

*Int. School on Nuclear Physics 2013: Neutrino Physics – Present and Future
Erice/Italy, September 16-24, 2013*

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Introduction

Direct Neutrino Mass determination

- Rhenium β decay and EC experiments
- Tritium β decay experiments

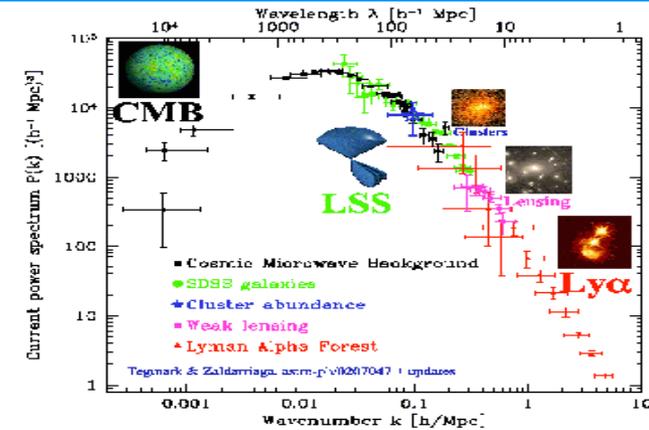
The Karlsruhe Tritium Neutrino experiment KATRIN

Summary and Outlook

Three complementary ways to the absolute neutrino mass scale

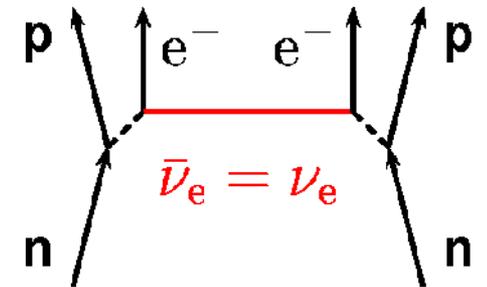
1) Cosmology

very sensitive, but model dependent
compares power at different scales
current sensitivity: $\Sigma m(\nu_i) \approx 0.5 \text{ eV}$



2) Search for $0\nu\beta\beta$

Sensitive to Majorana neutrinos
First upper limits by EXO-200, KamLAND-Zen, GERDA



3) Direct neutrino mass determination:

No further assumptions needed, use $E^2 = p^2c^2 + m^2c^4 \Rightarrow m^2(\nu)$ is observable mostly

Time-of-flight measurements (ν from supernova)

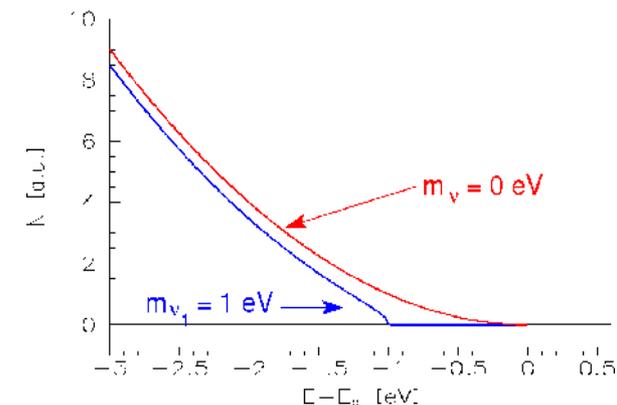
SN1987a (large Magellan cloud) $\Rightarrow m(\nu_e) < 5.7 \text{ eV}$

Kinematics of weak decays / beta decays

measure charged decay prod., E-, p-conservation

β -decay searches for $m(\nu_e)$ - tritium β spectrometers

- ^{187}Re , ^{163}Ho bolometers

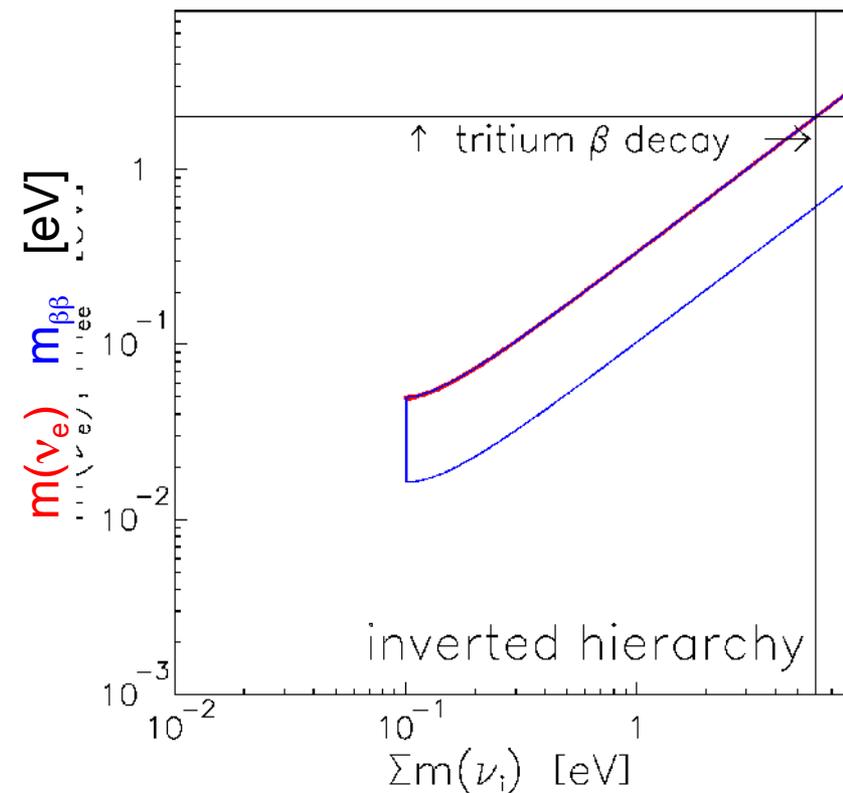
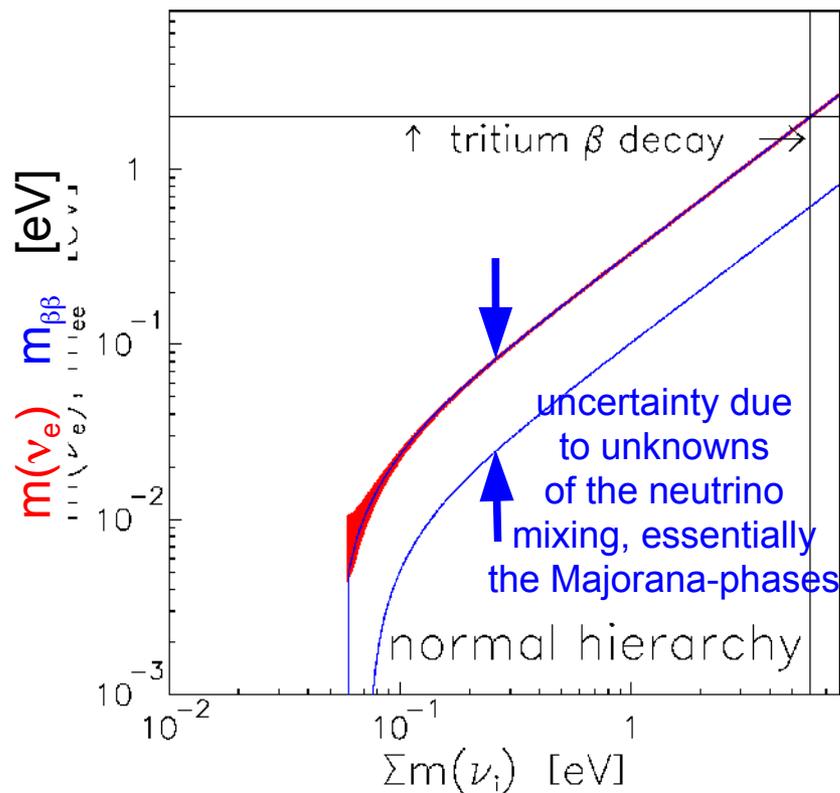


Comparison of the different approaches to the neutrino mass

Direct kinematic measurement: $m^2(\nu_e) = \sum |U_{ei}|^2 m^2(\nu_i)$ (incoherent)

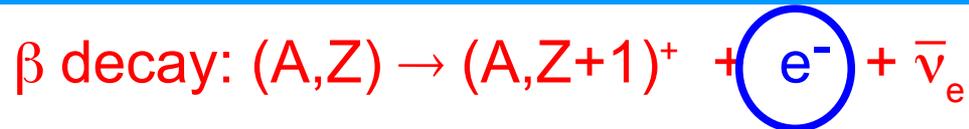
Neutrinoless double β decay: $m_{\beta\beta}(\nu) = |\sum |U_{ei}|^2 e^{i\alpha(i)} m(\nu_i)|$ (coherent)

if no other particle is exchanged (e.g. R-violating SUSY)
problems with uncertainty of nuclear matrix elements



⇒ absolute scale/cosmological relevant neutrino mass in the lab by single β decay

Direct determination of $m(\nu_e)$ from β decay

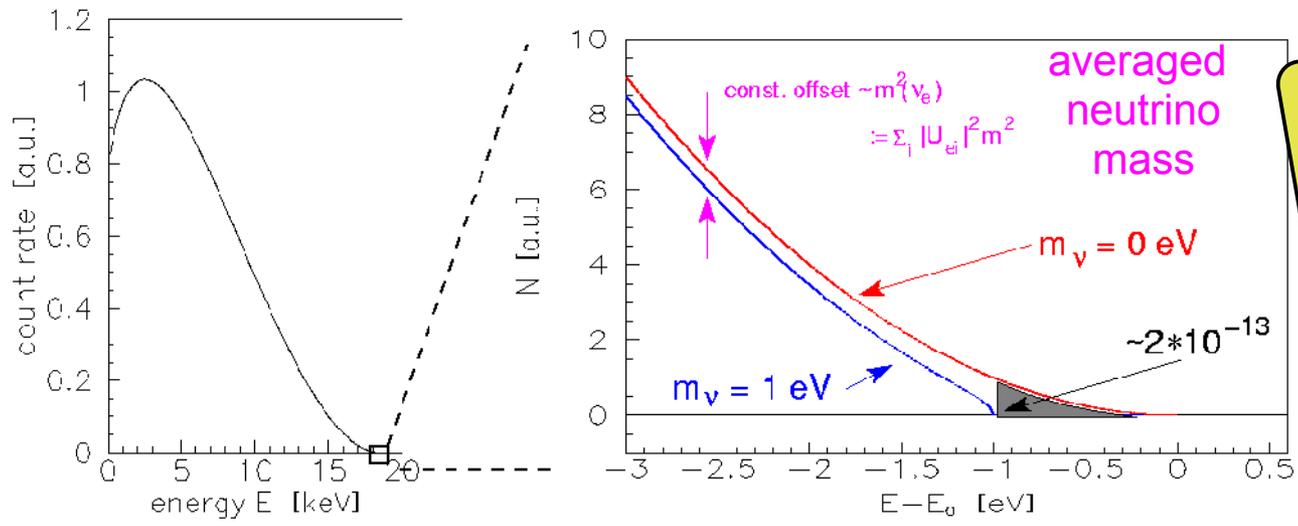


β electron energy spectrum:

$$dN/dE = K F(E,Z) p E_{\text{tot}} (E_0 - E_e) \sqrt{(E_0 - E_e)^2 - m(\nu_e)^2}$$

(modified by electronic final states, recoil corrections, radiative corrections)

Complementary to $0\nu\beta\beta$
and cosmology



E.W. Otten & C. Weinheimer
Rep. Prog. Phys.
71 (2008) 086201
G. Drexlin, V. Hannen, S. Mertens,
C. Weinheimer, Adv. High Energy
Phys., 2013 (2013) 293986

Need: low endpoint energy
very high energy resolution &
very high luminosity &
very low background } \Rightarrow Tritium ^3H , (^{187}Re)
 \Rightarrow MAC-E-Filter
(or bolometer for ^{187}Re)

Summary: β -spectrum incl. electronic final states + ν mixing

Including electronic excited final states of excitation energy V_j with probability W_j

$$W_j = |\langle \Psi_0 | \Psi_{f,j} \rangle|^2$$

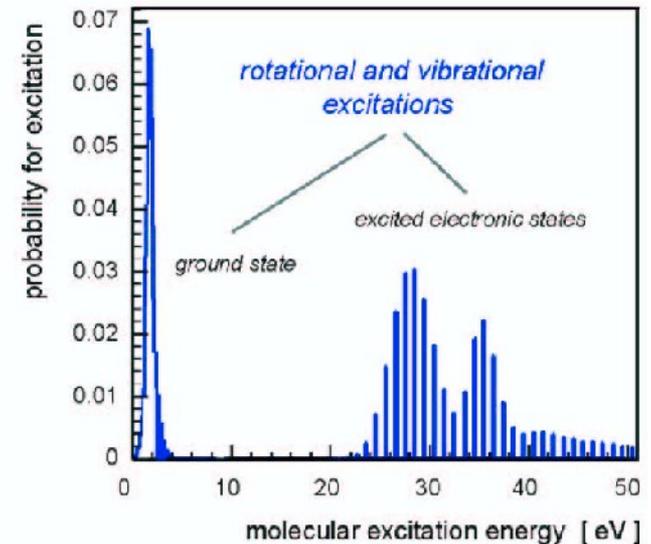
Using $\epsilon_j = E_0 - V_j - E$

$$\frac{d^2 N}{dt dE} = A \cdot F(E, Z + 1) \cdot p \cdot (E + m) \cdot \sum_j W_j \cdot \epsilon_j \cdot \sqrt{\epsilon_j^2 - m^2(\nu_c)} \cdot \Theta(\epsilon_j - m(\nu_c))$$

⇒ **electronic final states are important**

Final states of T_2 β -decay:

A. Saenz et al. Phys. Rev. Lett. 84 (2000) 242,
N. Doss et al., Phys. Rev. C73 (2006) 025502



Including neutrino mixing

$$\frac{d^2 N}{dt dE} = A \cdot F(E, Z + 1) \cdot p \cdot (E + m) \cdot \sum_j W_j \cdot \epsilon_j \cdot \left(\sum_i |U_{ci}|^2 \sqrt{\epsilon_j^2 - m^2(\nu_i)} \cdot \Theta(\epsilon_j - m(\nu_i)) \right)$$

⇒ "Electron neutrino mass"

$$m^2(\nu_c) := \sum_i |U_{ci}|^2 \cdot m^2(\nu_i)$$

⇒ **the different $m(\nu_i)$ are not important at present precision**

Summary: β -spectrum incl. electronic final states + ν mixing

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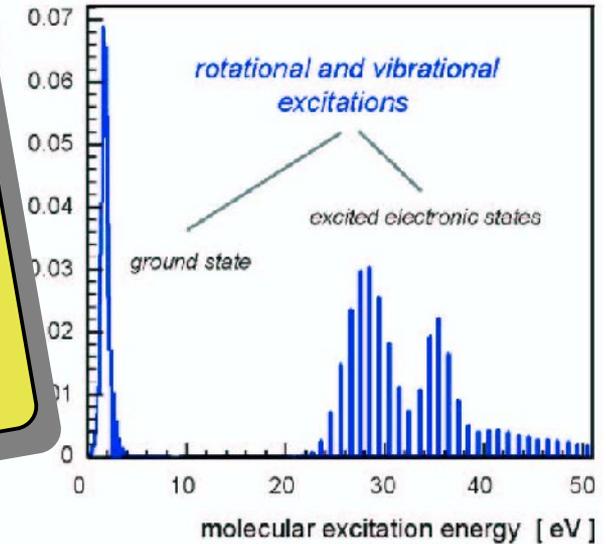
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⇒ **electronic final states are important**

Final states of
A. Saenz et al. P
N. Doss et al., P

The electron spectrum coming out of a β -source is even more complicated due to inelastic scattering, backscattering. ...



