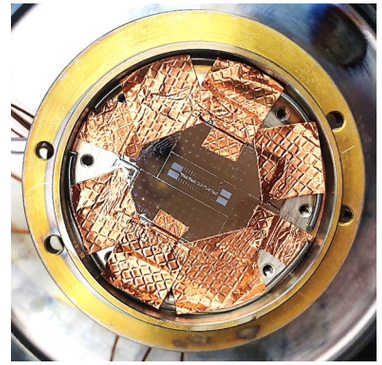


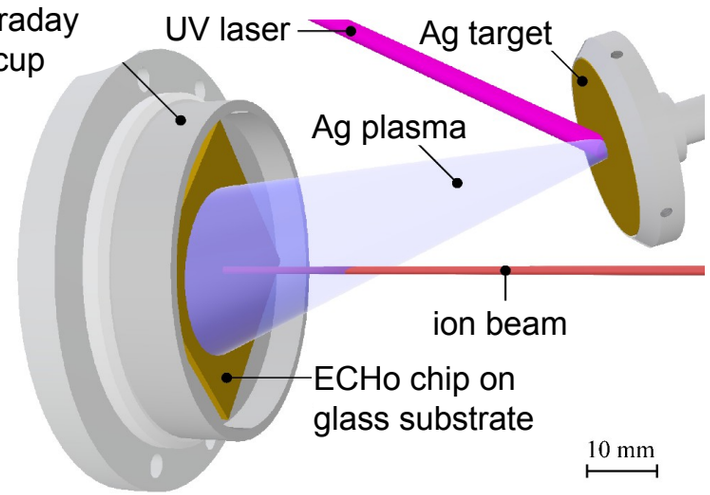
- end-point analysis in progress
- from preliminary end-point analysis
  - EC Brass & Haverkort theory
  - Flat background

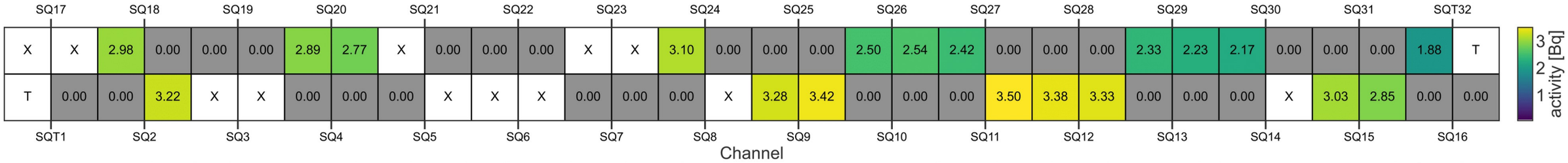
  → Q_{EC}=(2860 ± 2stat ± 5syst) eV

ECHo-100k progresses

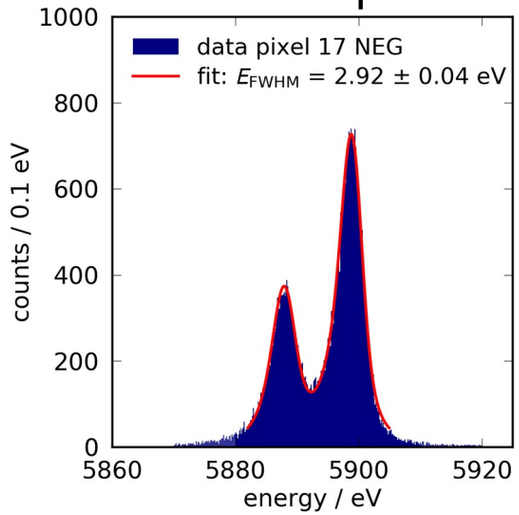
- 30 MBq of pure ^{163}Ho available
- implant system improvements for 10 Bq
  - beam positioning on ECHo-100k wafer
  - co-deposition through PLD in progress
- pixel design optimization to minimize *C*
- μmux read-out system development in progress

