# Sterile Neutrino Searches at J-PARC MLF

Takasumi Maruyama (KEK)

Thanks to J.Spitz (U of Michigan) for KPIPE experiment discussion

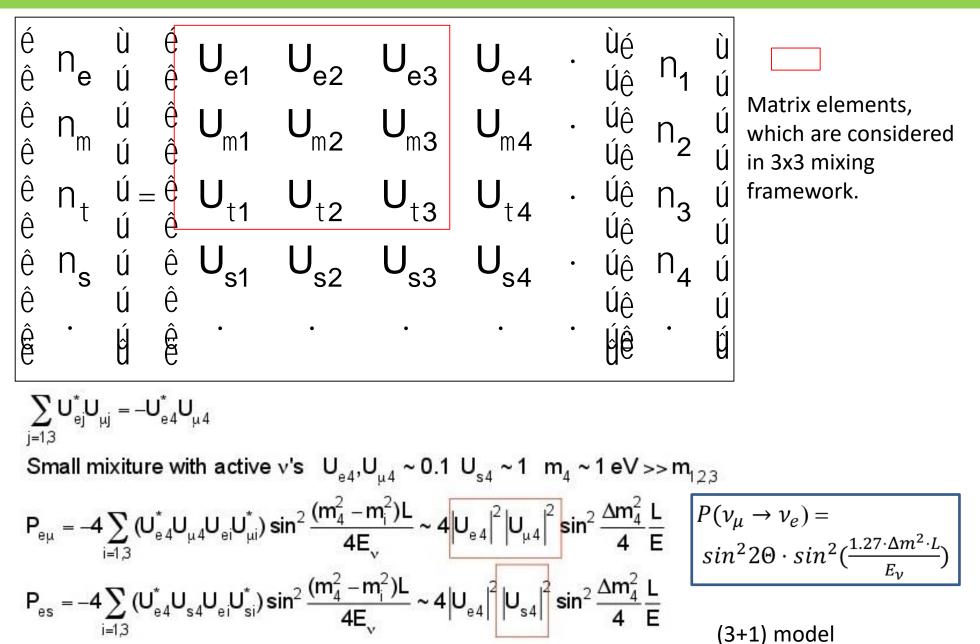
## indication of the sterile neutrino ( $\Delta m^2 \sim 1 eV^2$ ) ?

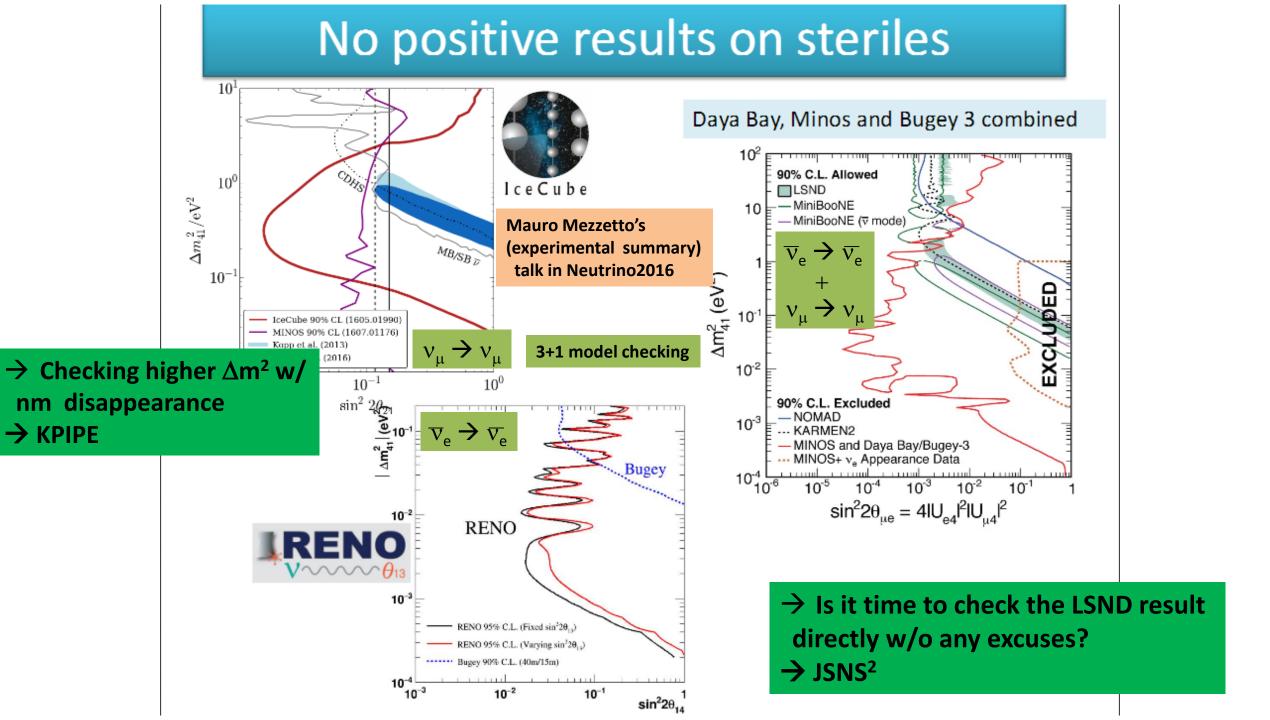
 Anomalies, which cannot be explained by standard neutrino oscillations for ~20 years are shown;

Experiments	Neutrino source	signal	significance	E(MeV),L(m)
LSND	$\mu$ Decay-At-Rest	$\overline{\nu}_{\mu} \rightarrow \overline{\nu_{e}}$	3.8σ	40,30
MiniBooNE	$\pi$ Decay-In-Flight	$\nu_{\mu} \rightarrow \nu_{e}$	3.4σ	800,600
		$\overline{\nu_{\mu}} \rightarrow \overline{\nu_{e}}$	2.8σ	
		combined	3.8σ	
Ga (calibration)	e capture	$v_e \rightarrow v_x$	2.7σ	<3,10
Reactors	Beta decay	$\overline{v}_{e} \rightarrow \overline{v}_{x}$	3.0σ	3,10-100

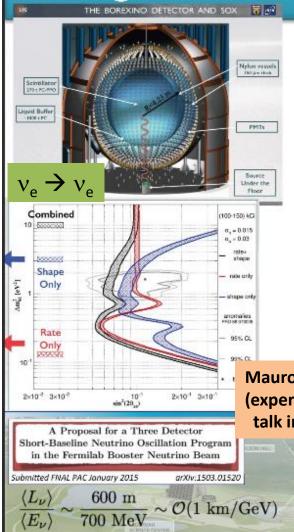
- Excess or deficit does really exist?
- The new oscillation between active and inactive (sterile) neutrinos?