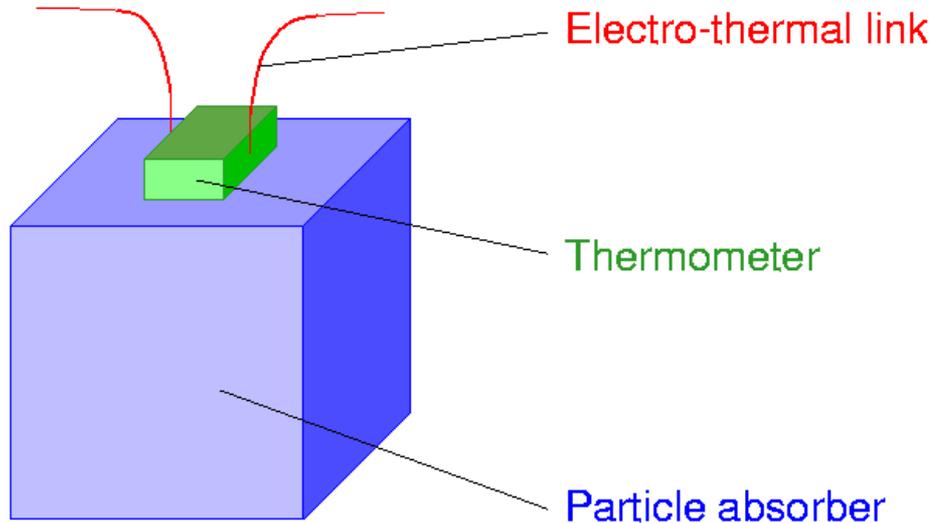
Including neutrino mixing

$$\frac{d^2 N}{dt dE} = A \cdot F(E, Z) \cdot \sum_j W_j \cdot \epsilon_j \cdot \left(\sum_i |U_{ci}|^2 \sqrt{\epsilon_j^2 - m^2(\nu_i)} \cdot \Theta(\epsilon_j - m(\nu_i)) \right)$$

⇒ "Electron neutrino mass"

$$m^2(\nu_c) := \sum_i |U_{ci}|^2 \cdot m^2(\nu_i)$$

⇒ **the different $m(\nu_i)$ are not important at present precision**



$$\Delta T = \Delta E / C$$

Debye:

$$C \sim T^3$$

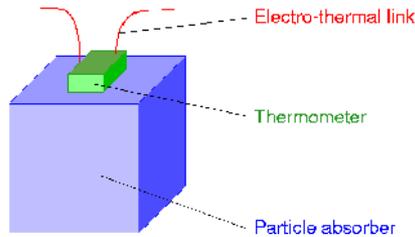
Measures temperature rise by β -decay in an absorber

- all energy except that of the neutrino is measured
- „single final state experiment“, no problems with inelastic scattering, backscattering, ...

Disadvantage:

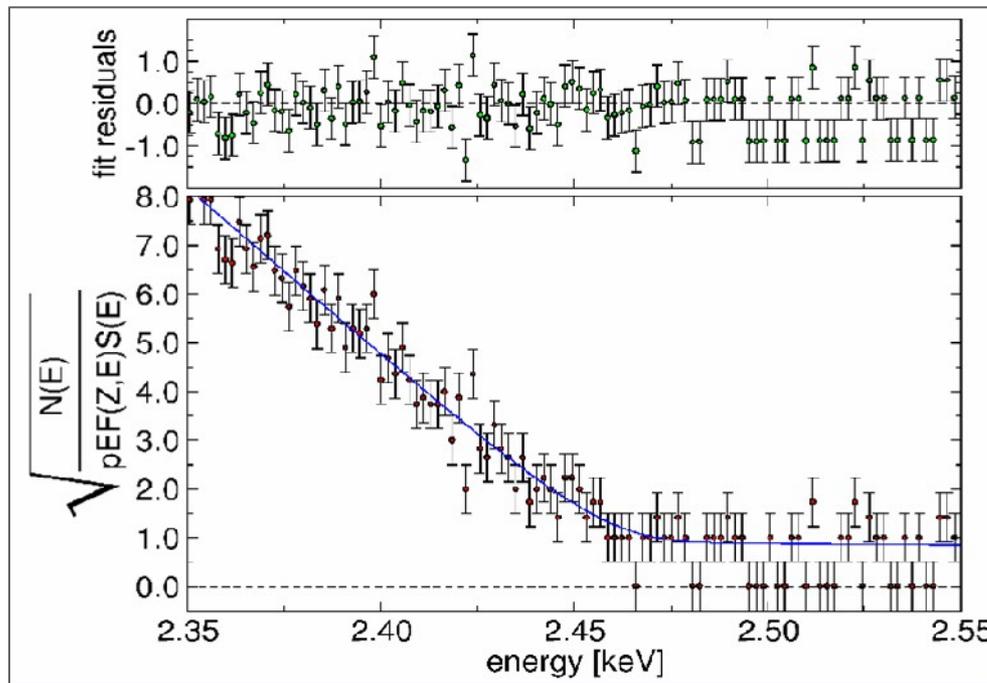
- measure whole spectrum at once
- pile-up problem
- need many detector pixels

Cryogenic bolometers with ^{187}Re MIBETA (Milano/Como)



$$\Delta T = \Delta E / C$$

Debye:
 $C \sim T^3$



Measures all energy except that
of the neutrino

detectors: 10 (AgReO_4)

rate each: 0.13 1/s

energy res.: $\Delta E = 28 \text{ eV}$

pile-up frac.: $1.7 \cdot 10^{-4}$

$$M_\nu^2 = -141 \pm 211_{\text{stat}} \pm 90_{\text{sys}} \text{ eV}^2$$

$$M_\nu < 15.6 \text{ eV (90\% c.l.)}$$

(M. Sisti et al., NIMA520 (2004) 125)

MANU (Genova)

- Re metallic crystal (1.5 mg)
- BEFS observed (F.Gatti et al., Nature 397 (1999) 137)
- sensitivity: $m(\nu) < 26 \text{ eV}$ (F.Gatti, Nucl. Phys. B91 (2001) 293)

MARE neutrino mass project: ^{167}Re beta decay with cryogenic bolometers

Advantages of cryogenic bolometers:

- measures all released energy except that of the neutrino
- no final atomic/molecular states
- no energy losses
- no back-scattering

Challenges of cryogenic bolometers:

- measures the full spectrum (pile-up)
- need large arrays to get statistics
- understanding spectrum
- still energy losses or trapping possible
- beta environmental fine structure

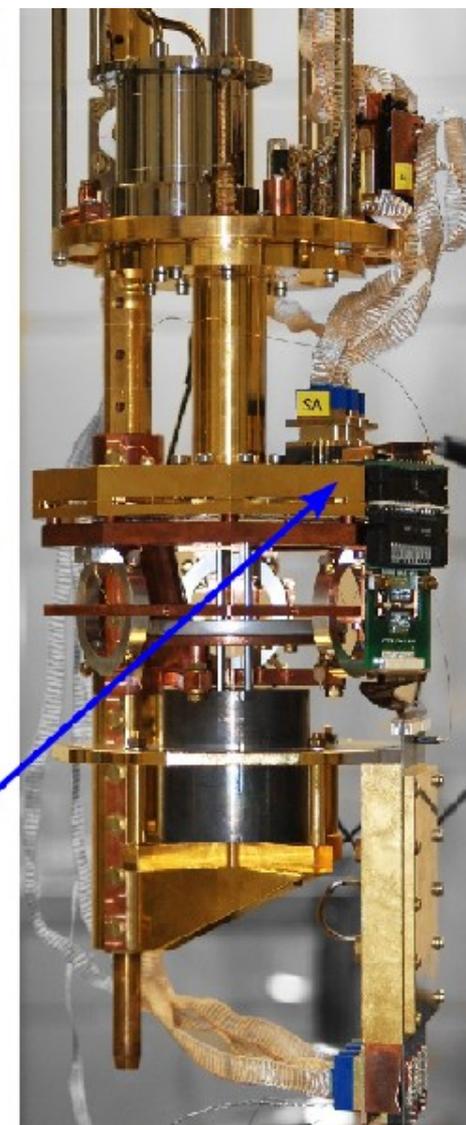
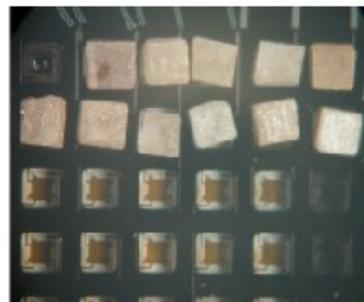
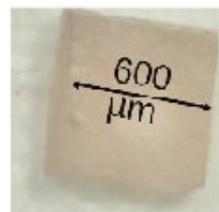
MARE-2 aims for 10^4 to 10^5 detectors with much more advanced time & energy resolution

MARE-1 @ Genova

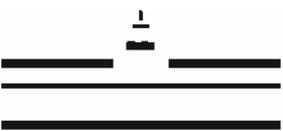
- R&D effort for Re single crystals on transition edge sensors (TES) → improve rise time to $\sim \mu\text{s}$ and energy resolution to few eV
- large arrays ($\approx 10^3$ pixels) for 10^4 - 10^5 detector experiment
- high bandwidth, multiplexed SQUID readout
- also used with ^{163}Ho loaded absorbers

MARE-1 @ Milano-Bicocca

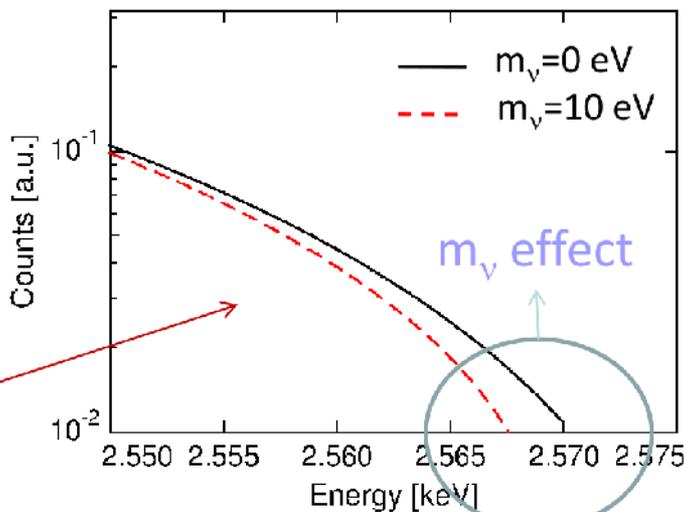
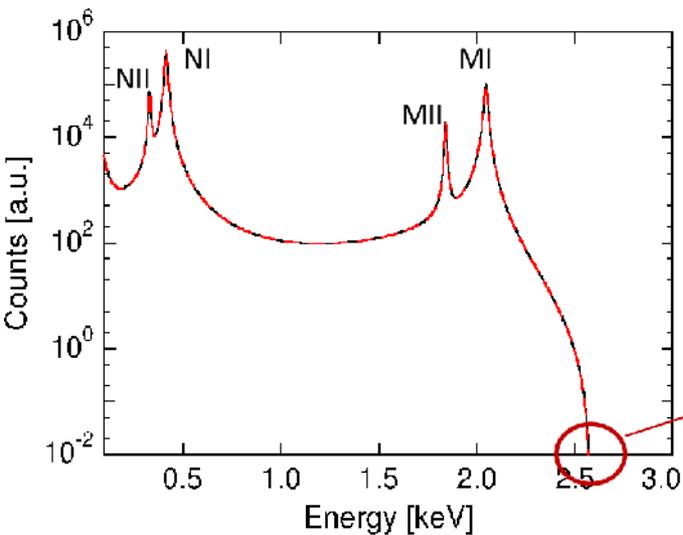
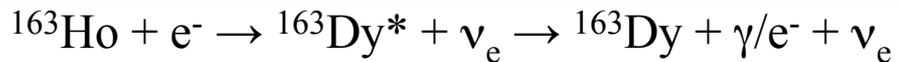
- 6x6 array of Si-implanted thermistors (NASA/GSFC)
- 0.5 mg AgReO_4 crystals
- $\Delta E \approx 30 \text{ eV}$, $\tau_R \approx 250 \mu\text{s}$
- experimental setup for up to 8 arrays completed
- starting with 72 pixels in 2011
- up to 10^{10} events in 4 years → $\sim 4 \text{ eV}$ sensitivity



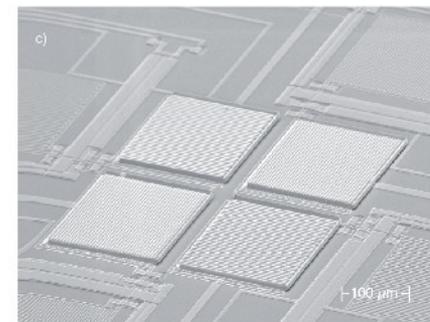
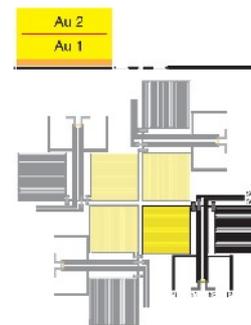
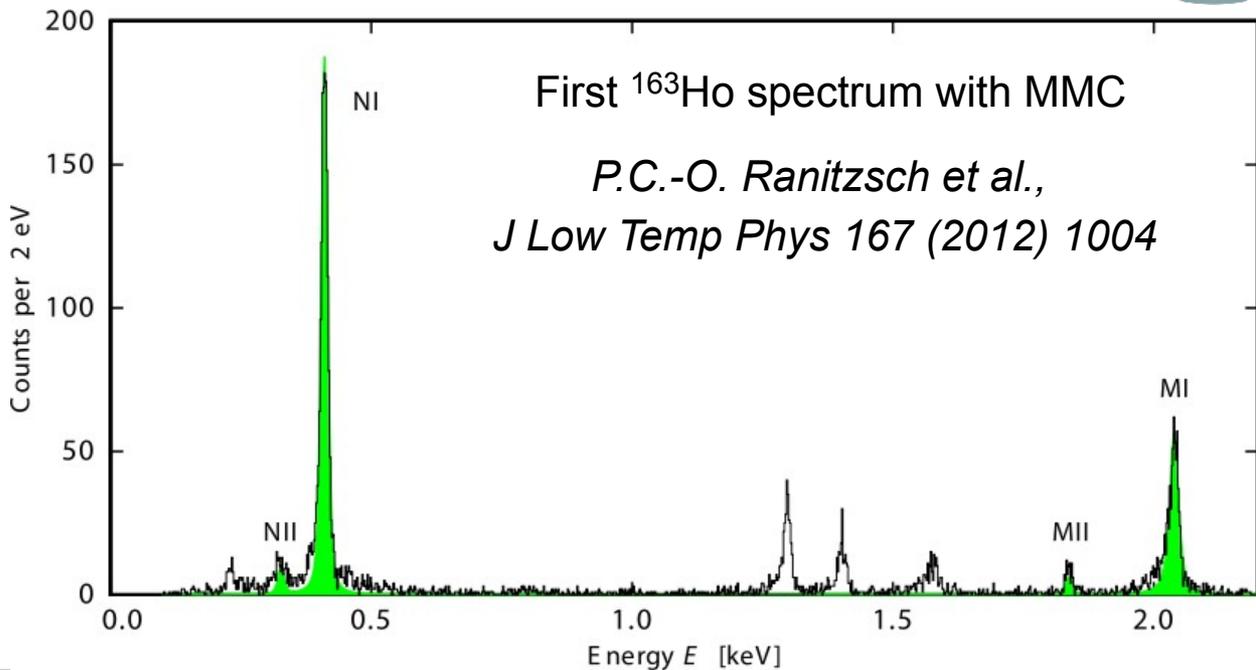
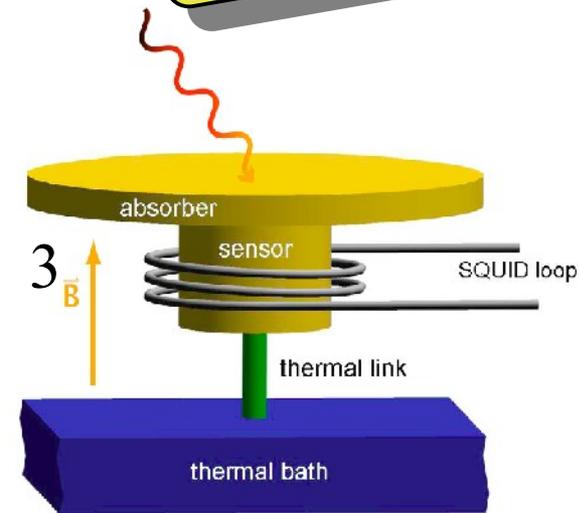
Angelo Nucciotti, Meudon 2011