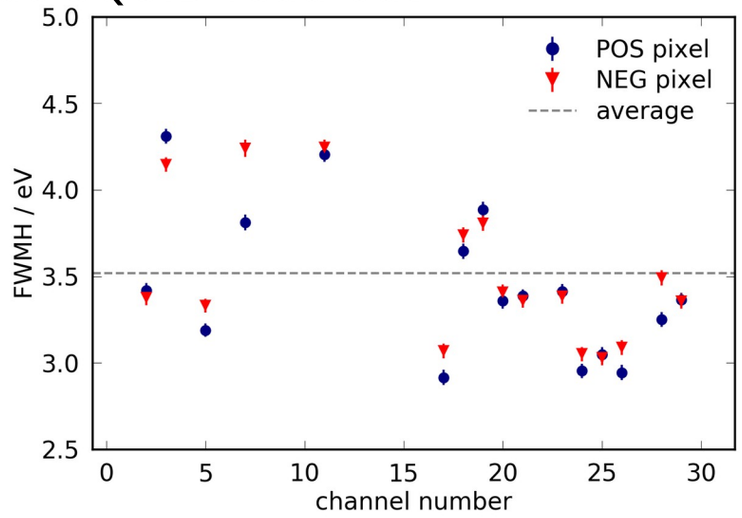




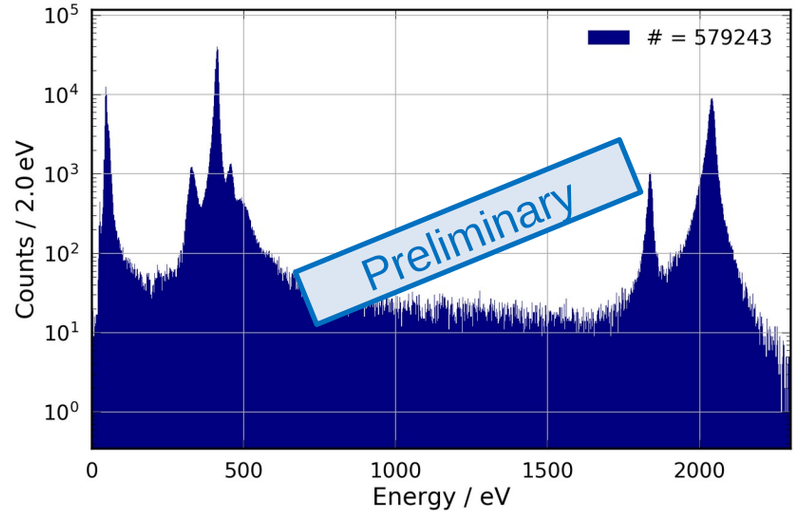


pixel characterization before ion implant