#### Neutrino oscillations with $\Delta m^2 \sim 1 eV^2$ region





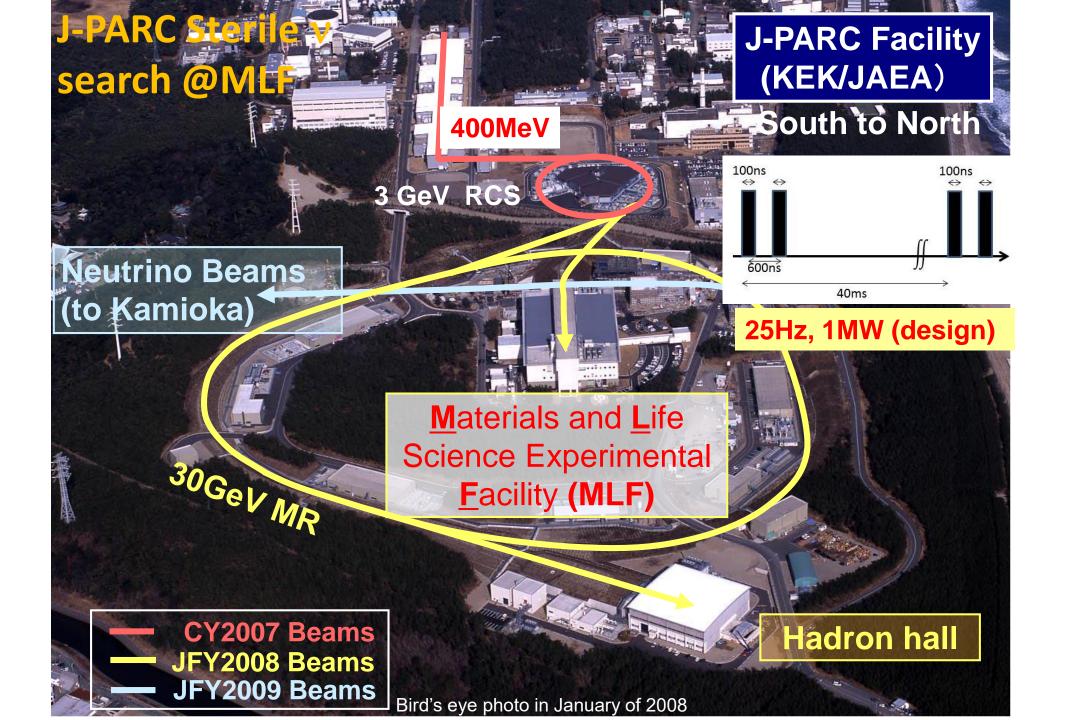
#### Next generation sterile experiments are almost ready



 $\sim v_{\mu} \rightarrow v_{e}$  (horn

Experimen	t Reactor Power/Fuel	Overburden (mwe)	Detection Material	Segmentation	Optical Readout	Particle ID Capability
DANSS (Russia)	3000 MW LEU fuel	~50	Inhomogeneous PS & Gd sheets	2D, ~5mm	WLS fibers.	Topology only
NEOS (South Korea)	2800 MW LEU fuel	~20	Homogeneous Gd-doped LS	none	Direct double ended PMT	recoil PSD only
nuLat (USA)	40 MW <sup>235</sup> U fuel	few	Homogeneous <sup>6</sup> Li doped PS	Quasi-3D, 5cm, 3-axis Opt. Latt	Direct PMT	Topology, reco & capture PSD
Neutrino4 (Russia)	100 MW <sup>235</sup> U fuel	~10	Homogeneous Gd-doped LS	2D, ~10cm	Direct single ended PMT	Topology only
PROSPECT (USA)	85 MW <sup>235</sup> U fuel	few	Homogeneous <sup>6</sup> Li-doped LS	2D, 15cm	Direct double ended PMT	Topology, reco & capture PSD
SoLid (UK Fr Bel US)	72 MW 235 U fuel	~10	Inhomogeneous <sup>6</sup> LiZnS & PS	Quasi-3D, 5cm multiplex	WLS fibers	topology, capture PSD
Chandler (USA)	72 MW 235 U fuel	~10	Inhomogeneous <sup>6</sup> LiZnS & PS	Quasi-3D, 5cm, 2-axis Opt. Latt	Direct PMT/ WLS Scint.	topology, capture PSD
Stereo (France)	57 MW <sup>235</sup> U fuel	~15	Homogeneous Gd-doped LS	1D, 25cm	Direct single ended PMT	recoil PSD
Mezzetto's nental sumr Neutrino201 Detec	.6 Distance		10 <sup>2</sup> Ve LAr Nass 10		MicroBooN	600, 6.6e+20 POT (600m IE, 1.32e+21 POT (470m ND, 6.6e+20 POT (100m v mode, CC Event
SBND Micro ICAR	BOONE 470	m 87	2 ton $(a_{2}^{2})^{2}$ ton $(a_{2}^{2})^{2}$ ton $(a_{2}^{2})^{2}$		Stot., X-3	Reconstructed Energ 80% v <sub>e</sub> Efficienc Sec., Flux, Cosmics, Di v <sub>e</sub> Only F — 90% CL — 30 CL
sed beam)		NEW HEAR DECISION	RECEIVE 10-1	50 ISND 99% CL + USND Bee Fit		5ơ CL

 $sin^2 2 \theta$ 



# JSNS<sup>2</sup> (J-PARC Sterile Neutrino Search at J-PARC Spallation Neutron Source)

#### Direct ultimate tests for LSND.

Collaboration meeting @ KEK 2017/May/11-12

JSNS<sup>2</sup> collaboration (53 collaborators)
10 Korean institutions (18 members)
6 Japanese institutions (28 members)
5 US institutions (7 members)

Technical Design Report (TDR): Searching for a Sterile Neutrino at J-PARC MLF  $(E56, JSNS^2)$ 