ECHO neutrino mass project: ^{163}Ho electron capture with metallic magnetic calorimeters

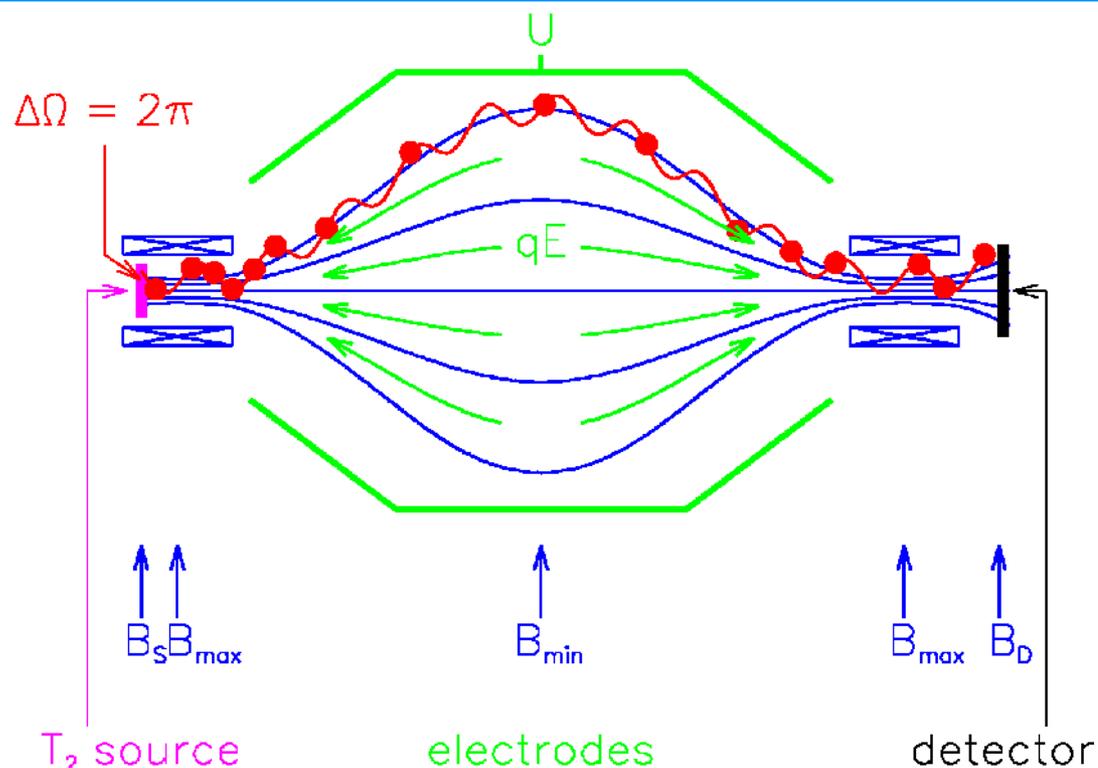


see talk by L. Gastaldo



courtesy L. Gastaldo

The classical way: Tritium β -spectroscopy with a MAC-E-Filter



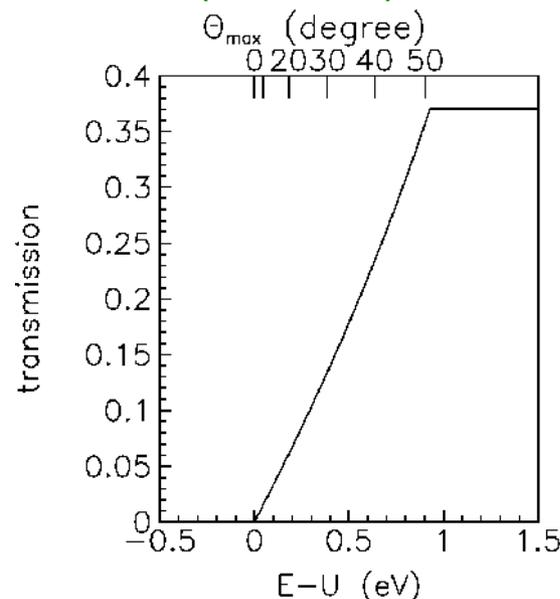
p_e (without E field)



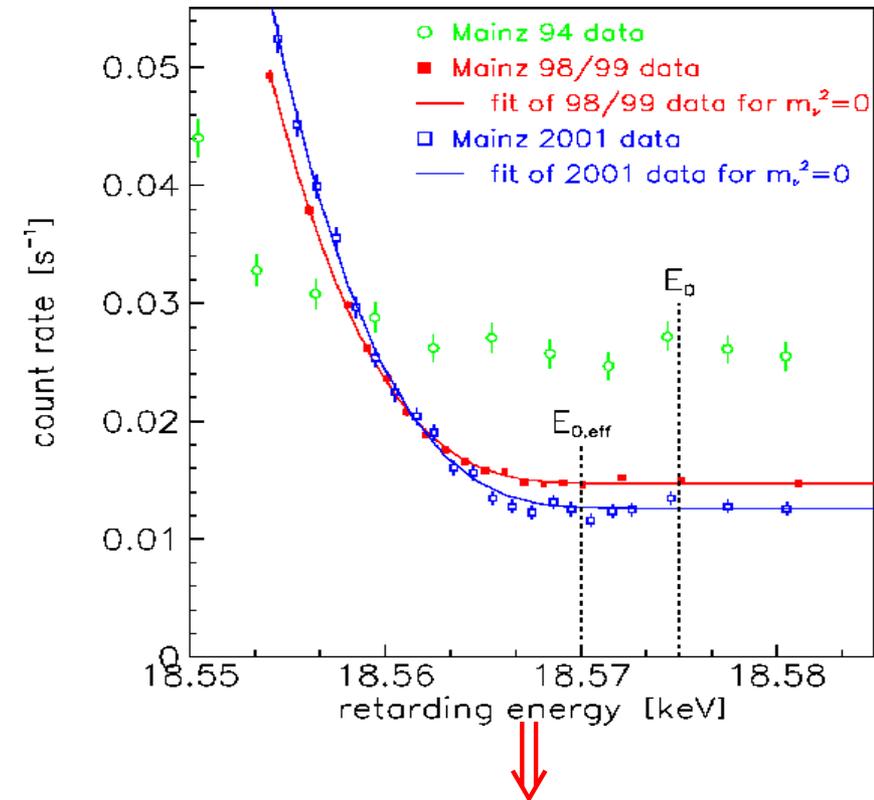
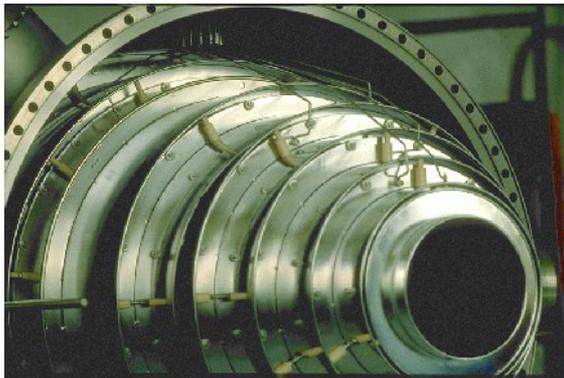
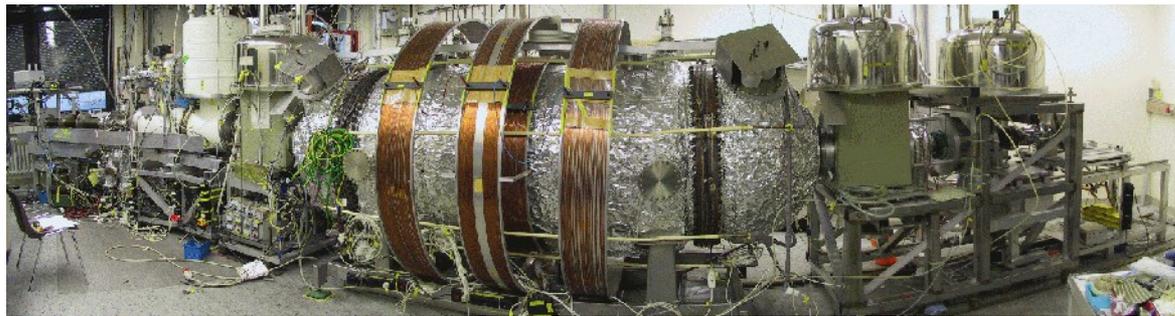
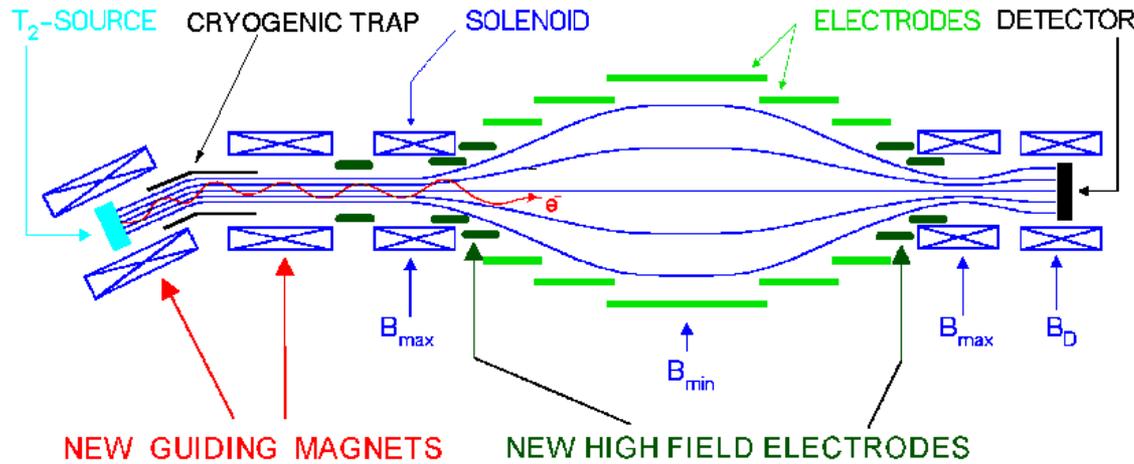
⇒ sharp integrating transmission function without tails →

Magnetic Adiabatic Collimation + Electrostatic Filter
(A. Picard et al., Nucl. Instr. Meth. 63 (1992) 345)

- Two supercond. solenoids compose magnetic guiding field
- adiabatic transformation:
 $\mu = E_{\perp} / B = \text{const.}$
⇒ parallel e^- beam
- Energy analysis by electrostat. retarding field
 $\Delta E = E \cdot B_{\min} / B_{\max}$
 $= 0.93 \text{ eV (KATRIN)}$



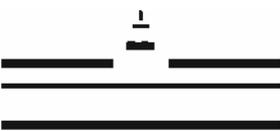
The Mainz Neutrino Mass Experiment Phase 2: 1997-2001



After all critical systematics measured by own experiment
(atomic physics, surface and solid state physics:
inelastic scattering, self-charging, neighbour excitation):

$$m^2(\nu) = -0.6 \pm 2.2 \pm 2.1 \text{ eV}^2 \Rightarrow m(\nu) < 2.3 \text{ eV (95\% C.L.)}$$

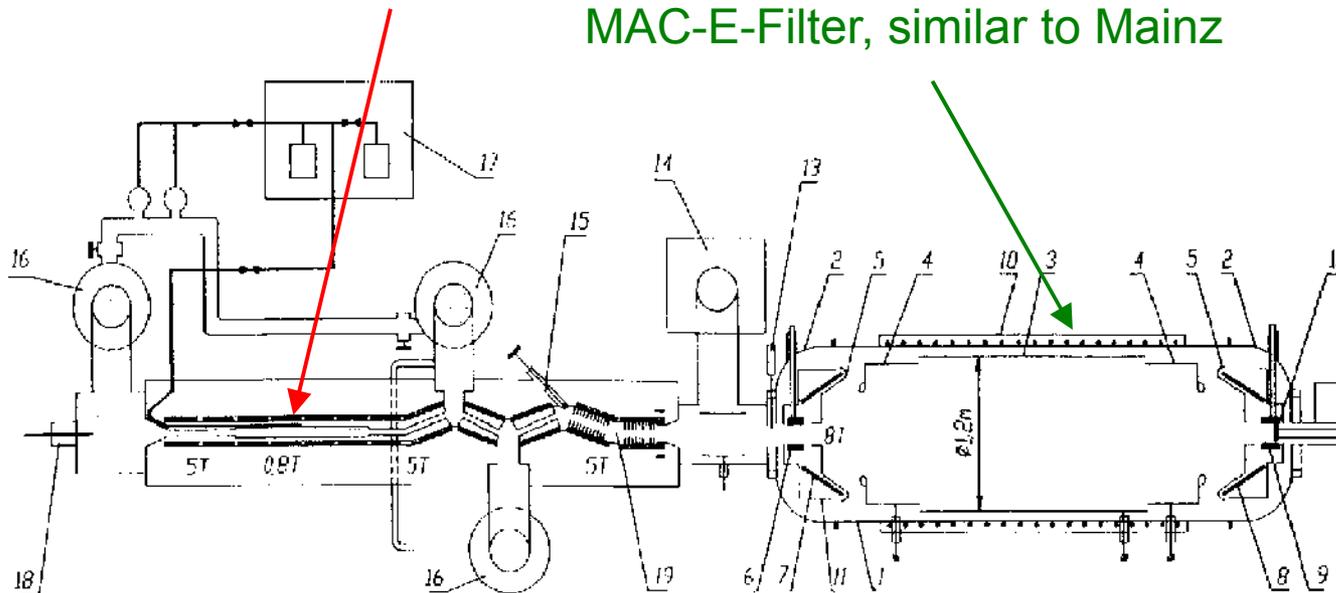
C. Kraus et al., Eur. Phys. J. C 40 (2005) 447



The Troitsk Neutrino Mass Experiment

windowless gaseous T_2 source, similar to LANL

MAC-E-Filter, similar to Mainz



Re-analysis of data
(better source thickness,
better run selection)
Aseev et al, Phys. Rev. D 84,
112003 (2011)
 $m_\beta < 2.05$ eV, 95% CL

Luminosity: $L = 0.6 \text{ cm}^2$
($L = \Delta\Omega/2\pi * A_{\text{source}}$)

Energy resolution: $\Delta E = 3.5 \text{ eV}$
3 electrode system in 1.5m
diameter UHV vessel ($p < 10^{-9}$ mbar)



upgraded
setup at Troitsk
for KATRIN
systematics
and sterile keV
neutrino search

The KATRIN experiment at KIT



Aim: $m(\nu_e)$ sensitivity of 200 meV (currently 2 eV)