parallel dc-SQUID read-out



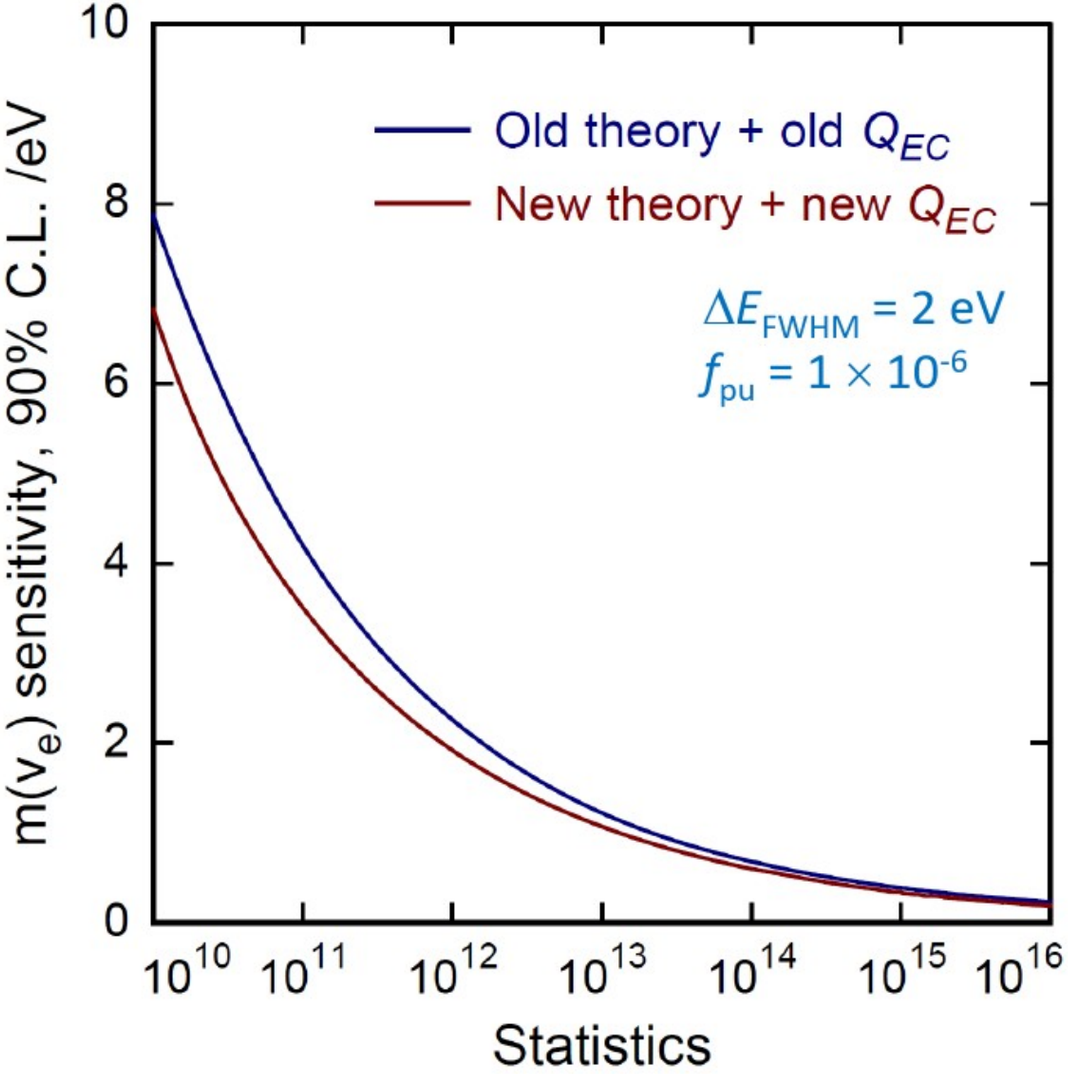


after ion implant without second Au layer
dc-SQUID read-out → ΔE≈3.4 eV at N1 line