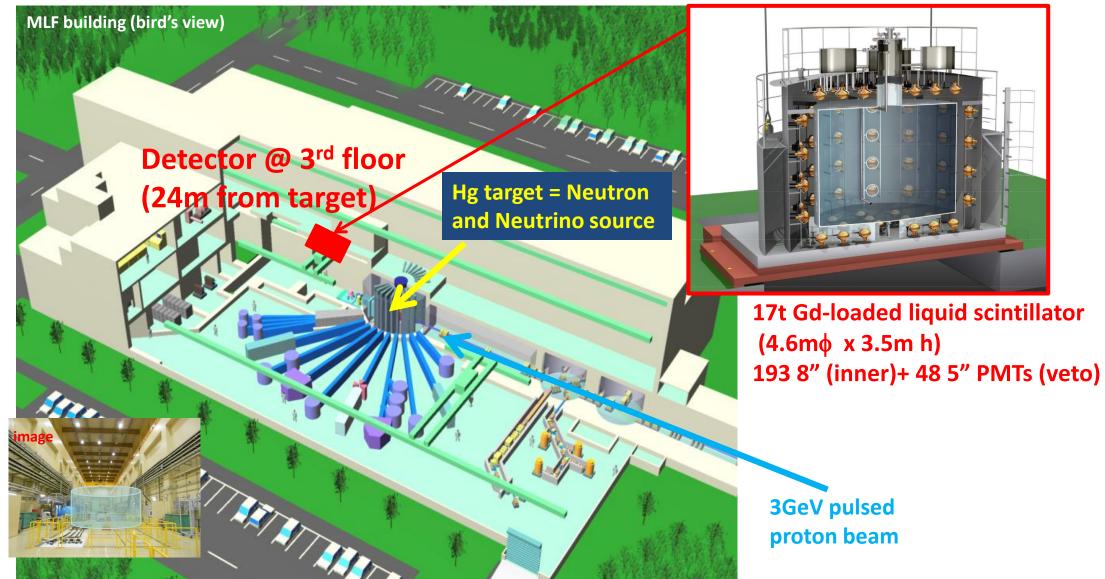
S. Ajimura<sup>1</sup>, M. K. Cheoun<sup>2</sup>, J. H. Choi<sup>3</sup>, H. Furuta<sup>4</sup>, M. Harada<sup>5</sup>, S. Hasegawa<sup>5</sup>,
Y. Hino<sup>4</sup>, T. Hiraiwa<sup>1</sup>, E. Iwai<sup>6</sup>, S. Iwata<sup>7</sup>, J. S. Jang<sup>8</sup>, H. I. Jang<sup>9</sup>, K. K. Joo<sup>10</sup>,
J. Jordan<sup>6</sup>, S. K. Kang<sup>11</sup>, T. Kawasaki<sup>7</sup>, Y. Kasugai<sup>5</sup>, E. J. Kim<sup>12</sup>, J. Y. Kim<sup>10</sup>,
S. B. Kim<sup>13</sup>, W. Kim<sup>14</sup>, K. Kuwata<sup>4</sup>, E. Kwon<sup>13</sup>, I. T. Lim<sup>10</sup>, T. Maruyama<sup>\*15</sup>,
T. Matsubara<sup>4</sup>, S. Meigo<sup>5</sup>, S. Monjushiro<sup>15</sup>, D. H. Moon<sup>10</sup>, T. Nakano<sup>1</sup>, M. Niiyama<sup>16</sup>,
K. Nishikawa<sup>15</sup>, M. Nomachi<sup>1</sup>, M. Y. Pac<sup>3</sup>, J. S. Park<sup>15</sup>, H. Ray<sup>17</sup>, C. Rott<sup>18</sup>, K. Sakai<sup>5</sup>,
S. Sakamoto<sup>5</sup>, H. Seo<sup>13</sup>, S. H. Seo<sup>13</sup>, A. Shibata<sup>7</sup>, T. Shima<sup>1</sup>, J. Spitz<sup>6</sup>, I. Stancu<sup>19</sup>,
F. Suekane<sup>4</sup>, Y. Sugaya<sup>1</sup>, K. Suzuya<sup>5</sup>, M. Taira<sup>15</sup>, W. Toki<sup>20</sup>, T. Torizawa<sup>7</sup>, M. Yeh<sup>21</sup>,

<sup>1</sup>Research Center for Nuclear Physics, Osaka University, Osaka, JAPAN <sup>2</sup>Department of Physics, Soongsil University, Seoul 06978, KOREA <sup>3</sup>Department of Radiology, Dongshin University, Chonnam 58245, KOREA <sup>4</sup>Research Center for Neutrino Science, Tohoku University, Sendai, Miyaqi, JAPAN <sup>5</sup>J-PARC Center, JAEA, Tokai, Ibaraki JAPAN <sup>6</sup>University of Michigan, Ann Arbor, MI, 48109, USA <sup>7</sup>Department of Physics, Kitasato University, Sagamihara 252-0373, Kanagawa, JAPAN <sup>8</sup>Gwanqju Institute of Science and Technology, Gwanqju, 61005, KOREA <sup>9</sup>Department of Fire Safety, Seoyeong University, Gwangju 61268, KOREA <sup>10</sup>Department of Physics, Chonnam National University, Gwangju, 61186, KOREA <sup>11</sup>School of Liberal Arts, Seoul National University of Science and Technology, Seoul, 139-743, KOREA <sup>12</sup>Division of Science Education, Physics major, Chonbuk National University, Jeonju, 54896, KOREA <sup>13</sup>Department of Physics and Astronomy, Seoul National University, Seoul 08826, KOREA <sup>14</sup>Department of Physics, Kyungpook National University, Daegu 41566, KOREA <sup>15</sup>High Energy Accelerator Research Organization (KEK), Tsukuba, Ibaraki, JAPAN <sup>16</sup>Department of Physics, Kyoto University, Kyoto, JAPAN <sup>17</sup>University of Florida, Gainesville, FL, 32611, USA <sup>18</sup>Department of Physics, Sungkyunkwan University, Gyeong Gi-do, KOREA <sup>19</sup>University of Alabama, Tuscaloosa, AL, 35487, USA <sup>20</sup>Colorado State University, Tuscaloosa, AL, 35487, USA <sup>21</sup>Brookhaven National Laboratory, Upton, NY, 11973-5000, USA

May 23, 2017

\*Spokesperson:(takasumi.maruyama@kek.jp)