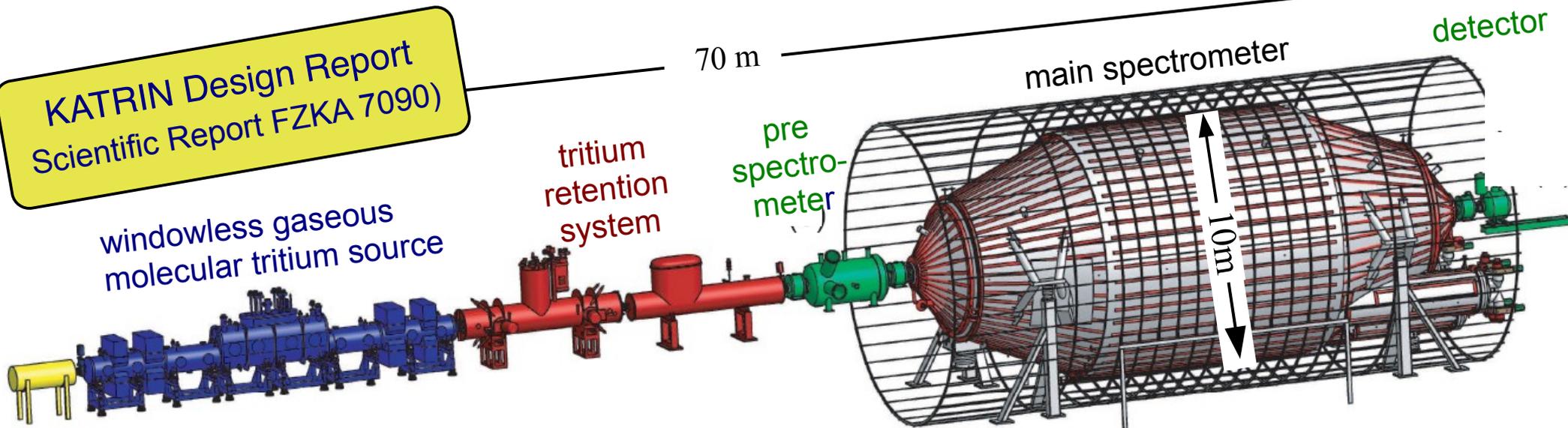
- very high energy resolution ($\Delta E \leq 1\text{eV}$, i.e. $\sigma = 0.3\text{ eV}$) \Rightarrow source \neq spectrometer concept
- strong, opaque source $\Rightarrow dN/dt \sim A_{\text{source}}$
- magnetic flux conservation (Liouville) \Rightarrow scaling law:

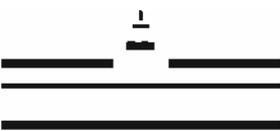
$$A_{\text{spectrometer}} / A_{\text{source}} = B_{\text{source}} / B_{\text{spectrometer}} = E / \Delta E =$$

20000 / 1

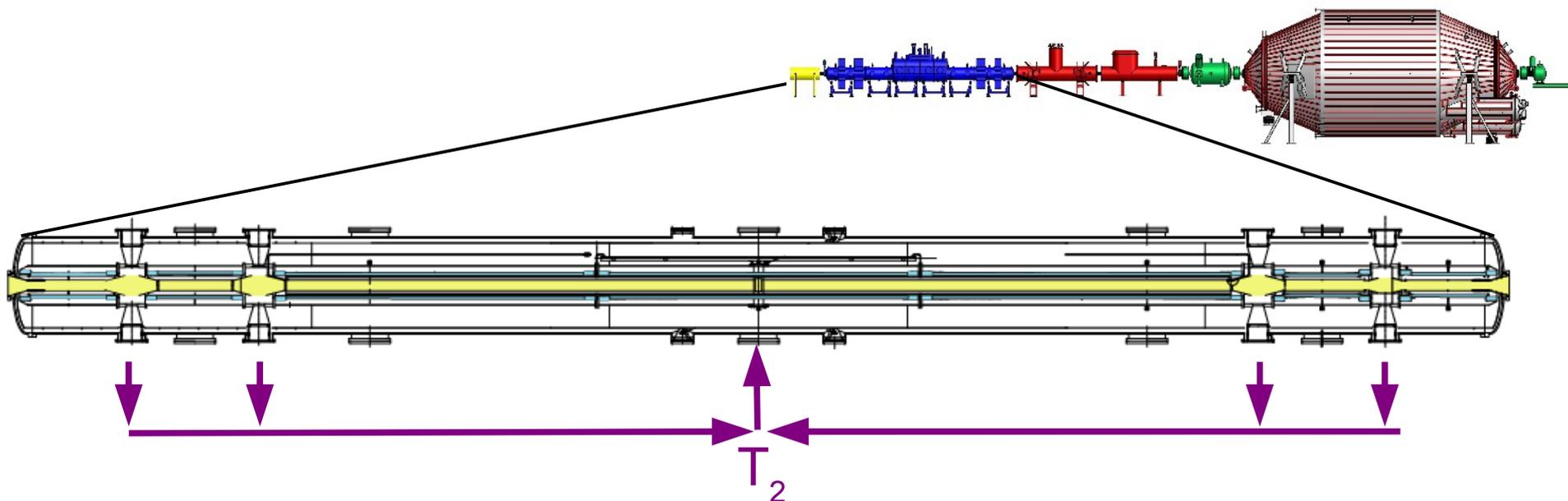
KATRIN Design Report
Scientific Report FZKA 7090

windowless gaseous
molecular tritium source





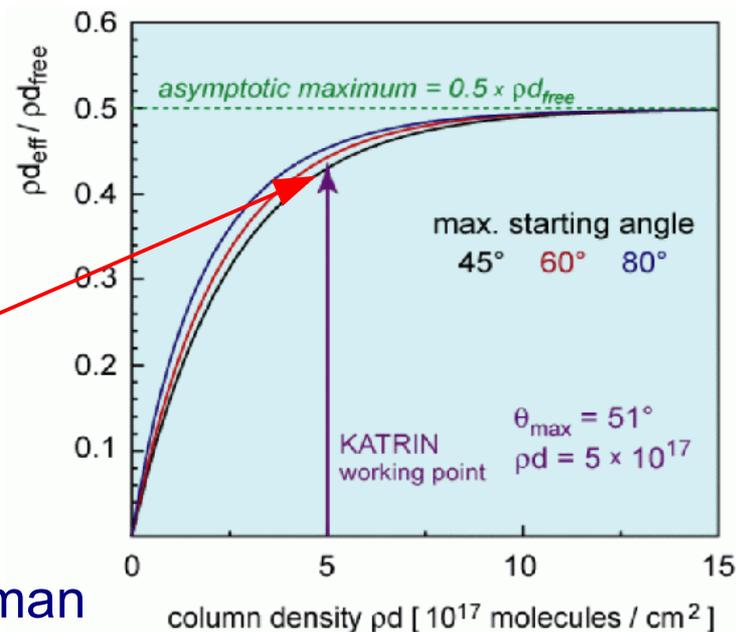
Molecular Windowless Gaseous Tritium Source WGTS



WGTS: tub in long superconducting solenoids
∅ 9cm, length: 10m, T = 30 K

Tritium recirculation (and purification)
 $p_{inj} = 0.003$ mbar, $q_{inj} = 4.7$ Ci/s

allows to measure with near to maximum count rate using
 $\rho d = 5 \cdot 10^{17}/\text{cm}^2$
with small systematics



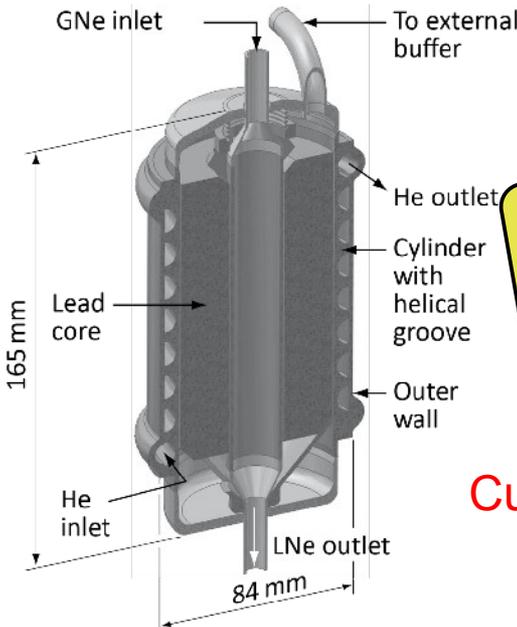
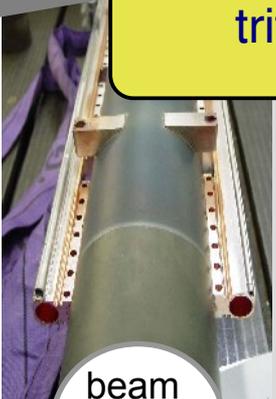
check column density by e-gun, T₂ purity by laser Raman

Very successful cool-down and stability tests of the WGTS demonstrator



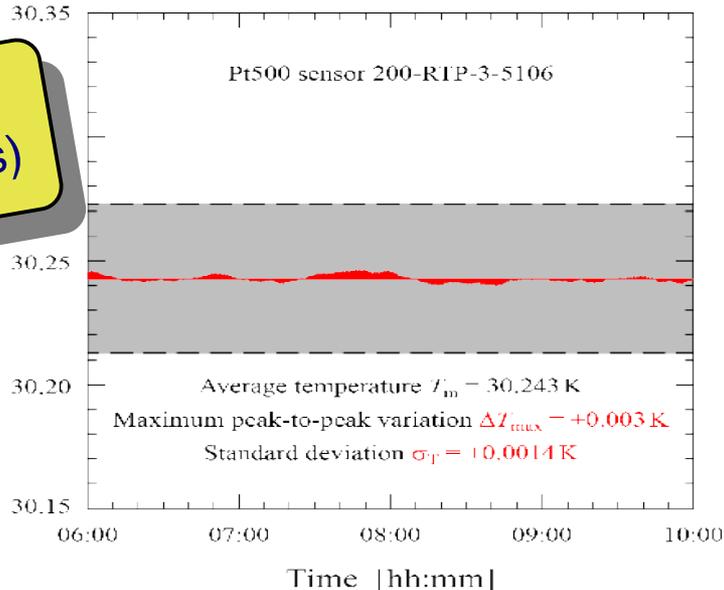
see talk by S. Fischer

per mill stability source strength request:
 $\overline{dN/dt} \sim f_T \cdot N / \tau \sim n = f_T \cdot p V / R T$
 tritium fraction f_T & ideal gas law

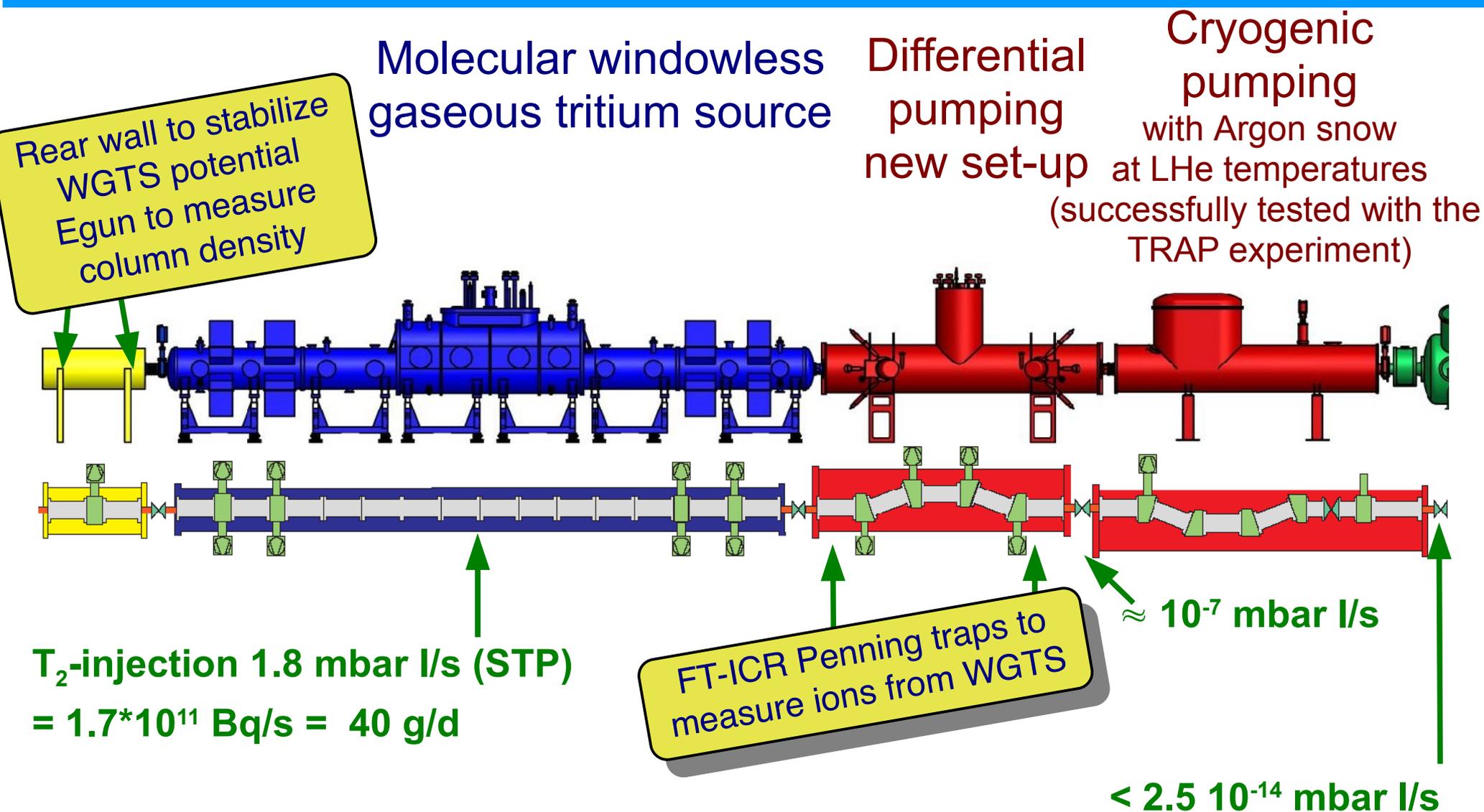


S. Grohmann et al.,
 Cryogenics, (2013, in press)

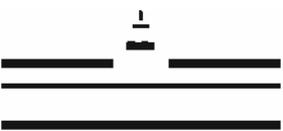
Currently: tests of sc magnets,
 constructing of WGTS
 out of demonstrator