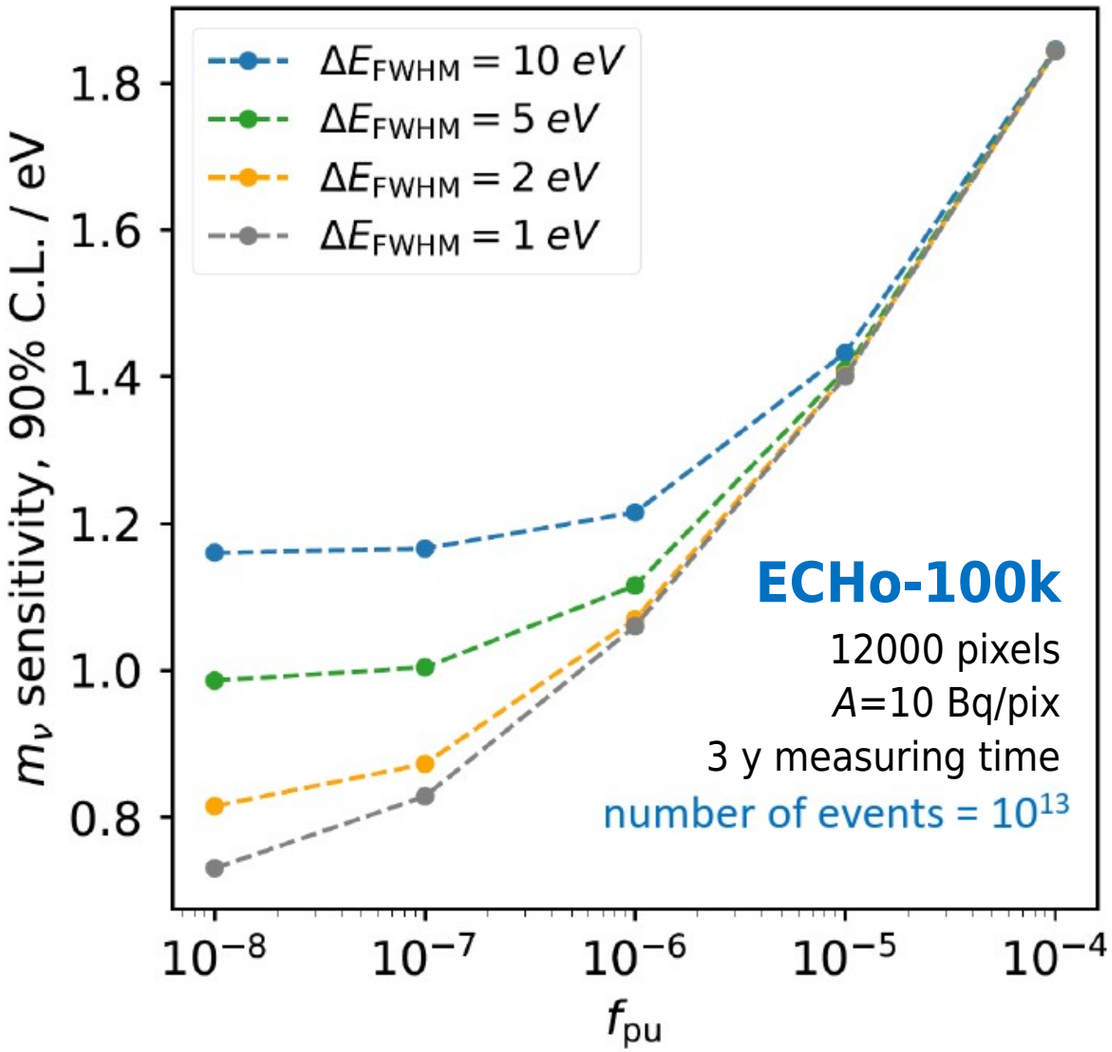


ECHo-100k updated expected sensitivity

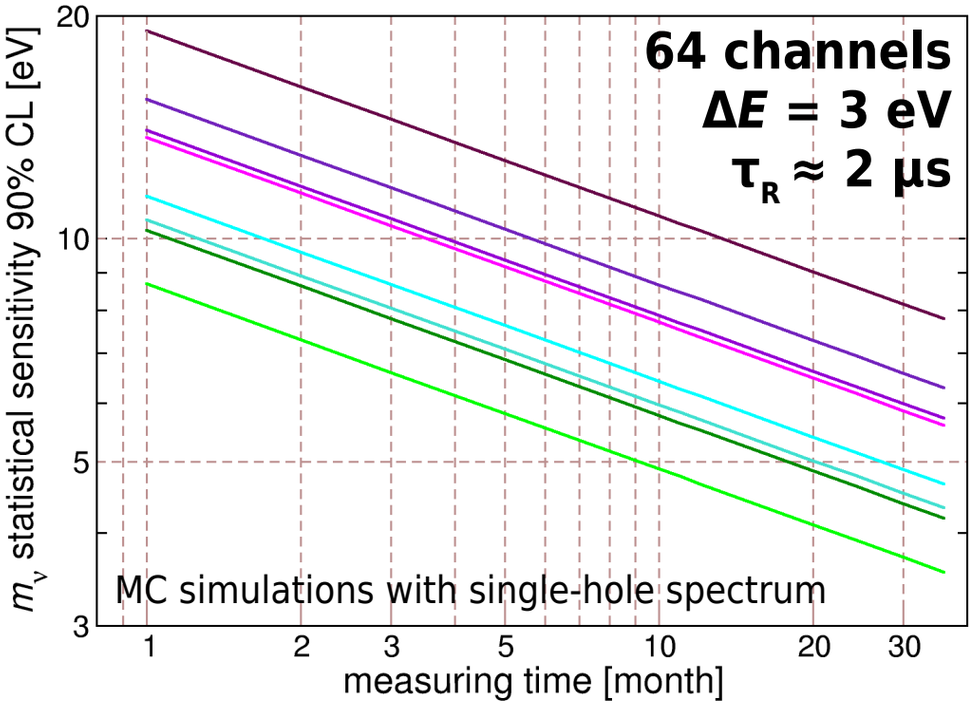
Brass & Haverkort theoretical model + new Q_{EC}-value