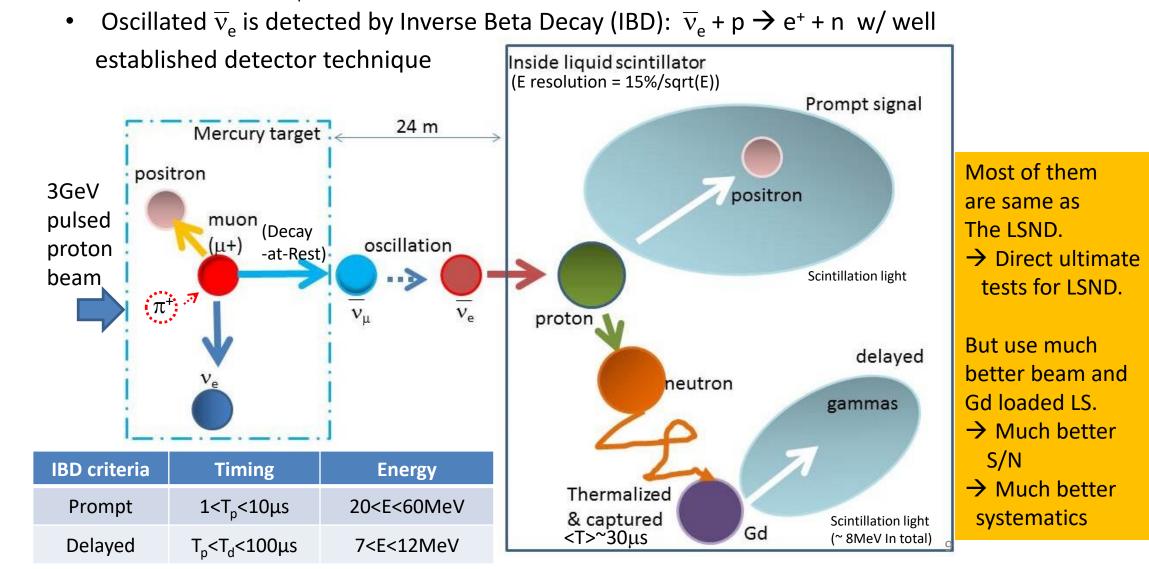
### JSNS<sup>2</sup> setup at J-PARC MLF



Searching for neutrino oscillation :  $\overline{v_{\mu}} \rightarrow \overline{v_{e}}$  with baseline of 24m. no new beamline, no new buildings are needed  $\rightarrow$  quick start-up

## Production / Detection

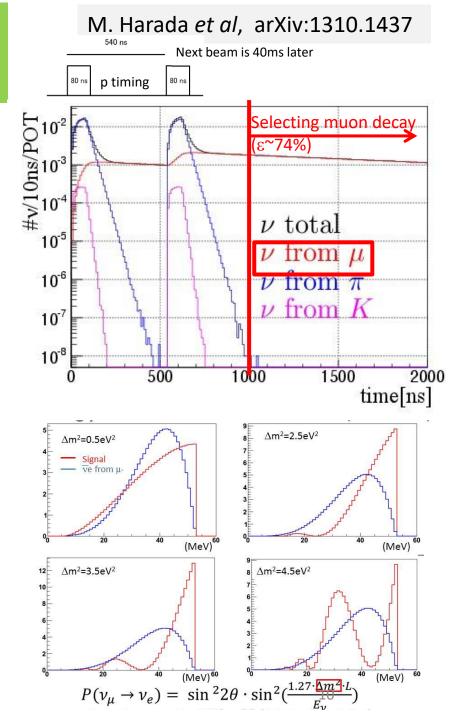
- Large amount of parent  $\mu$ + in Hg target  $\rightarrow \overline{v}_{\mu}$  are produced.
- If sterile v exist,  $\overline{v}_{\mu} \rightarrow \overline{v}_{e}$  oscillation is happened with 24m.

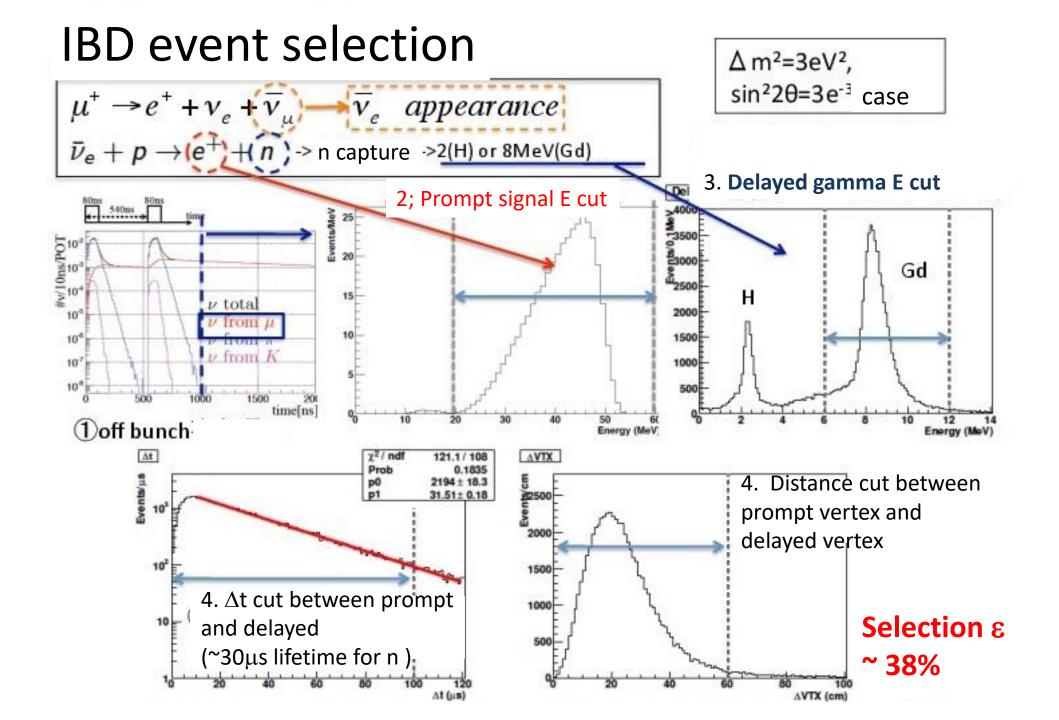


## **Timing and Energy**

#### Timing and Energy are friends of JSNS<sup>2</sup>

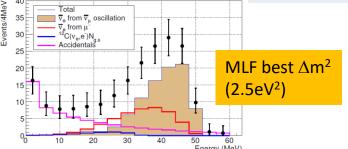
- > Timing: Ultra-pure v from  $\mu^+$  Decay-at-Rest
  - $\succ$  v from  $\pi$  and K -> removed with timing
  - Beam Fast neutrons -> removed w/ time
  - $\succ$  Cosmic ray BKG -> reduced by 9µs time window.
- Energy: signals / BKG separation by energy.
   ν from μ has well-known spectrum.
  - Energy reconstruction is very easy at the IBD. (Ev ~ Evis + 0.8MeV)
  - $\succ$  v from  $\mu$  is high suppressed.





