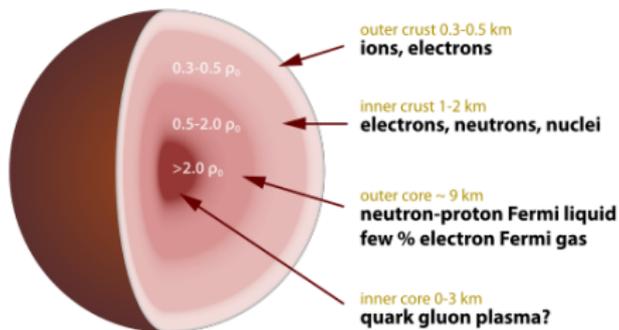


# The Polarizability of $^{208}\text{Pb}$

## A constraint on the Neutron Matter EoS

Jochen Wambach

TU-Darmstadt and GSI  
Germany



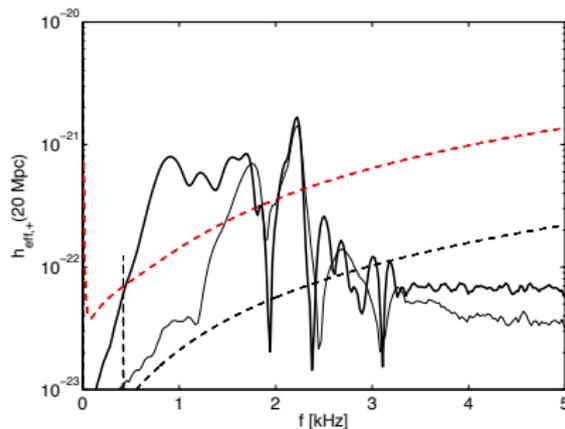
Compactness of a Neutron star ( $M/R$ )