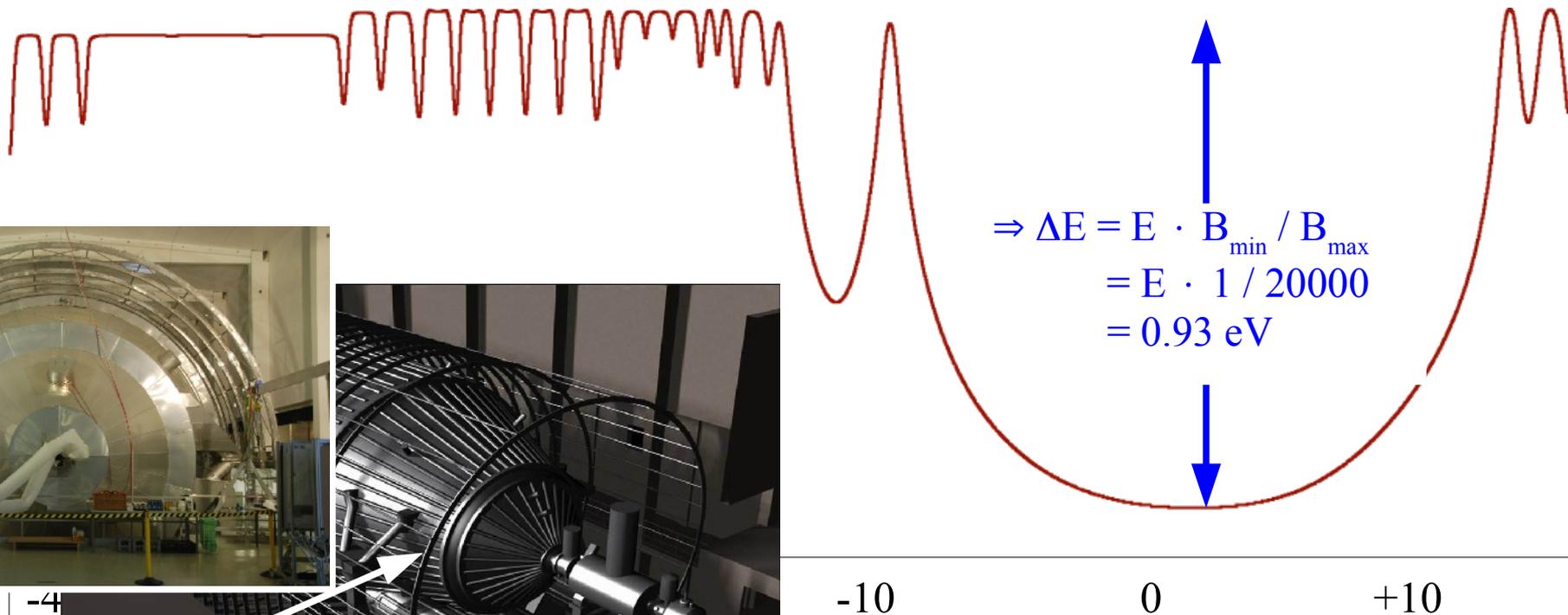
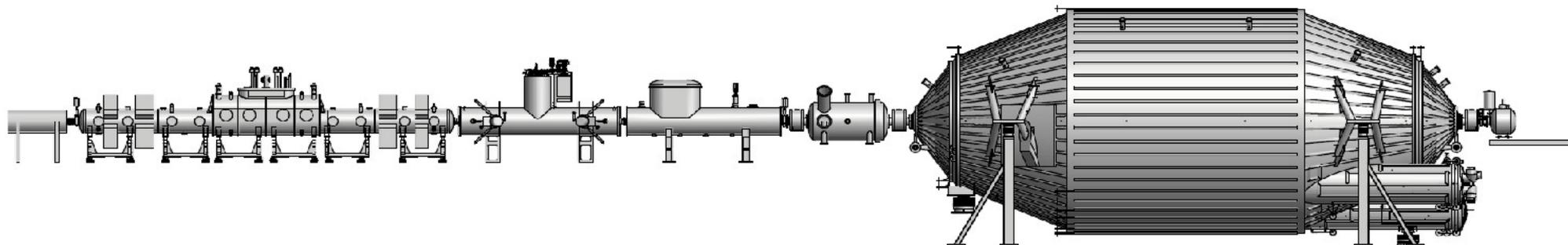
Transport and differential & cryo pumping sections



⇒ adiabatic electron guiding & T_2 reduction factor of $\sim 10^{14}$

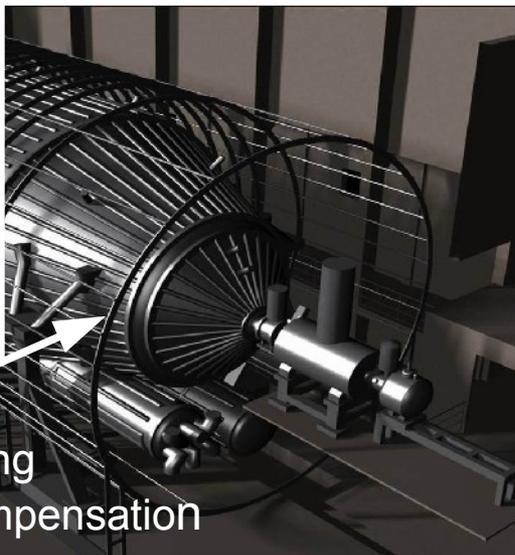


Electromagnetic design: magnetic fields

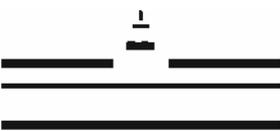


B-field [T]

-10 0 +10
distance from analysing plane [m]



-4
aircoils:
axial field shaping
+ earth field compensation



Main Spectrometer



Suppress secondary electron background from walls on high potential

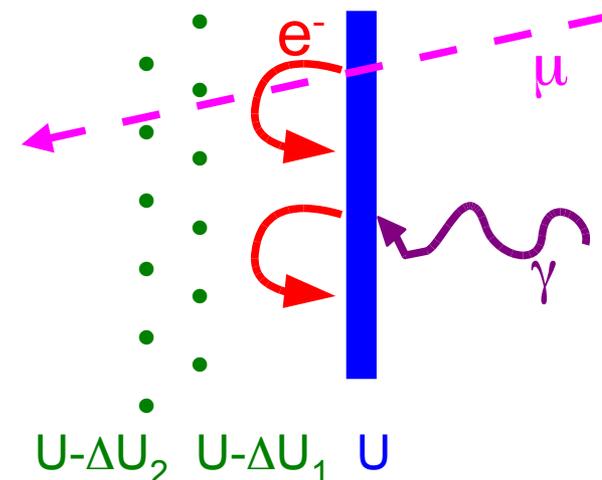
Secondary electrons from wall/electrode

by cosmic rays, environmental radioactivity, ...

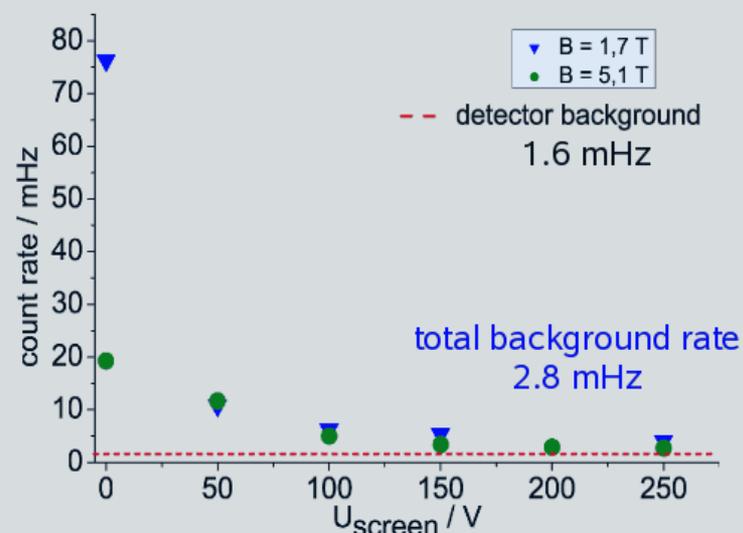
New: double layer wire electrode

on slightly more negative potential

(ca. 23,000 wires, 200 μm precision, UHV compatible)



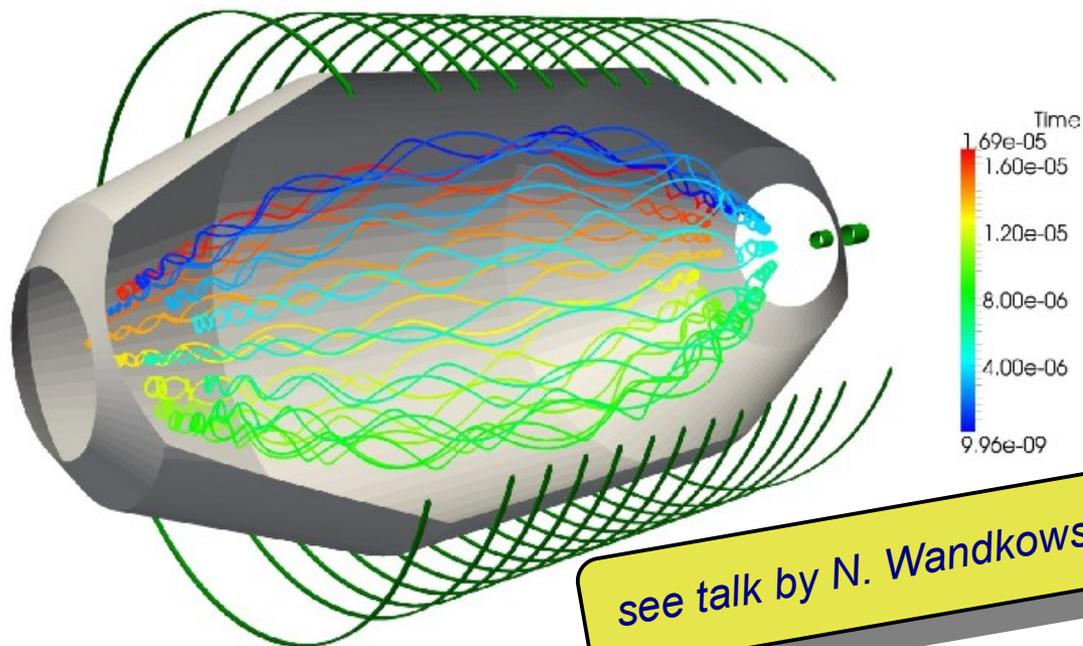
Background suppression **successfully tested** at the Mainz MAC-E filter:



Dipl. thesis B. Ostrick (U Mainz, 2002),
PhD thesis B. Flatt (U Mainz, 2004)

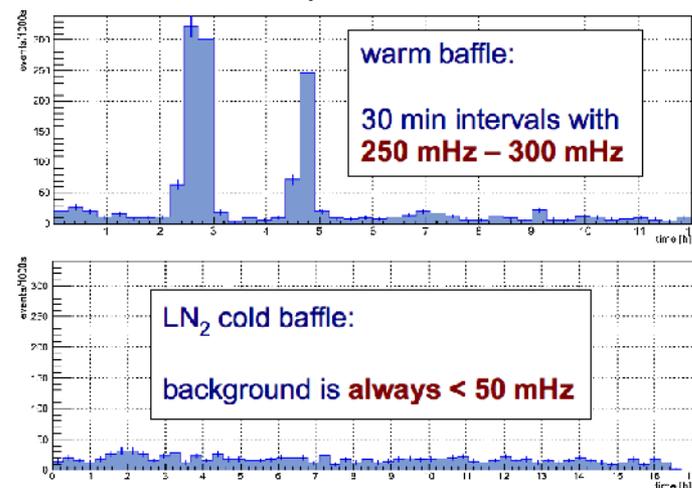
Background from stored electrons: methods to avoid or to eliminate them

Stored electron by magnetic mirrors
F. Fränkle et al., *Astropart. Phys.* 35 (2011) 128

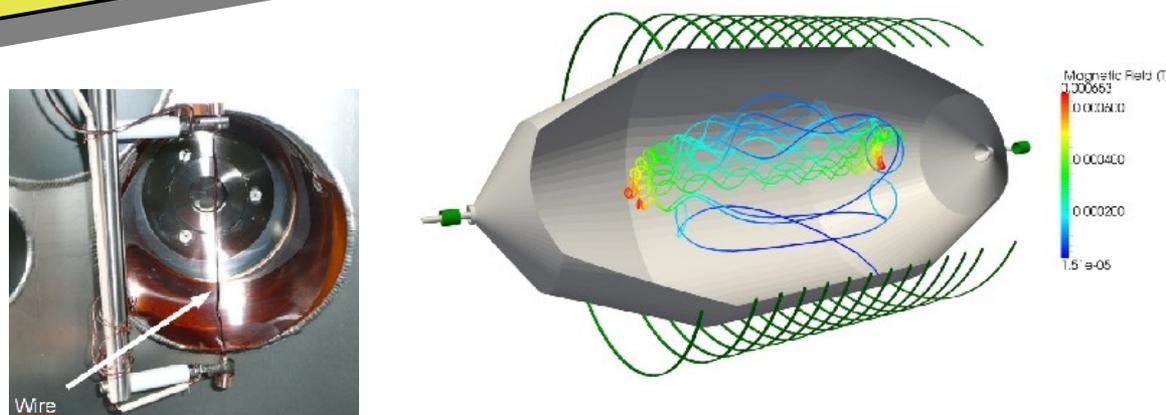


see talk by N. Wandkowsky

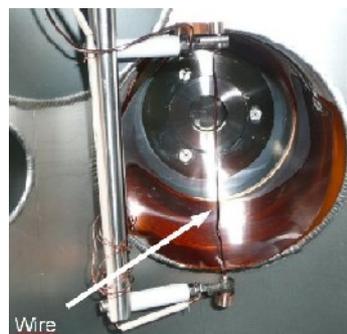
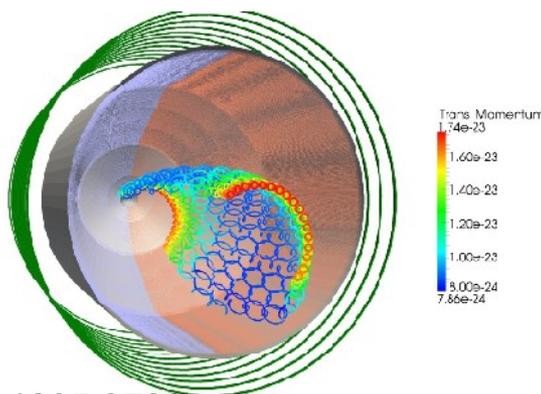
Radon suppression by LN₂ cooled baffle
S. Görhardt, diploma thesis, KIT



Nulling magnetic field by magn. pulse
B. Hillen, PhD thesis, Münster



radial E x B drift
due to electric
dipole pulse



Mechanical eliminating stored particles:
M. Beck et al, *Eur. Phys. J. A*44 (2010) 499

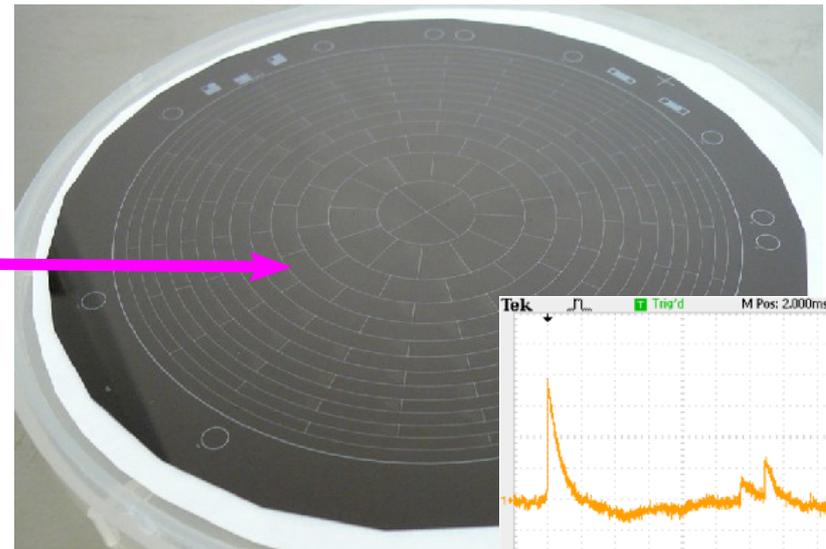
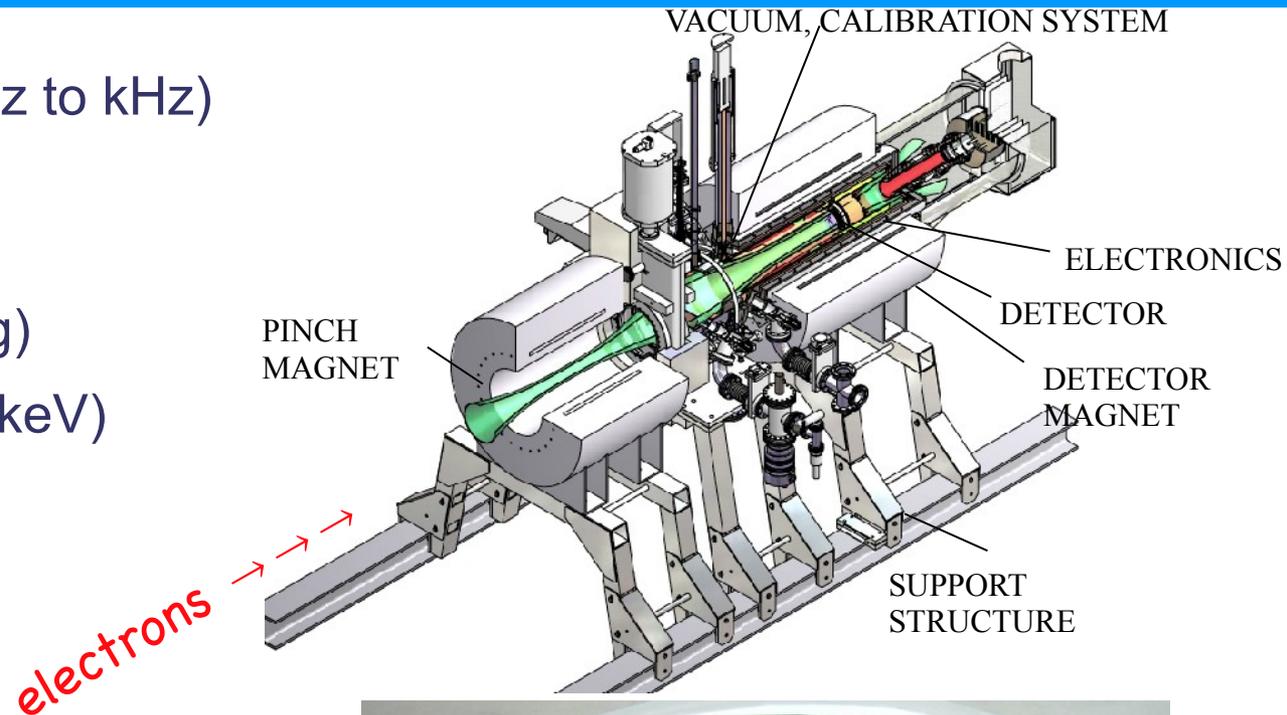
The detector