## #events (1MW x 3 years x 1 detector (17tons))

Source	contents	#ev.(17tons x 3years)	Reference : SR2014 (50tons x 5 years)	comments
background	$\overline{\nu_e}$ from $\mu$ -	43	237	Dominant BKG
	<sup>12</sup> C(v <sub>e</sub> ,e-) <sup>12</sup> N <sub>g.s.</sub>	3	16	
	Beam fast neutrons	Consistent with 0 < 2 ( <u>90%CL UL</u> )	<13	Based on real data
	Fast neutrons (cosmic)	~0	37	
	Accidental	20	32	Based on real data
signal		87	480	$\Delta m^2$ =2.5, sin <sup>2</sup> 2 $\theta$ =0.003
		62	342	$\Delta m^2$ =1.2, sin <sup>2</sup> 2 $\theta$ =0.003

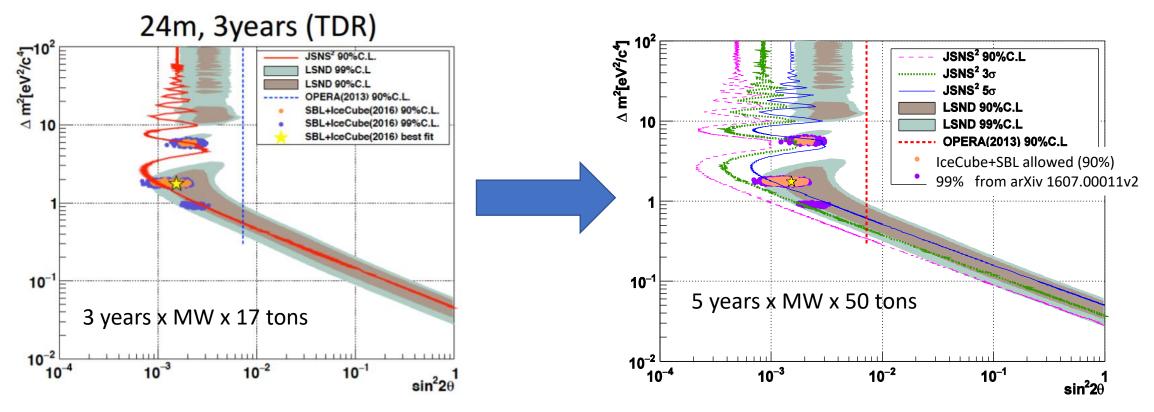


Accidental BKG is calculated by; R acc =  $\Sigma R_{prompt} \times \Sigma R_{delay} \times \Delta_{VTX} \times N_{spill}$ 

- $-\Sigma R_{prompt}$ ,  $\Sigma R_{delay}$  are probability of accidental BKG for prompt and delayed.
  - $\Delta_{\rm VTX}$  ; BKG rejection factor of **50**.
  - $N_{spill}$ (#spills / 5 years) = 1.9x10<sup>9</sup>

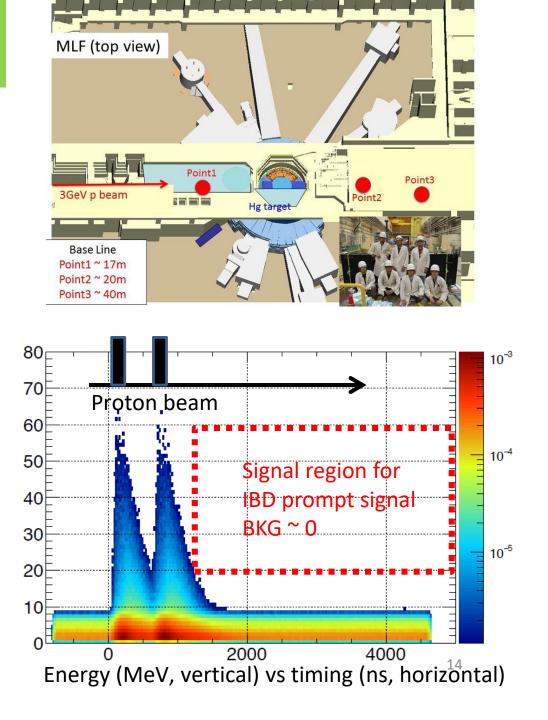
## Sensitivity / Upgrade

- To have a good international competition capability, we start the experiment with one detector (17tons fiducial volume).
- Even with one detector, we have a good 90% C.L constraints for the best fit point of global fit (of sterile neutrino searches) for 3 years. Left plot
- Meanwhile, we are making effort to obtain the budget to build the 2<sup>nd</sup> detector. (and enlarged acrylic tanks). This upgrade can make  $5\sigma$  significance test for the best fit point of the global fit.