# Binary NS mergers

## gravitational-wave signal

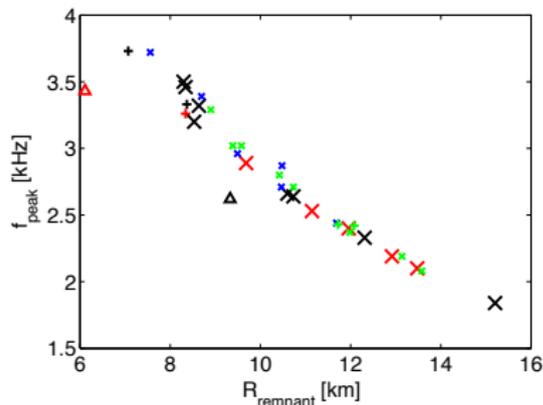
A. Bauswein et al. 2012

GW spectrum



merger of two  $M_* = 1.35 M_\odot$   
neutron stars

frequency-radius correlation

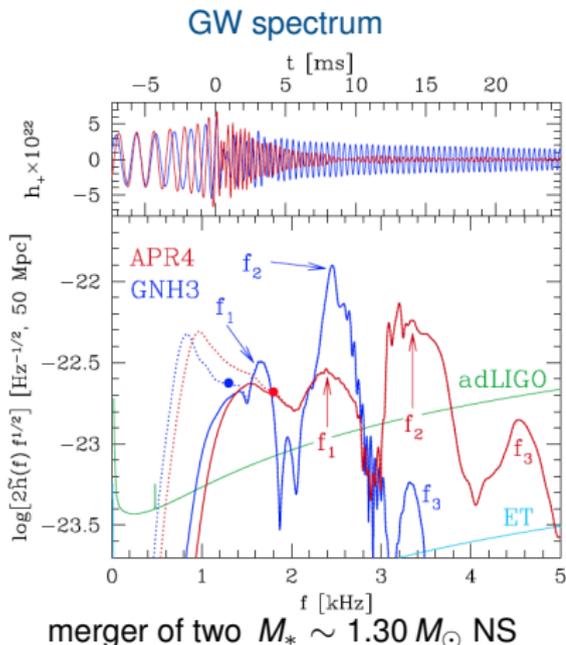


strong constraint on the  
high-density EoS

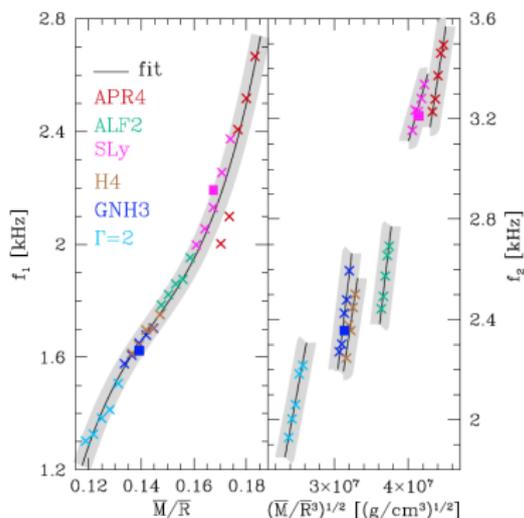
# Binary NS mergers

## gravitational-wave signal

K. Takami et al. 2014



frequency-compactness correlation

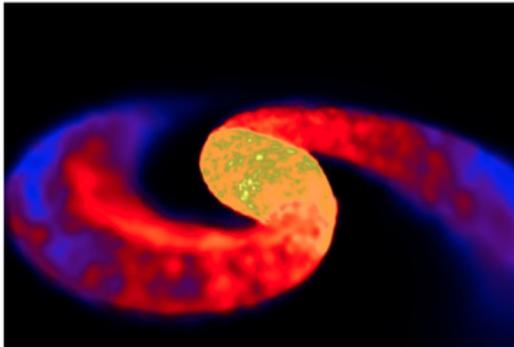


# Binary NS mergers

## r-process site

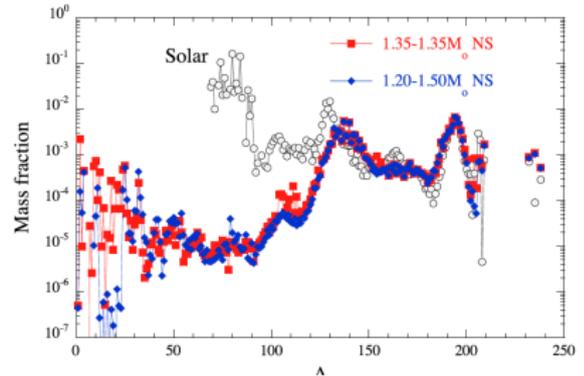
### mass ejection

D. Price and S. Roswog



### r-process abundance

S. Goriely et al. 2011



r-process abundance depends on amount of ejected material ( $\sim 0.01 M_{\odot}$ )

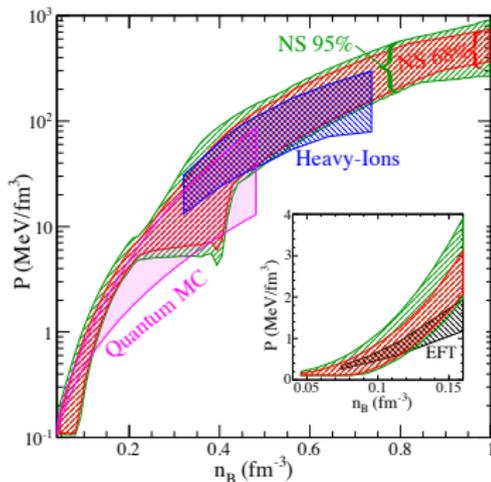
compactness crucial!