Requirements

- detection of β -electrons (mHz to kHz)
- high efficiency ($> 90\%$)
- low background (< 1 mHz)
(passive and active shielding)
- good energy resolution (< 1 keV)

Properties

- 90 mm \varnothing Si PIN diode
- thin entry window (50nm)
- detector magnet 3 - 6 T
- post acceleration (30kV)
(to lower background in signal region)
- segmented wafer (148 pixels)
 - record azimuthal and radial profile of the flux tube
 - investigate systematic effects
 - compensate field inhomogeneities



As smaller $m(\nu)$ as smaller the region of interest below endpoint E_0
→ quantum mechanical thresholds help a lot !

A few contributions with $\Delta m_\nu^2 \leq 0.007 \text{ eV}^2$ each:

1. inelastic scatterings of β 's inside WGTS

- **dedicated e-gun measurements**, unfolding of response fct.

2. fluctuations of WGTS column density (required $< 0.1\%$)

- rear detector, Laser-Raman spectroscopy, T=30K stabilisation,
e-gun measurements

3. WGTS charging due to remaining ions (MC: $\varphi < 20\text{mV}$)

- **monocrystalline rear plate short-cuts potential differences**

4. final state distribution

- **reliable quantum chem. calculations**

5. transmission function

- detailed simulations, **angular-selective e-gun measurements**

6. HV stability of retarding potential on $\sim 3\text{ppm}$ level required

- **precision HV divider (with PTB), monitor spectrometer beamline**

tritium
source

spectrometer

Systematic uncertainties

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Measuring the last 25 or 30 eV only
 KATRIN becomes nearly
 a „single final state“ experiment
 as the cryo-bolometers

tritium
source

spectrometer

Example of KATRIN simulation & fit
(last 25eV below endpoint, reference):

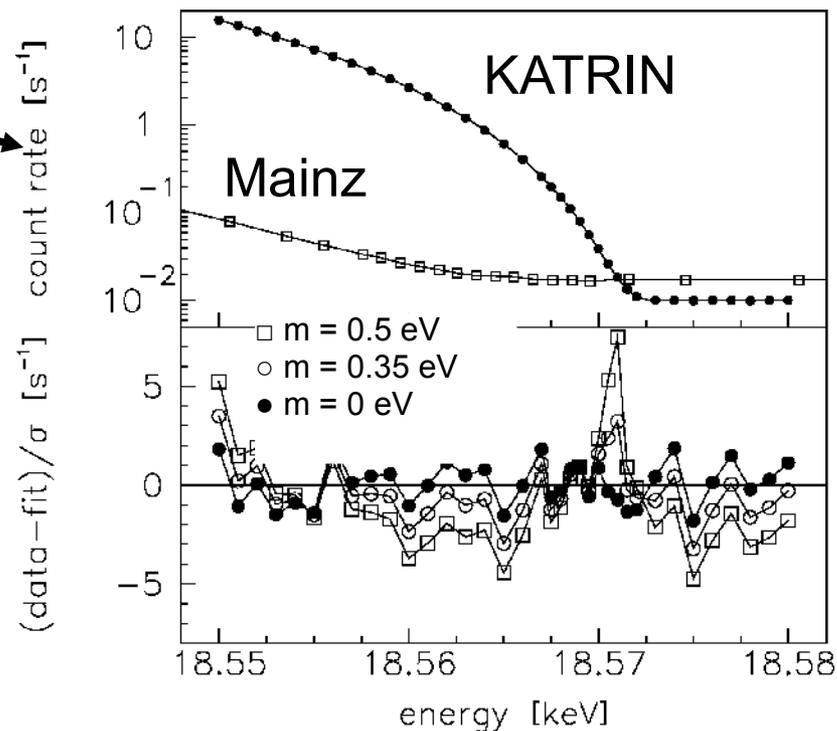
sensitivity:

$$m_\nu < 0.2\text{eV (90\%CL)}$$

discovery potential:

$$m_\nu = 0.3\text{eV} \quad (3\sigma)$$

$$m_\nu = 0.35\text{eV} \quad (5\sigma)$$



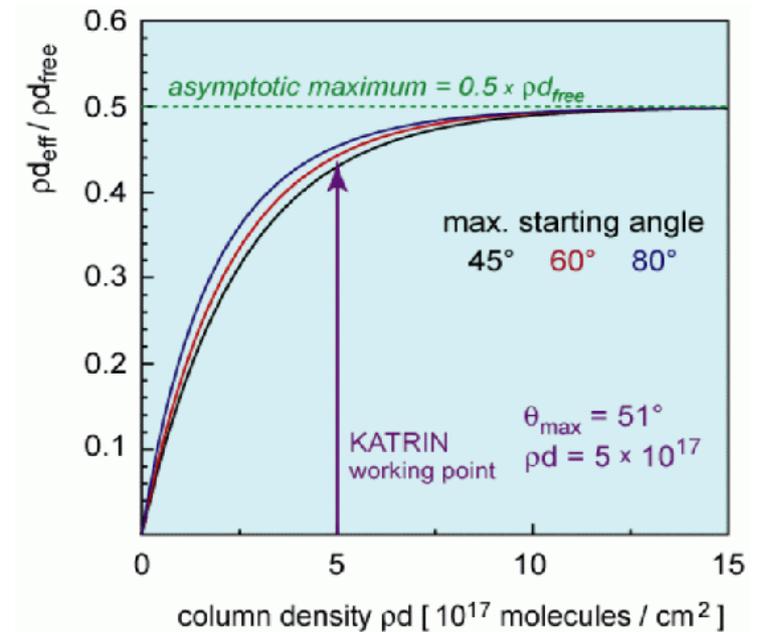
Expectation for 3 full data taking years: $\sigma_{\text{sys}} \sim \sigma_{\text{stat}}$

⇒ **KATRIN** will improve the sensitivity by 1 order of magnitude
will check the whole cosmological relevant mass range
will detect degenerate neutrinos (if they are degen.)

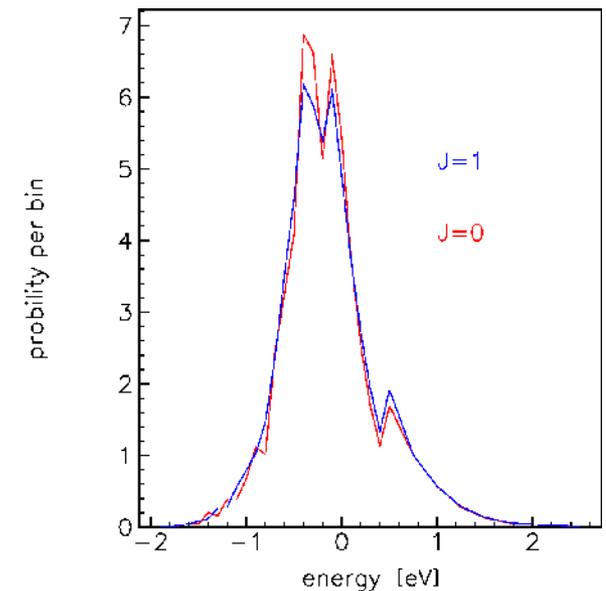
Can KATRIN be largely improved ?

Problems

- 1) The source is already opaque
→ need to increase size transversally
but a $\varnothing 100\text{m}$ spectrometer is not feasible



- 2) Resolution is limited to $\sigma = 0.34$ eV by the excitation of ro-vibrational states in the final state when using molecular tritium



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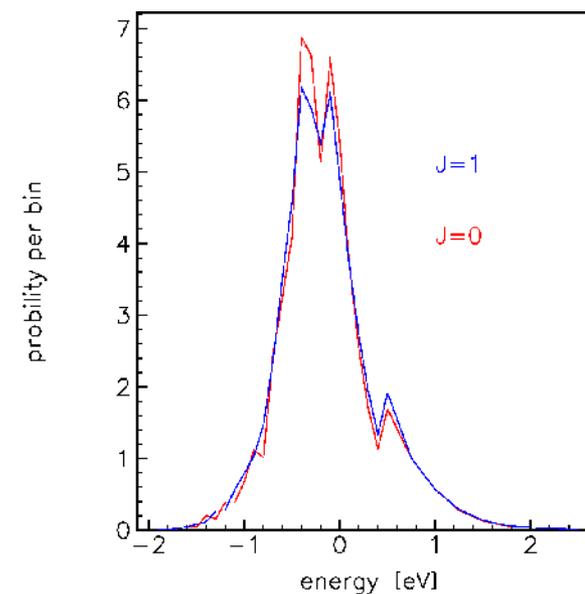
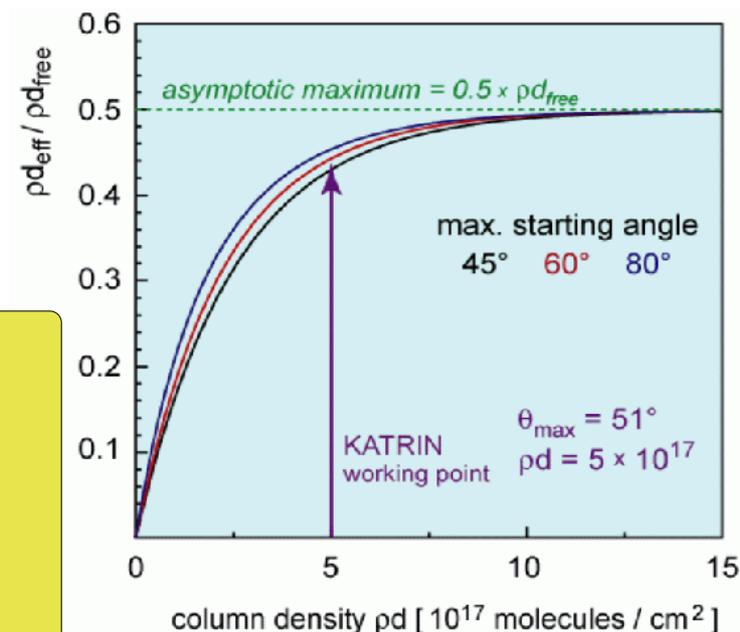
Two possible ways out:

- a) source inside detector
using cryogenic bolometers (MARE, ECHO)
- b) hand-over energy information of β electron
to other particle (photon),
which can escape tritium source (Project 8)

- 2) Resolution is limited to $\sigma = 0.34$ eV by the
excitation of ro-vibrational states in the final state
when using molecular tritium

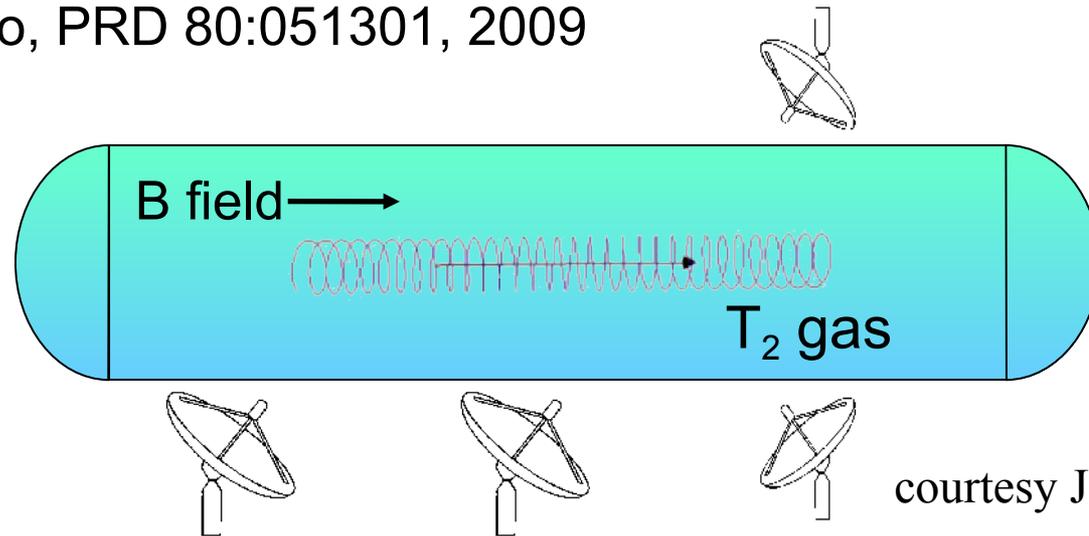
Not really realistic yet:

- c) atomic tritium source
Unfortunately, it is technically really a challenge



Outlook - Project 8: Measure coherent cyclotron radiation of tritium β electrons

B. Monreal and J. Formaggio, PRD 80:051301, 2009



courtesy J. Formaggio

General idea:

- Source = KATRIN tritium source technology :

uniform B field
low pressure T₂ gas

**β electron radiates coherent
cyclotron radiation**

$$\omega(\gamma) = \frac{\omega_0}{\gamma} = \frac{eB}{K + m_e}$$

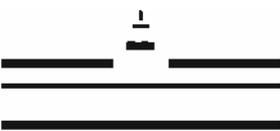
- Antenna array (interferometry) for cyclotron radiation detection since cyclotron radiation can leave the source and carries the information of the β -electron energy

A lot of R&D necessary and has started

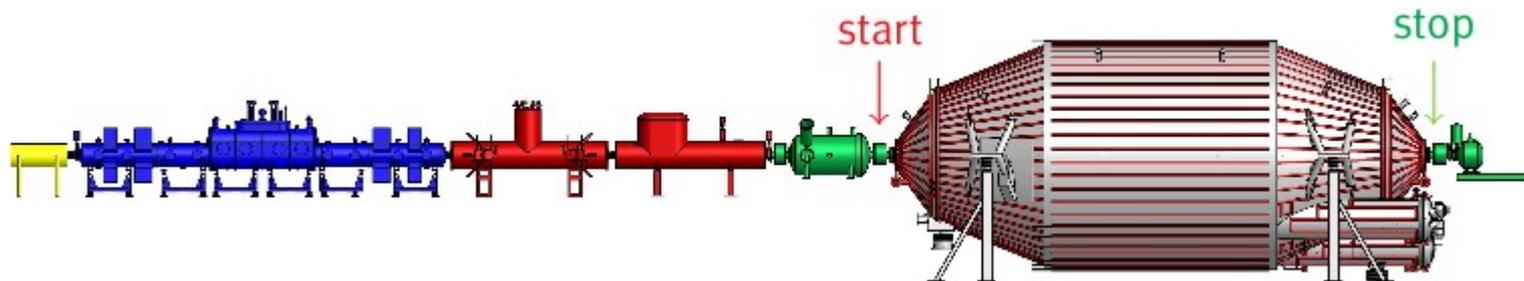
- Is it really possible ?
- What are the systematic uncertainties ?