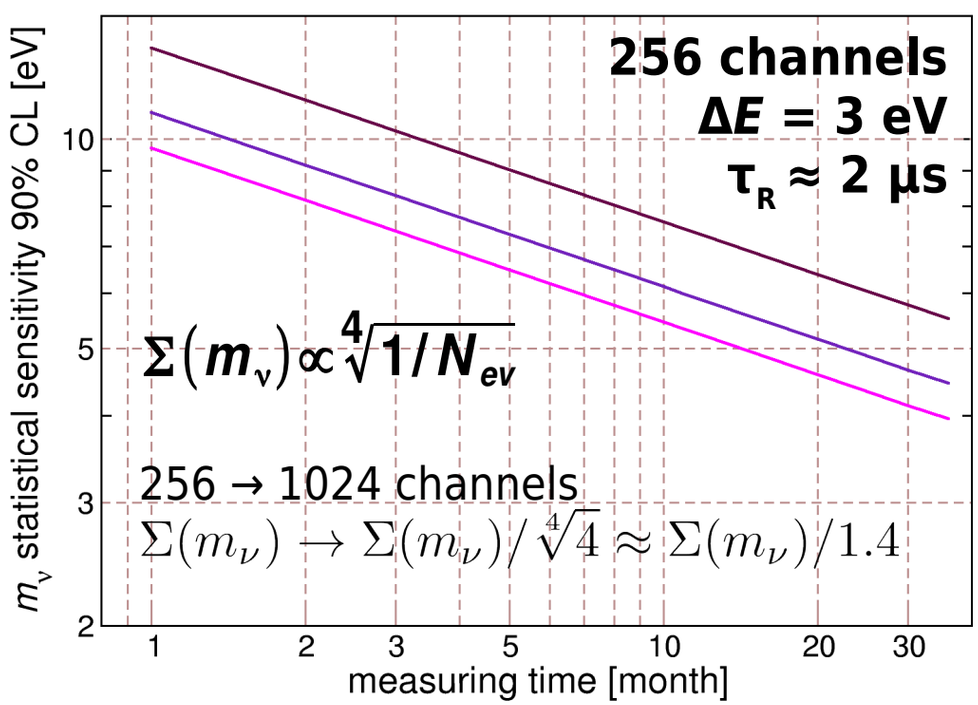


Sensitivity for the coming phase of ECHo



HOLMES sensitivity evolution vs. pixel activity



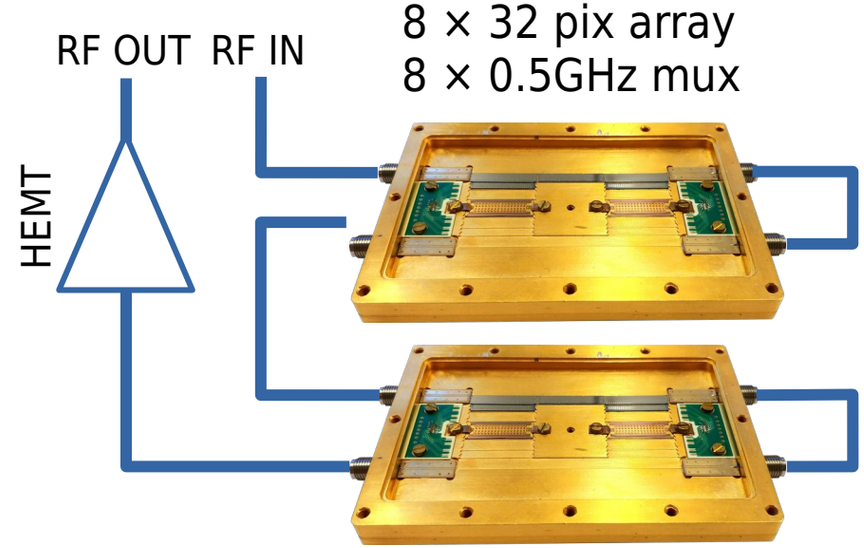


A = 1 Hz/det
A = 3 Hz/det
low dose

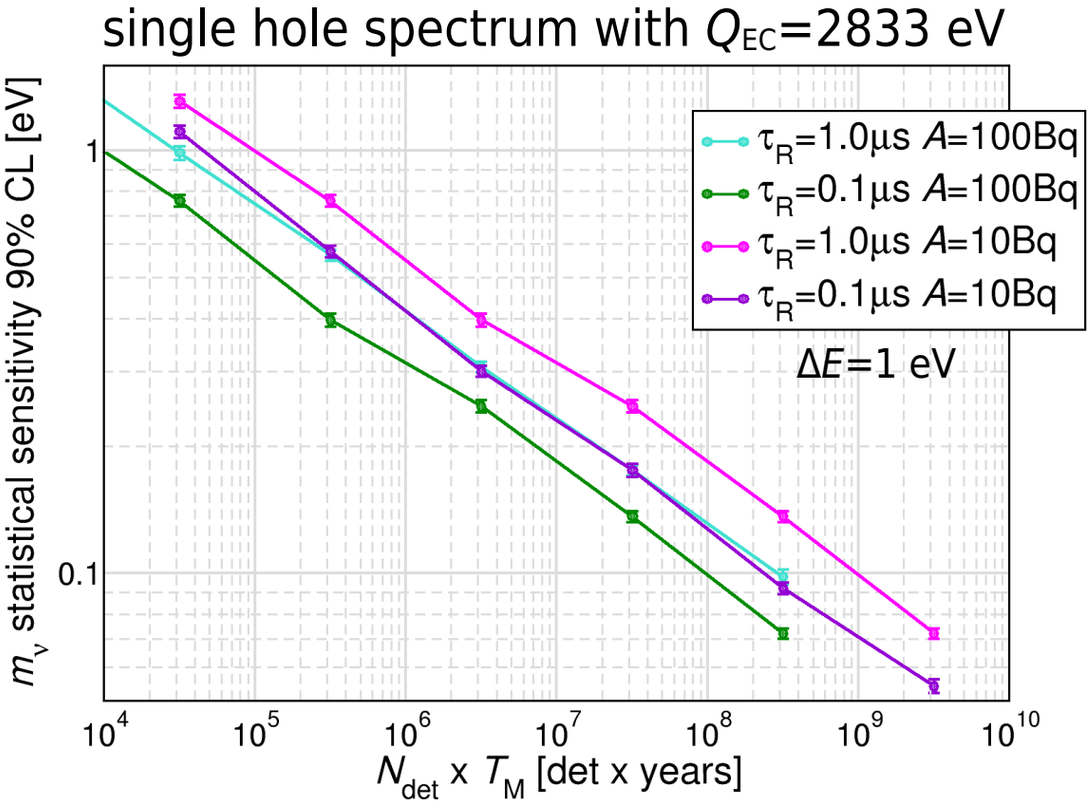
A = 5 Hz/det
A = 10 Hz/det
A = 30 Hz/det
A = 50 Hz/det
A = 100 Hz/det
A = 300 Hz/det
with focusing

?

8 × ROACH2+IF boards



Beyond ECHo and HOLMES: a 0.1eV experiment



10 years measuring time

***A*/det [Bq]**	10	10	100	100	1000	1000
τ_R [us]	0.1	1	0.1	1	0.1	1
f_{pp}	1.0E-06	1.0E-05	1.0E-05	1.0E-04	1.0E-04	1.0E-03
N_{det}	3.0E+07	9.8E+07	9.8E+06	3.0E+07	3.0E+06	1.3E+07
***A* total [Bq]**	3.0E+08	9.8E+08	9.8E+08	3.0E+09	3.0E+09	1.3E+10
^{162}Er [mg] *	2274	7429	7429	22742	22742	98548

* ^{162}Er/A(^{163}Ho) = 3790 mg/GBq + 50% usage efficiency

- pixel activity ≥100Bq/det ↔ ^{163}Ho heat capacity
- time resolution below 0.1μs ↔ multiplexing and DAQ bandwidth
- about 10M pixels ↔ multiplexing and DAQ bandwidth

Conclusions

- **ECHo and HOLMES are reaching a statistical sensitivity of order of 1 eV in few years**
  - many technical challenges faced successfully (also separately by ECHo and HOLMES)
    - production of large amounts of clean 163Ho samples
    - efficient ion implantation
    - high resolution detectors with multiplexed read-out
    - sophisticated analysis tools
  - some activities are still required to fully assess the potential of holmium experiments
    - understanding the holmium decay spectrum
    - effect of high activities on detector performances
    - investigating systematic effects
- **longer term plans: next generation experiments for sub-eV sensitivities require**
  - larger international collaboration
  - increased single pixel activity
  - cost reduction (isotope production and efficient usage, readout electronics)

Collaborations

HOLMES

Università di Milano-Bicocca, Italy
INFN Milano-Bicocca, Italy
INFN Genova, Italy
INFN Roma, Italy
INFN LNGS, Italy
NIST, Boulder, USA
PSI, Villigen, Switzerland
ILL, Grenoble, France

ECHo

Heidelberg University, Germany
Johannes Gutenberg University, Mainz, Germany
GSI. Darmstadt, Germany
Helmholtz Institute Mainz, Mainz, Germany
University of Tübingen, Tübingen, Germany
MPI für Kernphysik„ Heidelberg, Germany
KIT, Karlsruhe, Germany
Petersburg Nuclear Physics Institute, Gatchina, Russia
Laboratoire Souterrain de Modane, Modane, France
CERN, Geneva, Switzerland
ILL, Grenoble, France