# Neutron Stars

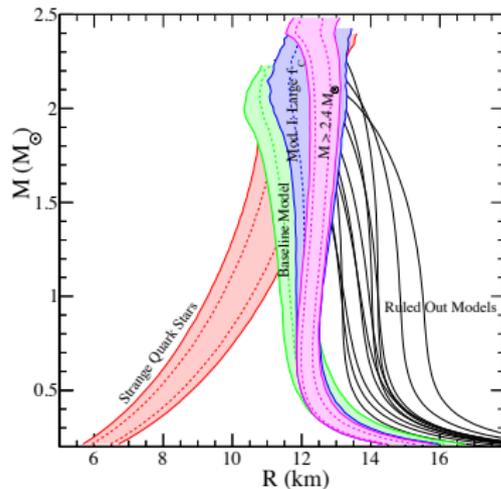
## equation of state

recent analysis A. Steiner et al. 2012

pressure vs density



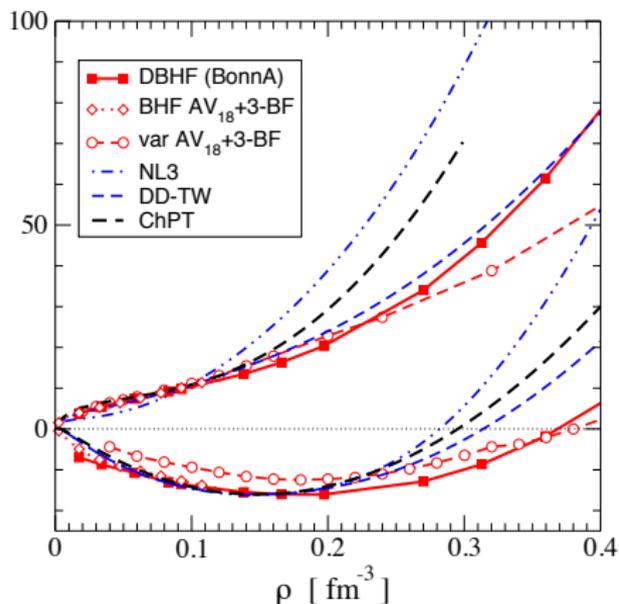
mass-radius constraints



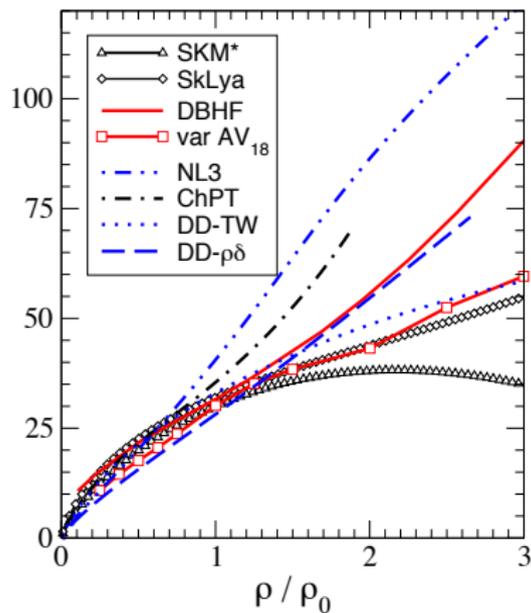
PSR J1614-2230  $M_* = 1.97 \pm 0.04 M_\odot$   
PSR J0348+0432  $M_* = 2.01 \pm 0.04 M_\odot$

more info from binary NS mergers  
through GW signals

nuclear EoS



symmetry energy



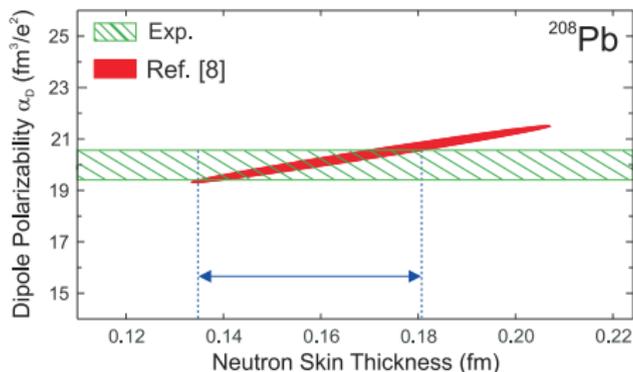


# Neutron Stars

## nuclear polarizability

correlation between 'neutron skin' of  $^{208}\text{Pb}$   
and the dipole polarizability

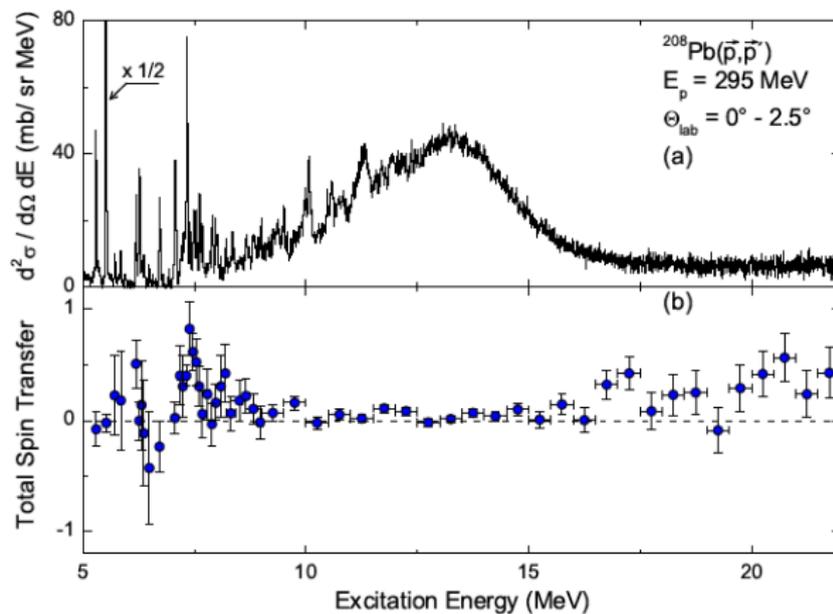
$$\alpha_D = \frac{8\pi}{9} e^2 \int d\omega \frac{S_D(\omega)}{\omega}$$