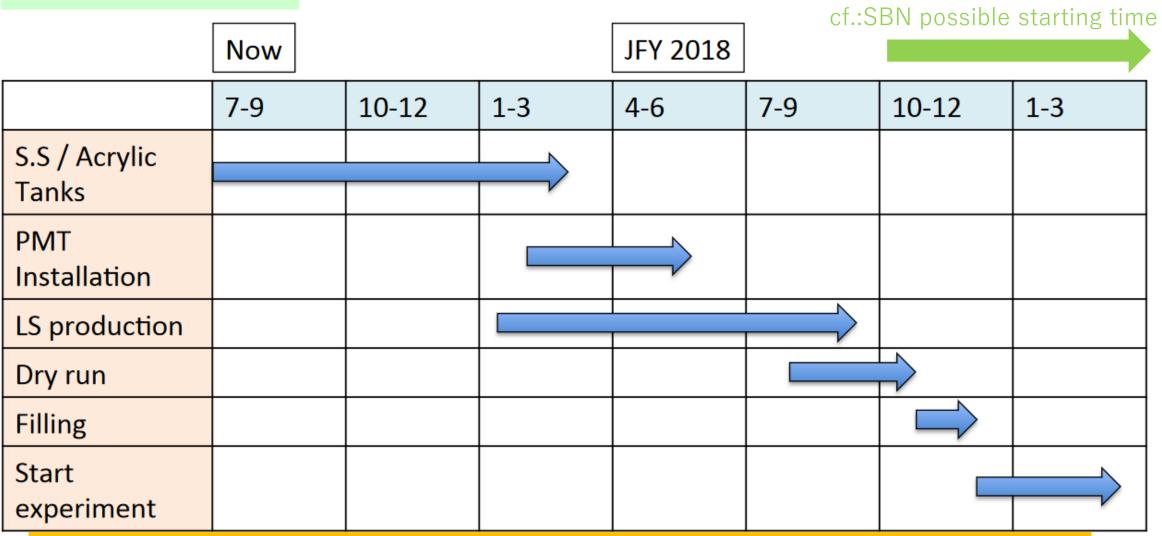
## Achievements so far

- 2013 Sep; A proposal was submitted to the J-PARC PAC
- 2014 Apr-Jul; We measured the BKG rate on 3<sup>rd</sup> floor. -> manageable beam /cosmic BKGs to perform JSNS<sup>2</sup> PTEP 2015 6, 063C01 / arXiv:1502.02255
- 2014-Dec; The result was reported to J-PARC PAC. → the stage-1 status was obtained from J-PARC /KEK
- The performance check of detector and safety discussions are being performed.
- 2016-June: The grant-in-aid was approved for one detector construction
- 2017-May: Technical Design Report was submitted to J-PARC PAC and arXiv (arXiv:1705.08629 [physics.ins-det])
- We aim to start JSNS<sup>2</sup> in JFY2018



## Schedule & Status

#### **Overall schedule**



(1) A stainless tank will be produced in JFY2017.(2) PMTs can be made within a half of year.(3) The bid for the acrylic tank will be started soon.(4) LS will be produced at Korea.

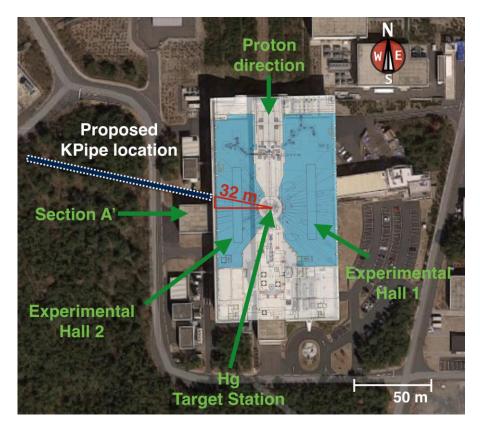
# KPIPE

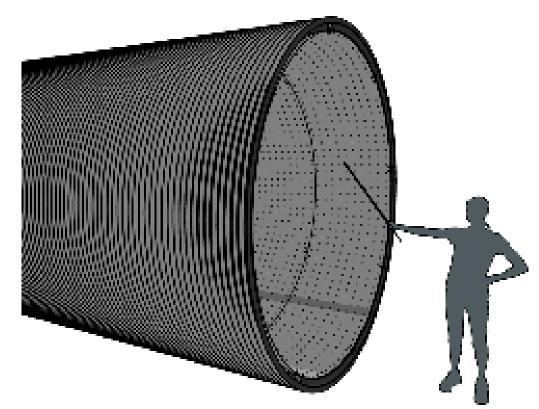
 $v_{\mu} \rightarrow v_{\mu}$  is also important.

# **KPIPE**

Axani, Collin, Conrad, Shaevitz, Spitz, Wongjirad, Phys Rev. D 92 092010 (2015)

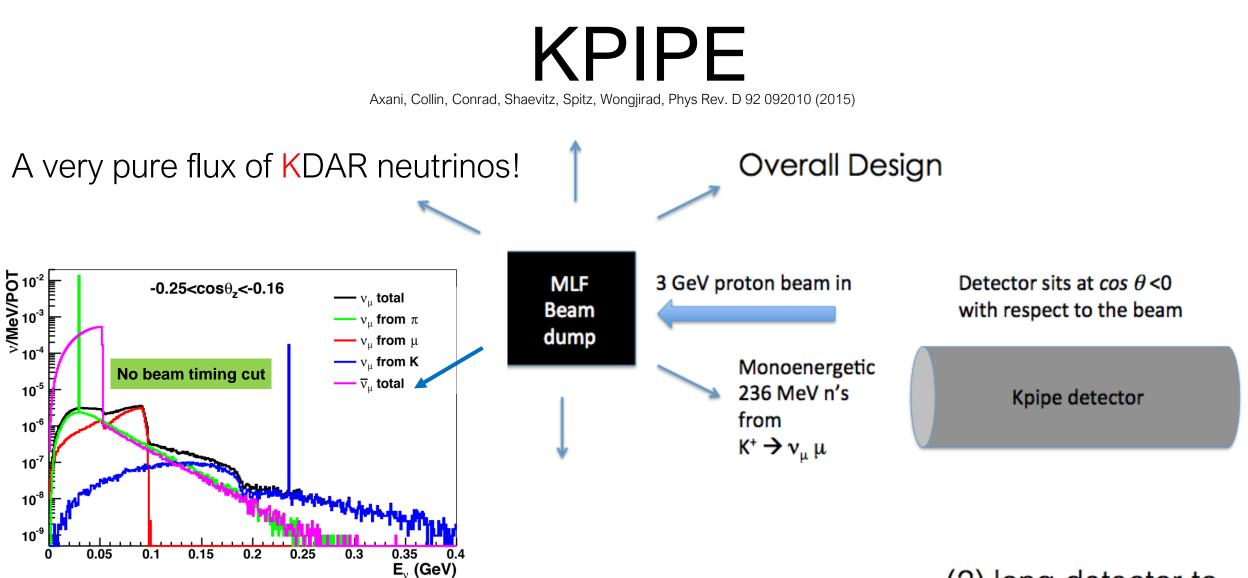
#### The idea: Use a very long liquid scintillator detector to look for $v_{\mu}$ disappearance (in *L*) using 236 MeV KDAR $v_{\mu}$ CC events





Long LS detector surrounded by SiPMs

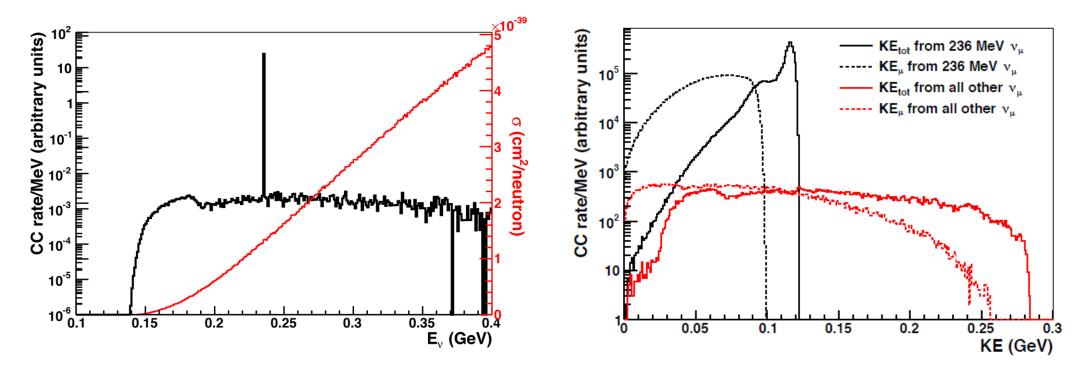
@ J-PARC MLF



(1) pure, mono-energetic flux of muon neutrinos (2) long detector to measure the oscillation wave