Can KATRIN be improved a bit ?

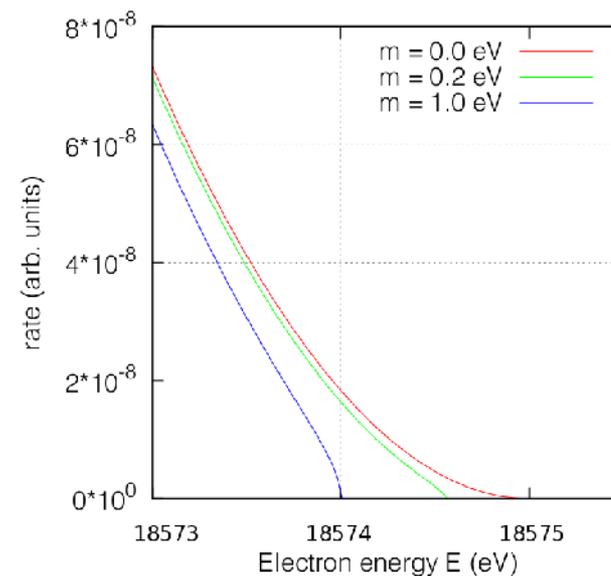
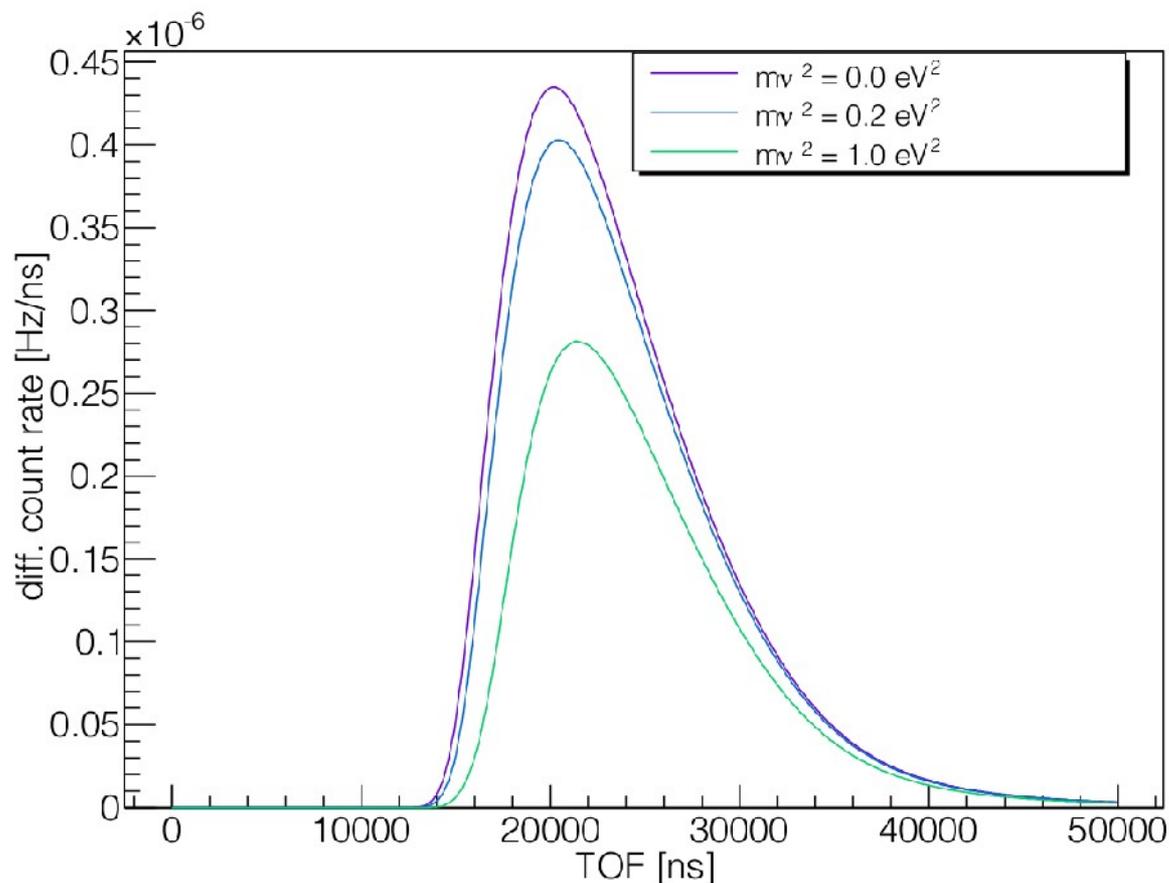
?



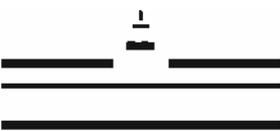
Alternative spectroscopy: measure time-of-flight TOF through KATRIN spectrometer



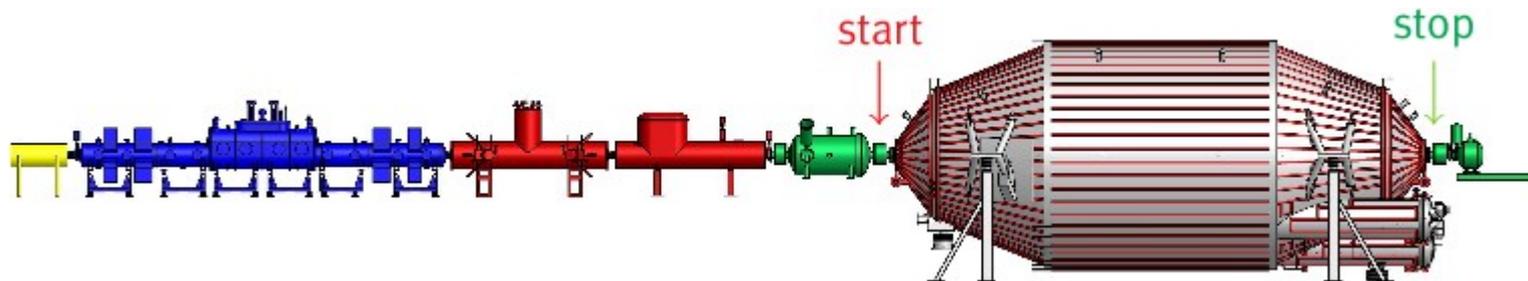
Comparison of TOF spectra for different neutrino masses for $E_0 = 18574.0$ eV, $U_{ret} = -18570.0$ eV



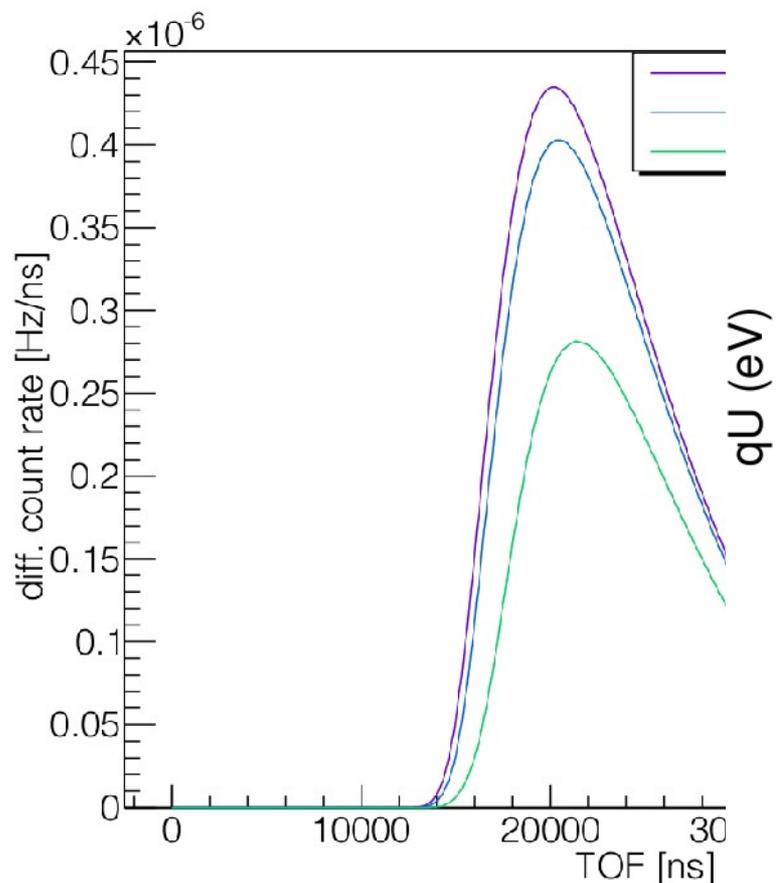
Time-of-flight spectrum is sensitive to the neutrino mass
require one retardation potential only
not integral but differential β -spectrum



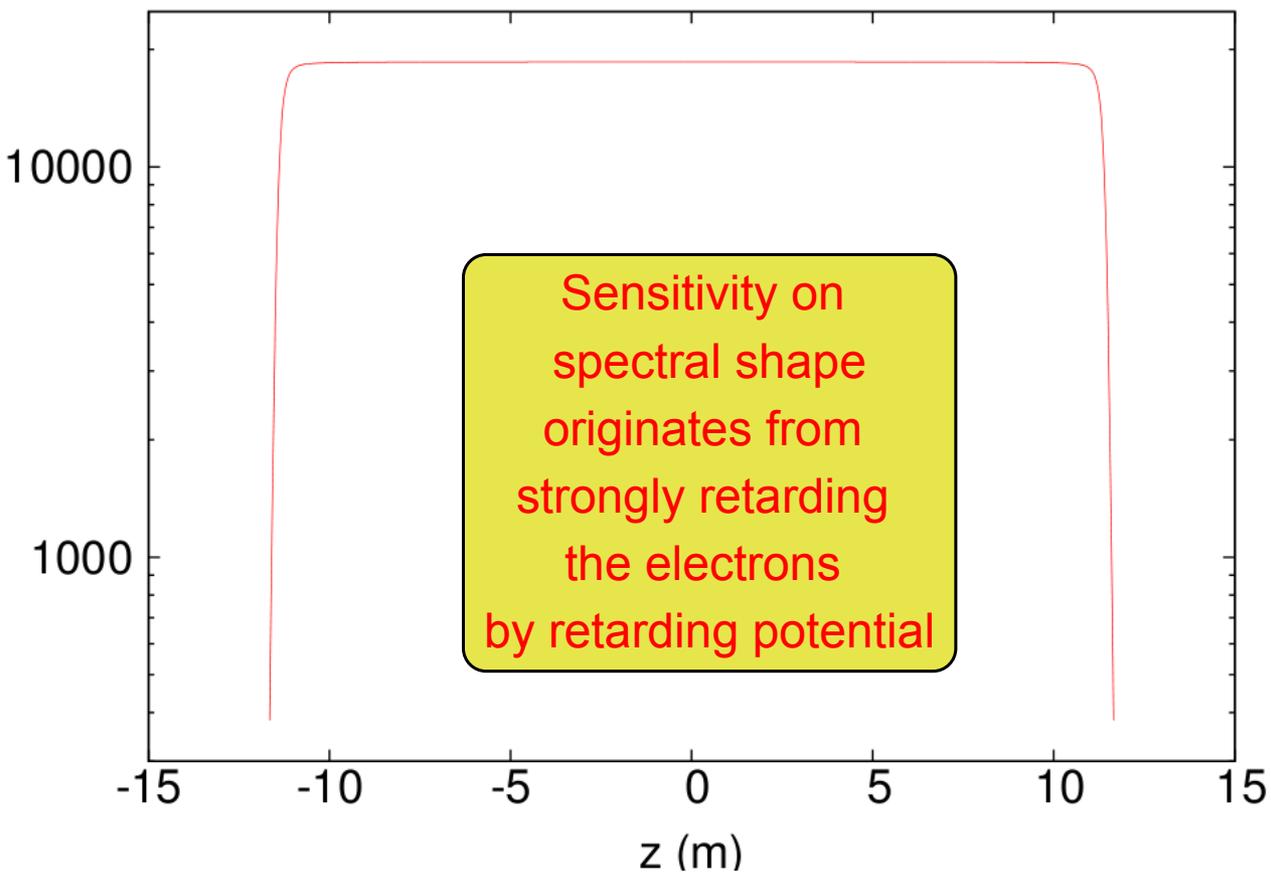
Alternative spectroscopy: measure time-of-flight TOF through KATRIN spectrometer



Comparison of TOF spectra for different neutrino masses for



Electric potential on main spectrometer z axis

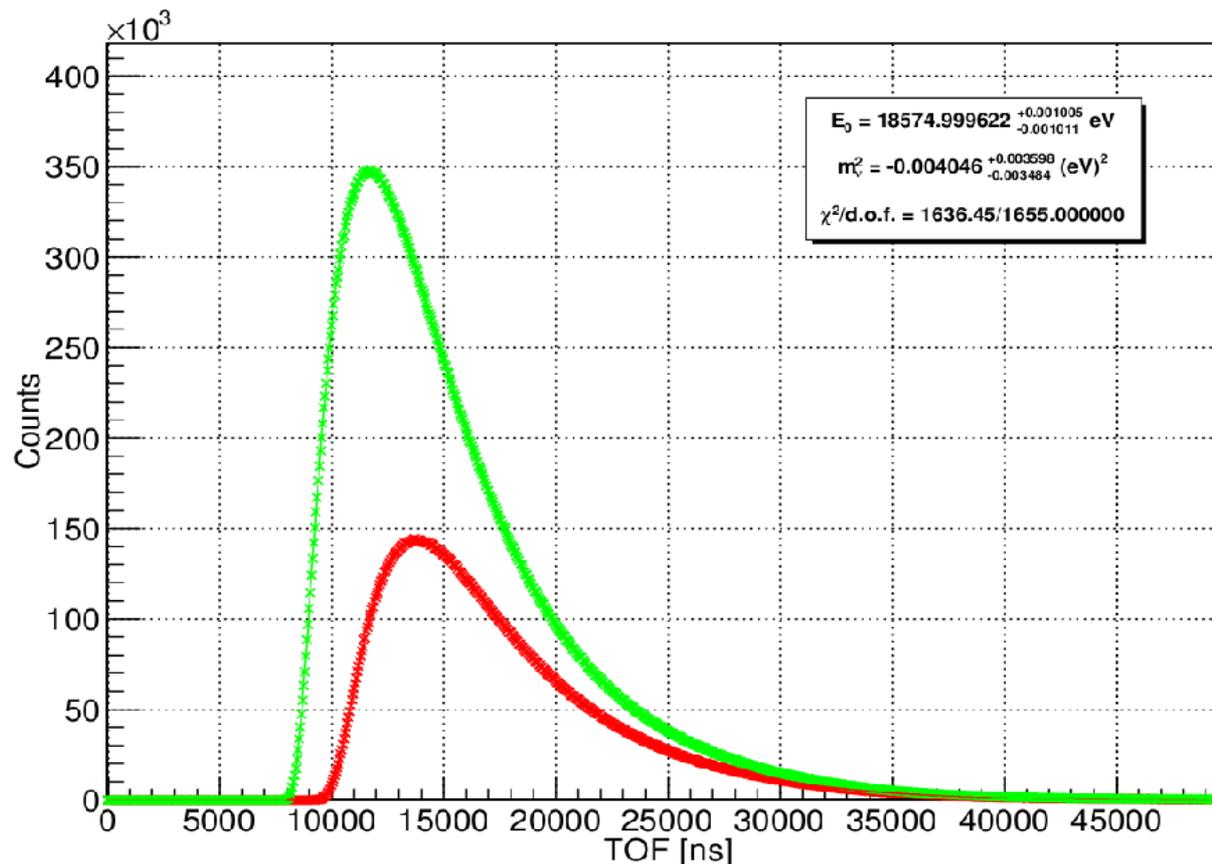


Sensitivity improvement on $m^2(\nu_e)$ by ideal TOF determination

Measure at 2 (instead of ≈ 30) different retarding potentials
since TOF spectra contain all the information

Coincidence request between start and stop signal \rightarrow nice background suppression

\rightarrow Factor 5 improvement in m_n^2 w.r.t. standard KATRIN, but ideal case !