P. G. Reinhard and W. Nazarewicz 2010

# Neutron Stars

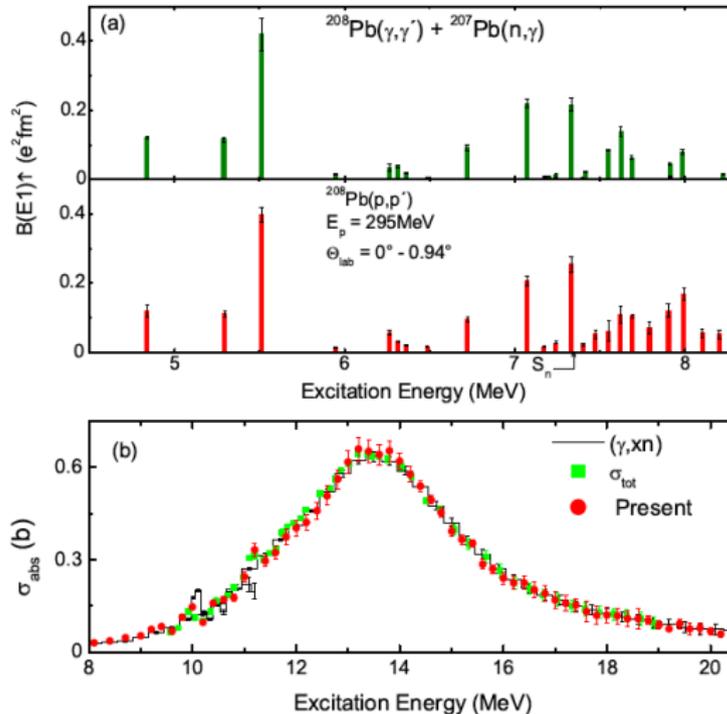
## The experiment



A. Tamii et al. 2011

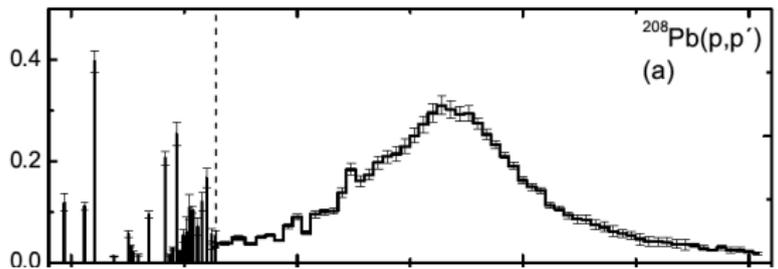
# Neutron Stars

## Experimental Consistency

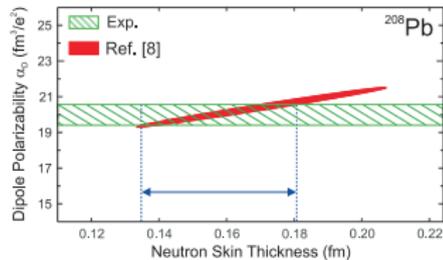


# Neutron Stars

## Results



$$\alpha_D = \frac{8\pi}{9} e^2 \int d\omega \frac{S_D(\omega)}{\omega} \quad \rightarrow \quad \alpha_D = 20.1 \pm 0.6 \quad \text{A. Tamii et al. 2011}$$



$\Delta r_{np} = 0.156 \pm 0.025 \text{ fm}$   
most precise value to date!

# Neutron Stars

## systematics