# The beauty of KPIPE

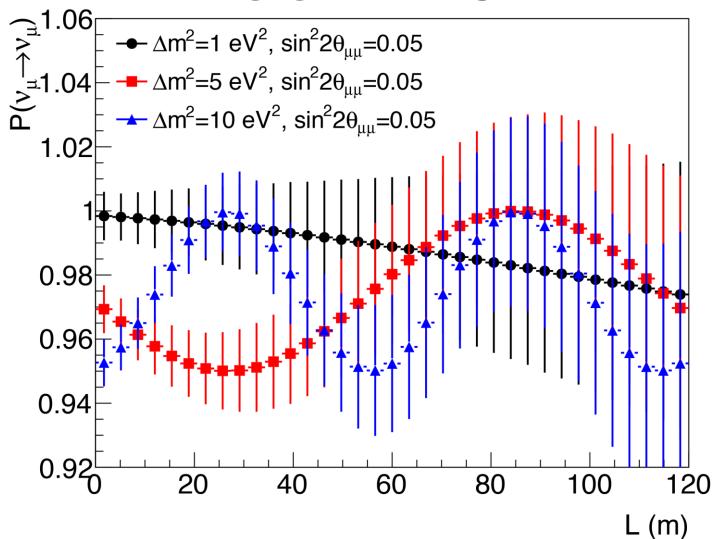
If you detect a numu CC event, you can be 98.5% sure that it was a 236 MeV muon neutrino!



Since you know the energy of the neutrino, you don't need to worry about energy resolution. KPIPE calls for 0.4% photocoverage. Estimated cost of experiment: \$4.5M

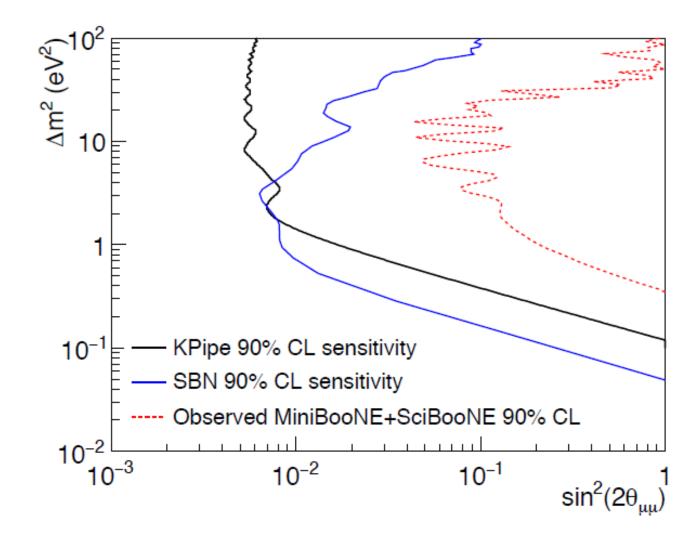
KPIPE cost document: <u>http://hdl.handle.net/1721.1/98388</u>

# KPIPE; what would a signal loc



# **KPIPE** sensitivity

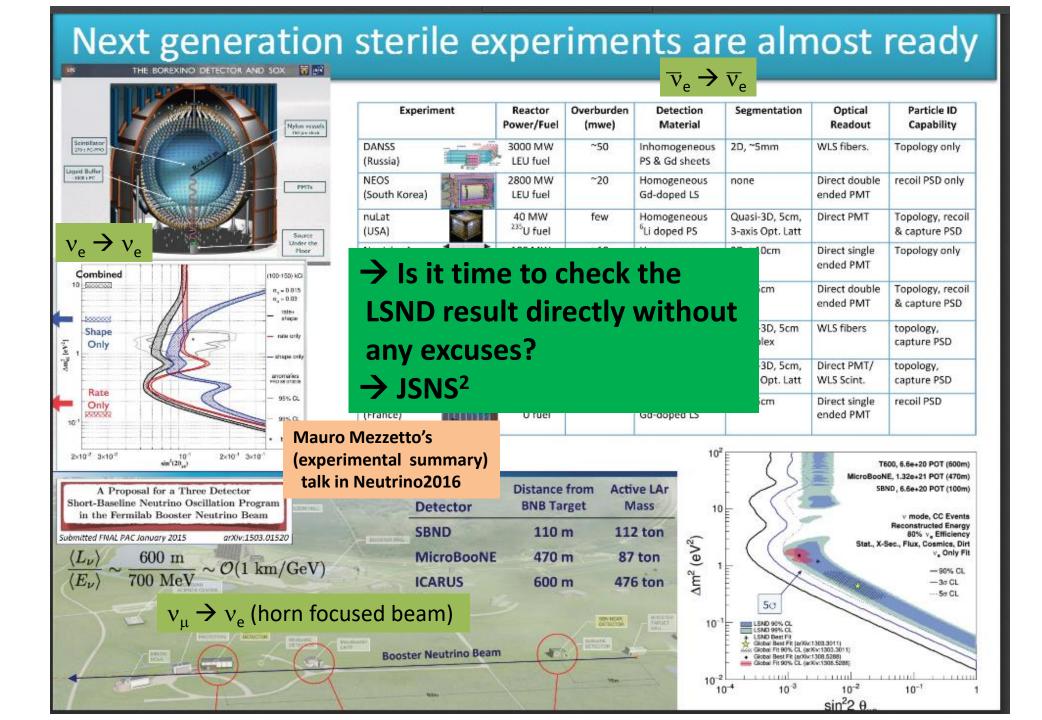
- 6 years of running
- Extends limit at high-Δm<sup>2</sup> by an order of magnitude.
- Highly complementary to SBN program.
  - 6 years of MicroBooNE
  - 3 years of T600 and SBND.