N. Steinbrink et al.
arXiv:1308.0532

How to measure time-of-flight at KATRIN ?

→ gated-filter

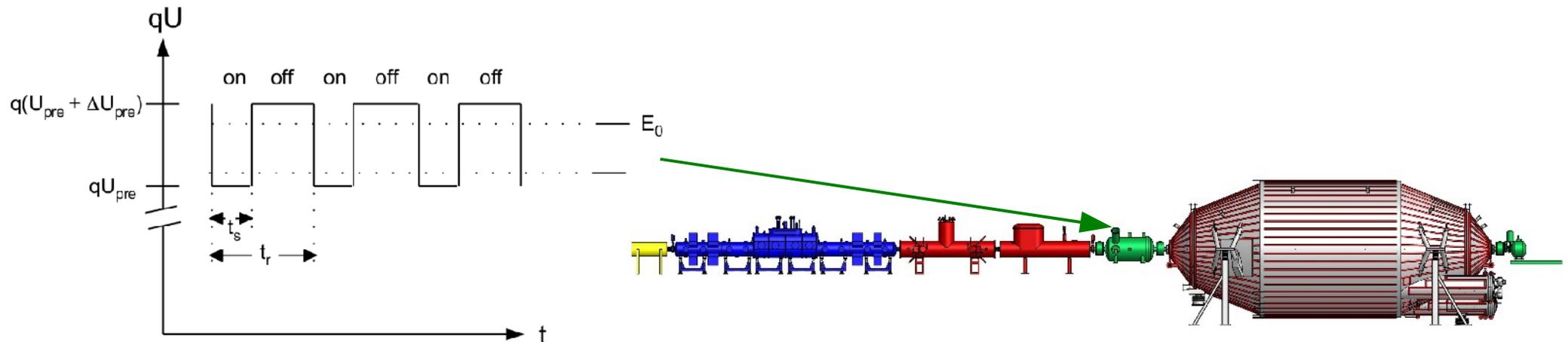
- 1) Can measure time-of-arrival with KATRIN detector with $\Delta t = 50 \text{ ns}$ → ok
- 2) Need to determine time-of-passing-by of beta electron before main spectrometer **without disturbing energy and momentum** by more than 10 meV !
→ Need „detector“ with 10 meV threshold
This seems not to be prohibited in principle but it is unrealistic for the near future !
- 2') Use **pre spectrometer** as a „gated-filter“ by switching fast the retarding voltage

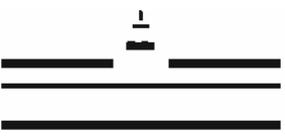
MAC-E-TOF demonstrated: *J. Bonn et al., Nucl. Instr. Meth. A421 (1999) 256*

no problem with transmission properties: *M. Prall et al., NJP 14 (2012) 073054*

About as sensitive on the neutrino mass as standard KATRIN:

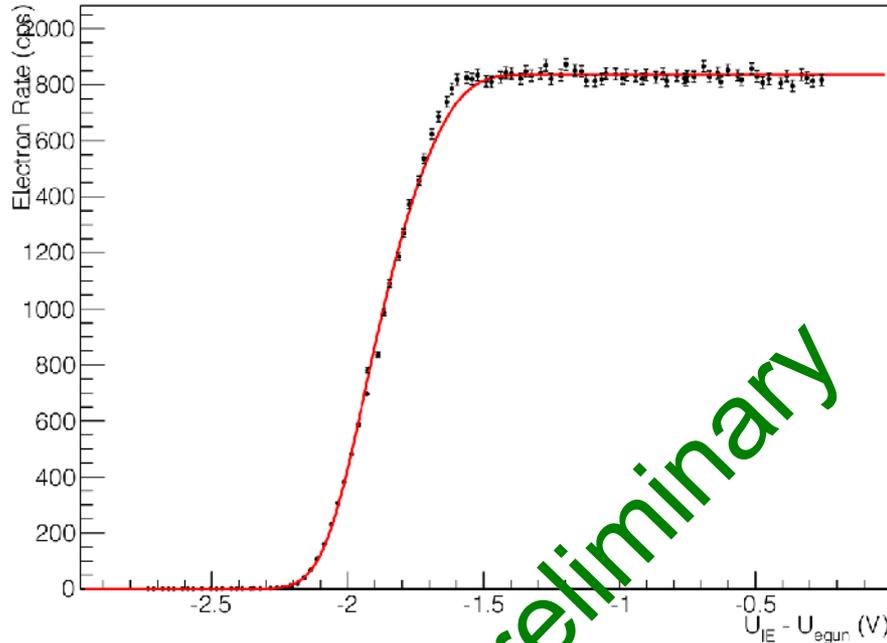
N. Steinbrink et al., arXiv:1308.0532



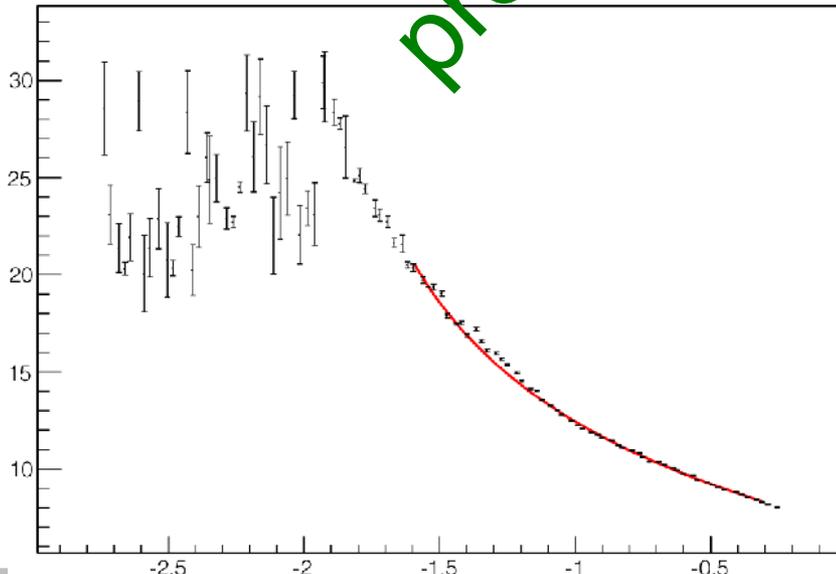


Test of TOF method at main spectrometer during SDS commissioning with pulsed egun

Transmission Function Measurement

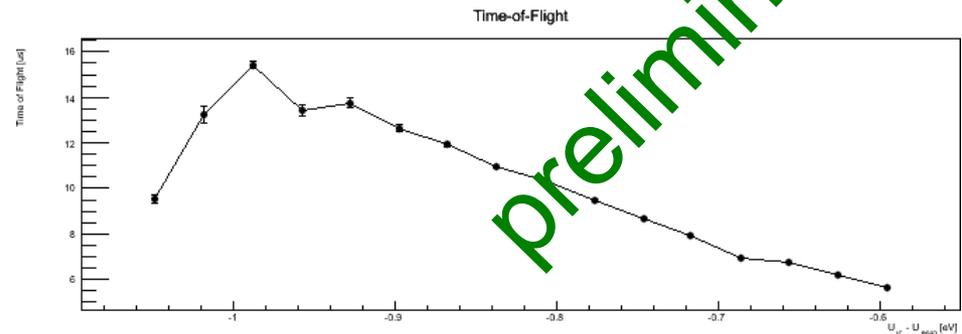
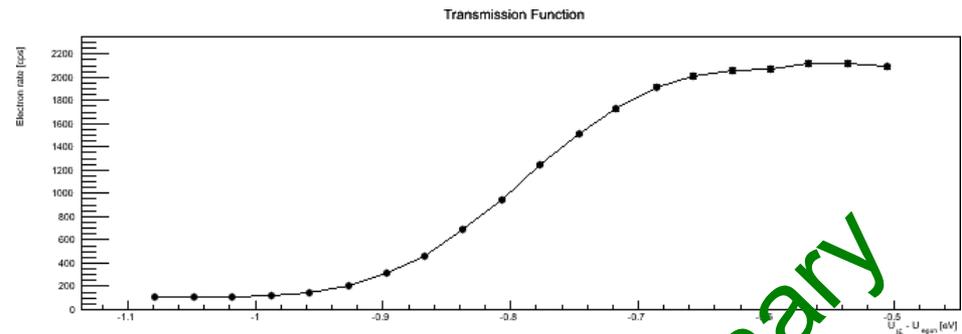


preliminary



Integral transmission function at -200 V

measured TOF agrees with
calculated TOF by N. Steinbrink
from electrical potential
and magnetic field maps



preliminary

Integral & TOF transmission function at -18600 V

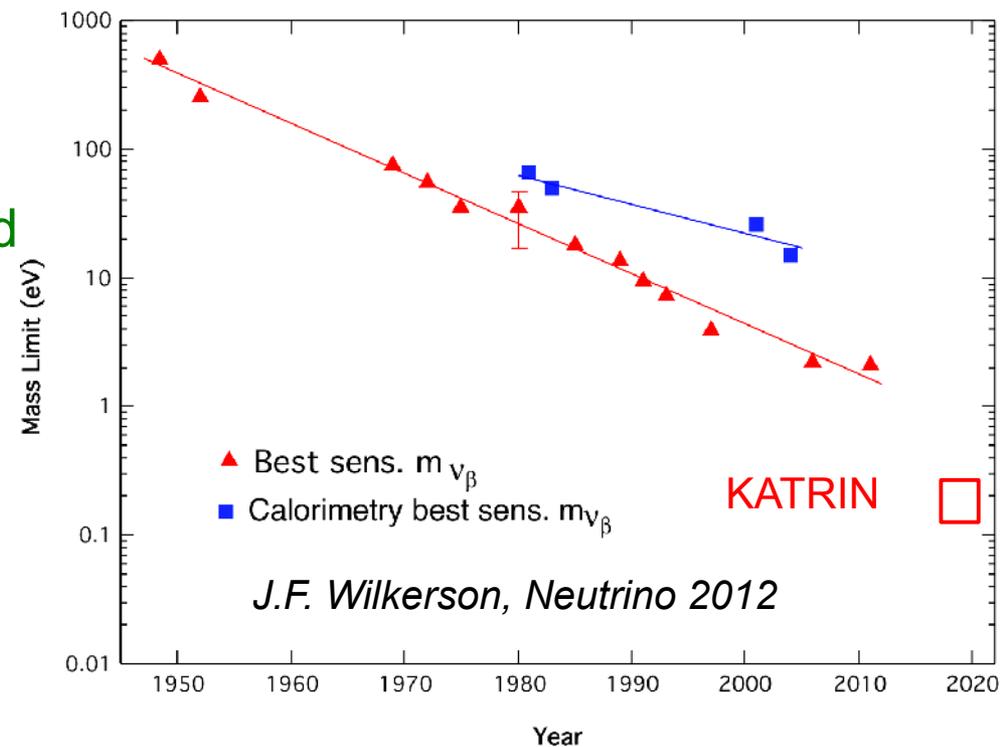
Different ways for a direct neutrino mass measurement from β -decay

- cryogenic bolometers investigating ^{187}Re β -decay (\rightarrow MARE)
- cryogenic bolometers investigating ^{163}Ho EC (\rightarrow MARE, ECHO)
- tritium β -decay using MAC-E-Filter (\rightarrow KATRIN)
- detection of synchrotron radiation (\rightarrow Projekt 8)
- ...

KATRIN is using a complex
but established method:
 \rightarrow sensitivity: 2 eV \rightarrow 200 meV
main spectrometer is being commissioned

Cryobolometers seem to have
a large potential
but need large arrays \rightarrow multiplexing
what are the systematics ?

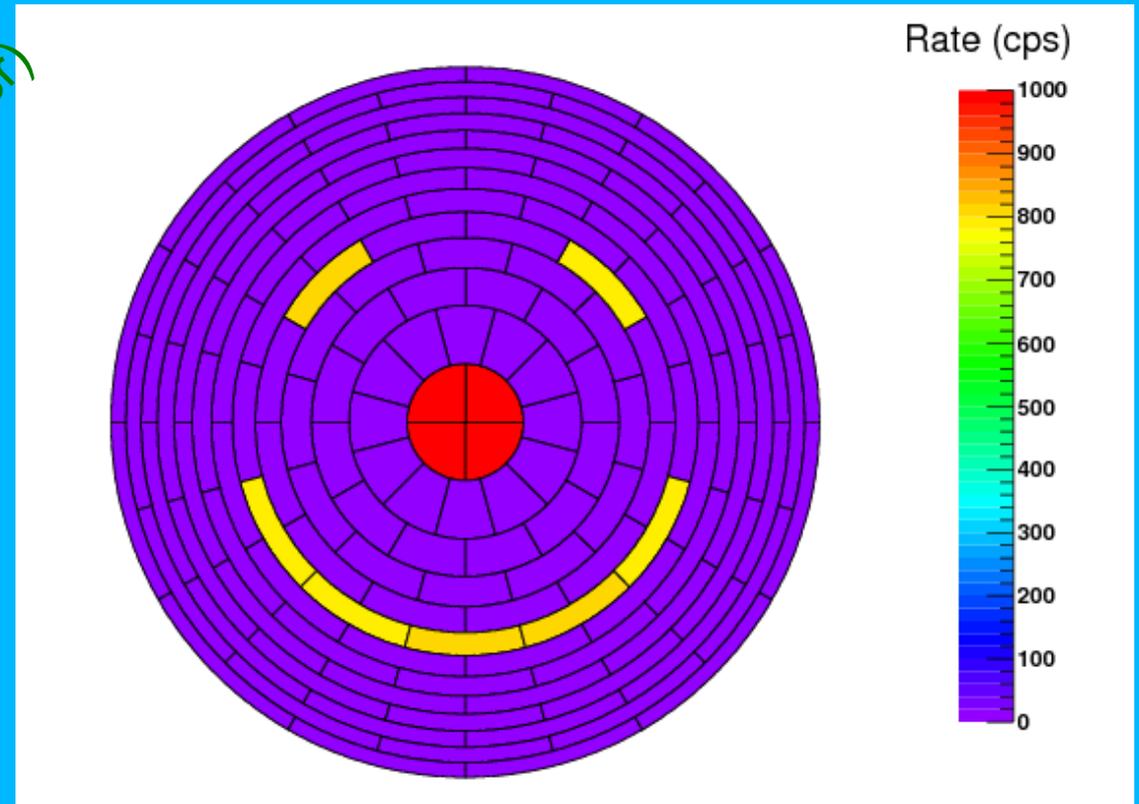
Are alternative methods feasible ?



And finally ...

**real data from the
KATRIN experiment**
(scannable egun + main spectrometer + detector)

The KATRIN 148-pixel detector is smiling



when being hit by electrons
from 11 subsequent positions
of the scanning photoelectron source