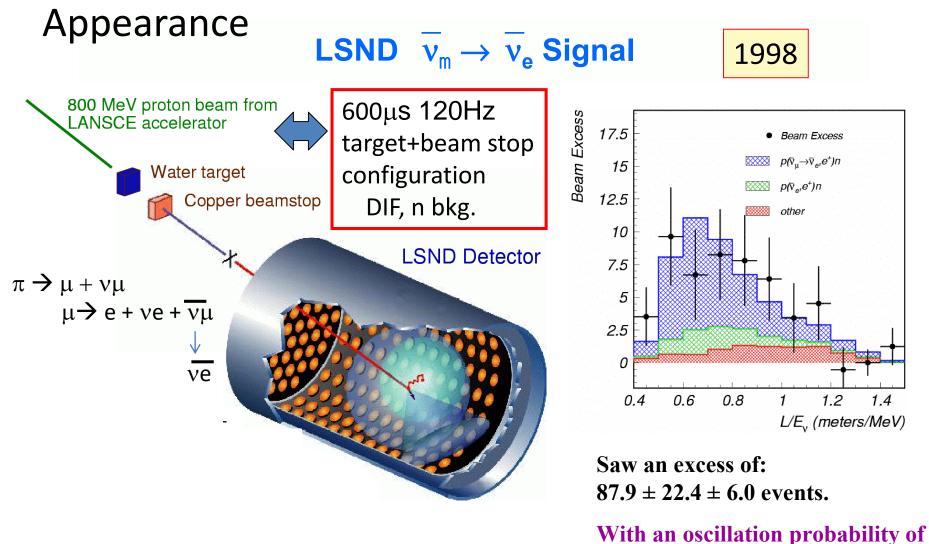


# Summary

- There are fascinate programs for sterile neutrino searches at J-PARC MLF. ← large number of neutrinos are produced at the mercury target.
  - JSNS<sup>2</sup>: Direct test for the LSND anomaly without any excuses  $(\overline{v_{\mu}} \rightarrow \overline{v_{e}})$ 
    - Using ultra pure  $\mu$  Decay-At-Rest neutrinos due to nice beam timing.
    - Gd loaded LS.
    - Improves S/N (by more than 100) and systematic uncertainties.
    - Budget to build one detector was granted in 2016.
    - Technical Design Report (TDR) was submitted to PAC and arXiv:1705.08629.
    - Aim to start the experiment in JFY2018.
  - KPIPE:  $v_{\mu}$  disappearance experiment
    - Using 236 MeV monochromatic neutrinos from Kaon Decay-At-Rest.
    - Oscillation curve can be seen as a function of L. (E is one number.)
    - Comparable sensitivity as SBN program with reasonable cost.
    - Letter of Intent (LoI) was submitted to J-PARC PAC (in 2016).







 $\pi^-$ ,  $\mu^-$  absorbed before decay into v's there should not be ve at the level of 7x10<sup>-4</sup>

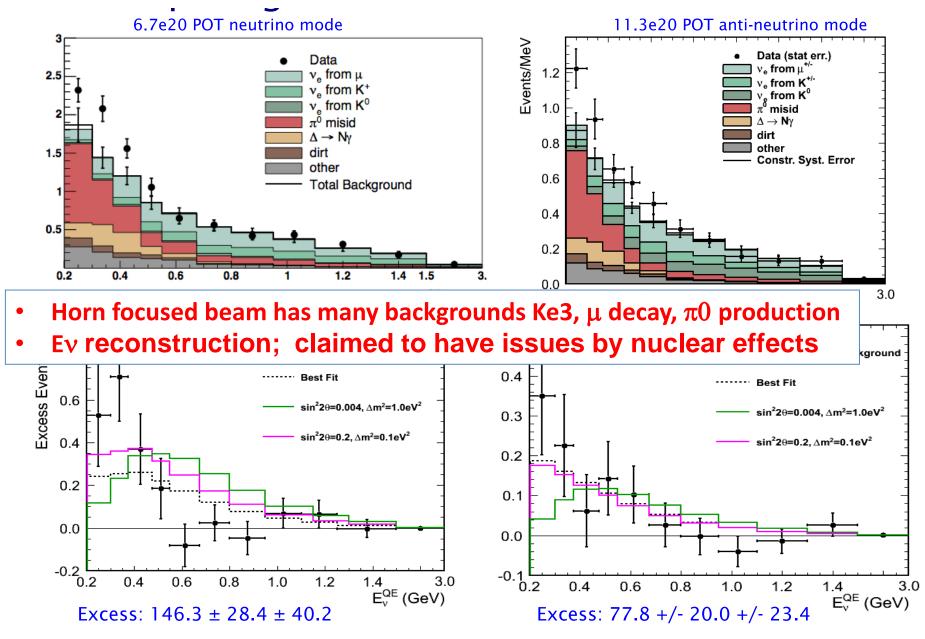
Signal :  $\overline{ve} p \rightarrow e^+ n np \rightarrow d\gamma(2.2 MeV)$ 

 $(0.264 \pm 0.067 \pm 0.045)\%$ .

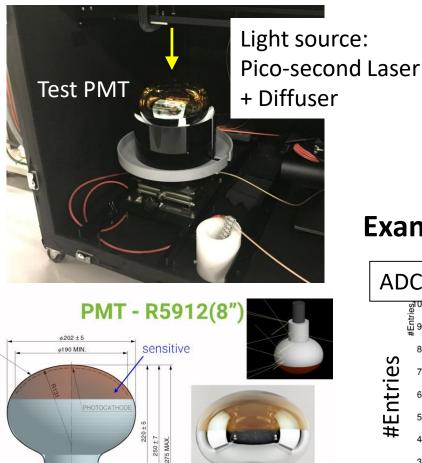
**3.8** S evidence for oscillation.

#### Trial by MiniBooNE @ Fermilab

#### Chris Polly NEUTRINO2012



# 8" PMT pre-calibration



BASE

C No. B20-102

nd (D) he

\$51.2 ± 0.5

#### Test item:

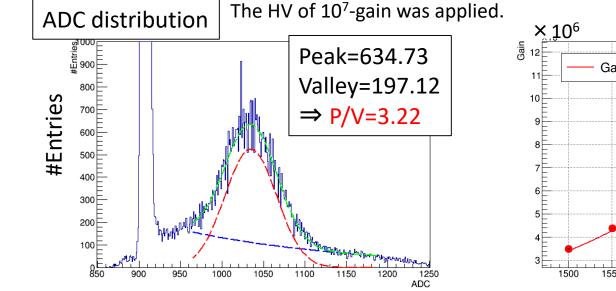
Gain, Peak-to-Valley ratio, Transit time spread, QE × CE Status:

Test conditions are adjusting using an 8"PMT sample.

- Light intensity, PMT position, etc...

#### **Examples of measurement**

There are almost consistent with Hamamatsu evaluations.



**Second Second S**